



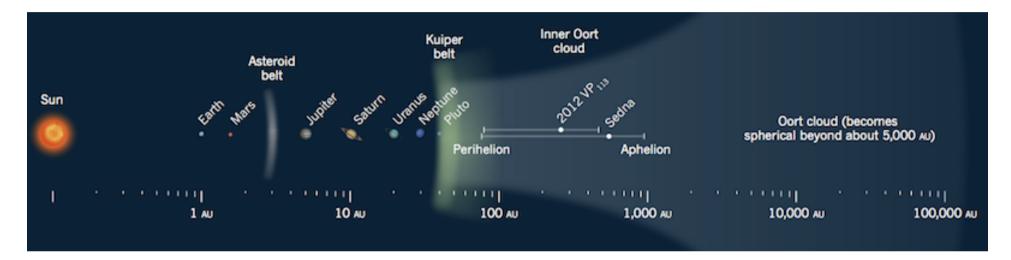
# The Chemical Composition of an Extrasolar Kuiper-Belt-Object

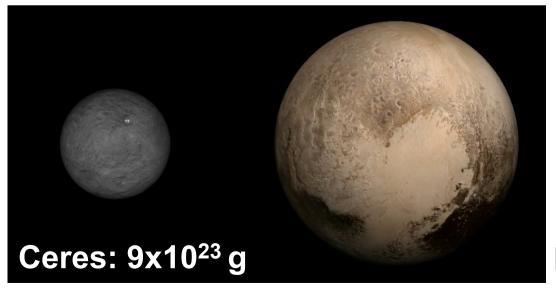
# Siyi Xu (许偲艺) ESO Fellow

In collaboration with... B. Zuckerman, P. Dufour, E. D. Young, B. Klein & M. Jura 2017, ApJL, 836, L7



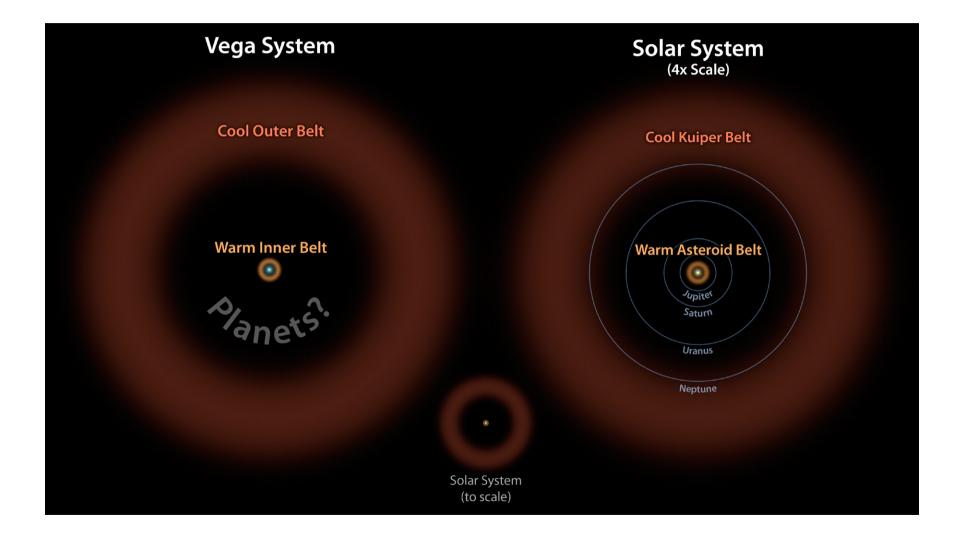
### **Minor Planets in the Solar System**



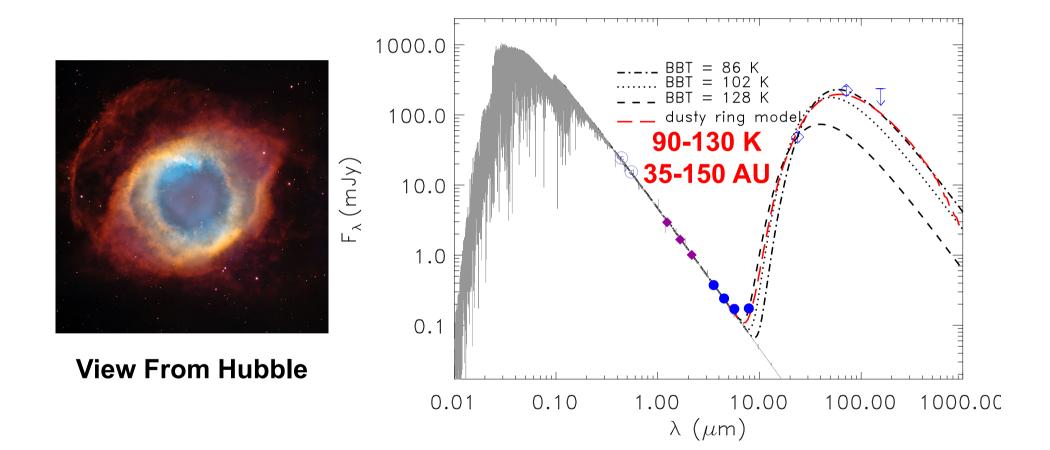


Pluto: 10<sup>25</sup> g

# **KBO Analogs around Other Stars**



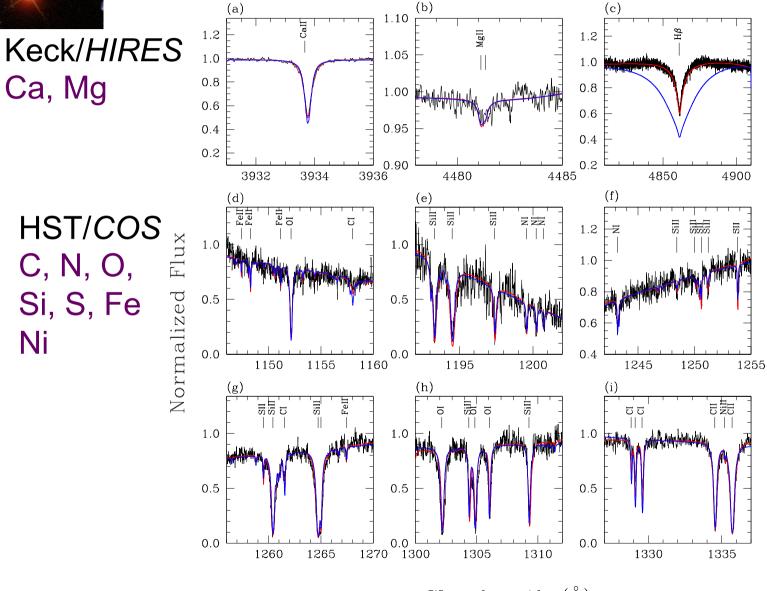
#### Helix Nebula: Collisions between KBOs?



Su+2007

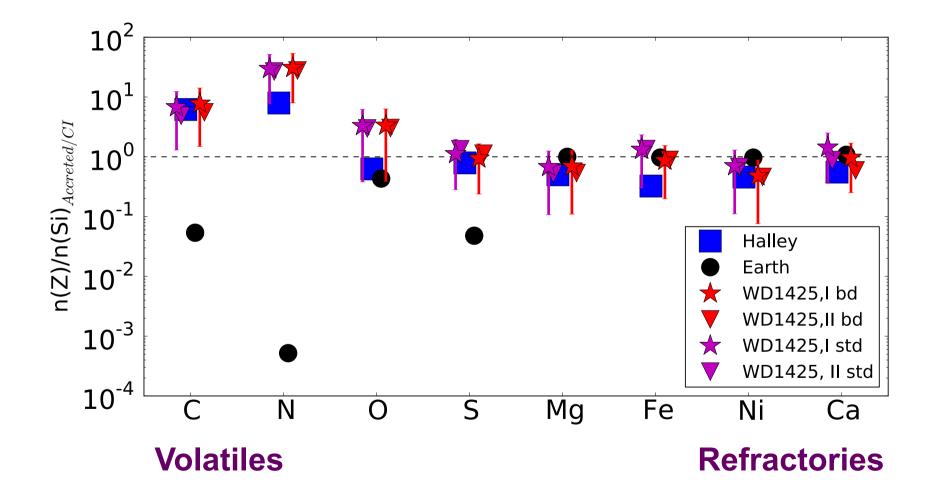


#### WD 1425+540: Spectroscopic Observations

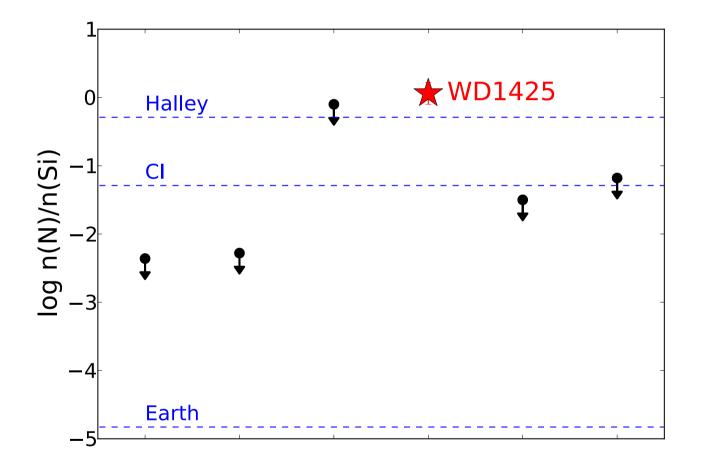


Wavelength (Å)

# Accretion from Comet Halley Analog but at least ~100,000 more massive

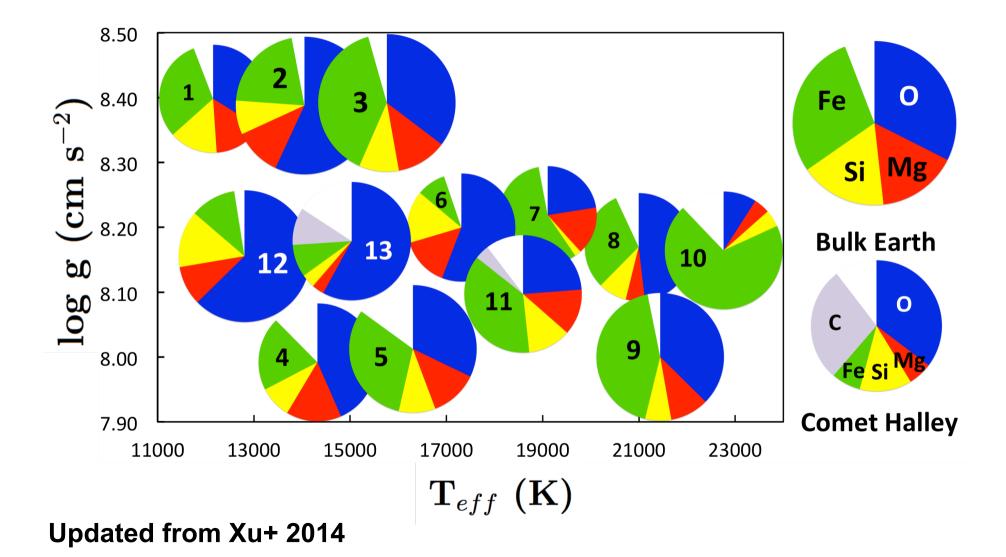


#### The Mass Fraction of Nitrogen is ~2%

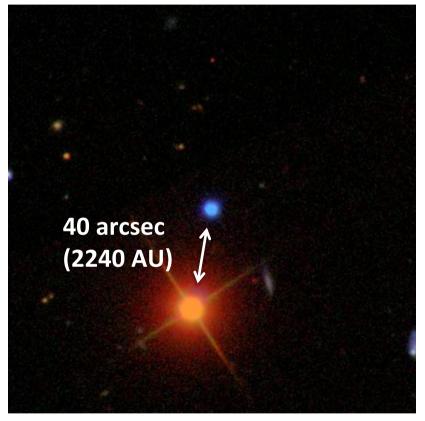


• Other polluted WDs

#### #13 The Most Carbon-rich Extrasolar Planetary Material



# **Binary Companion**

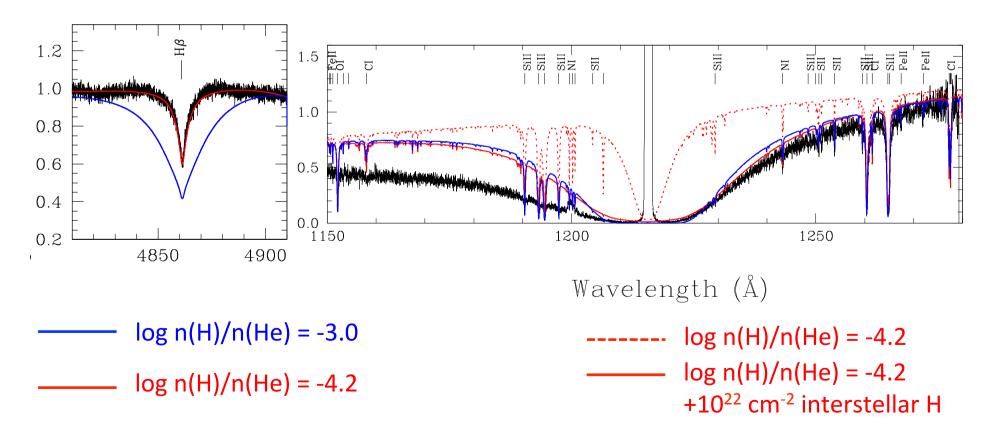


From SDSS

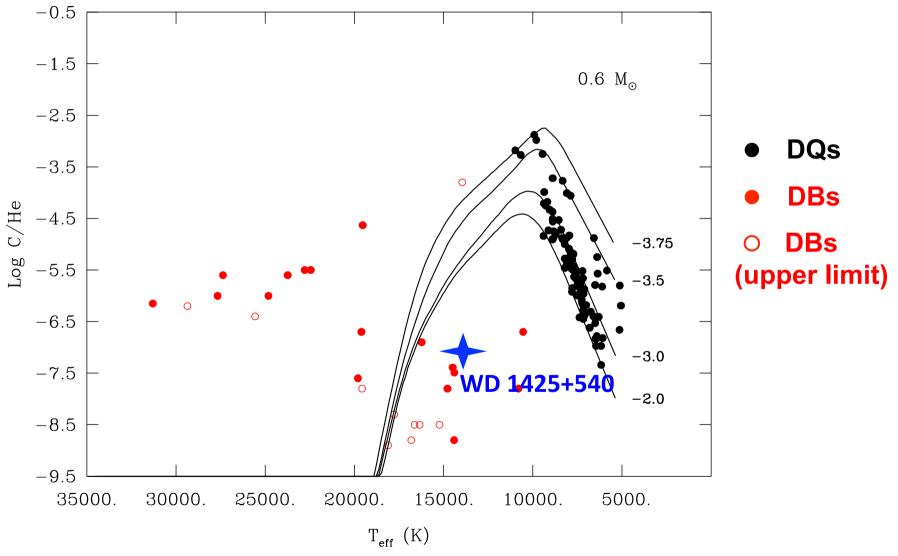
- WD 1425+540 has a K dwarf companion
- KBO analog originally located ~120 AU
- Galactic Tides (Bonsor & Veras 2015)? Kozai-Lidov Mechanism?

#### Future Work #1: Hydrogen Abundance

- Optical (H $\alpha$  & H $\beta$ ): log n(H)/n(He) = -4.2
- UV (Ly $\alpha$ ): log n(H)/n(He) = -3.0 (asymmetric)

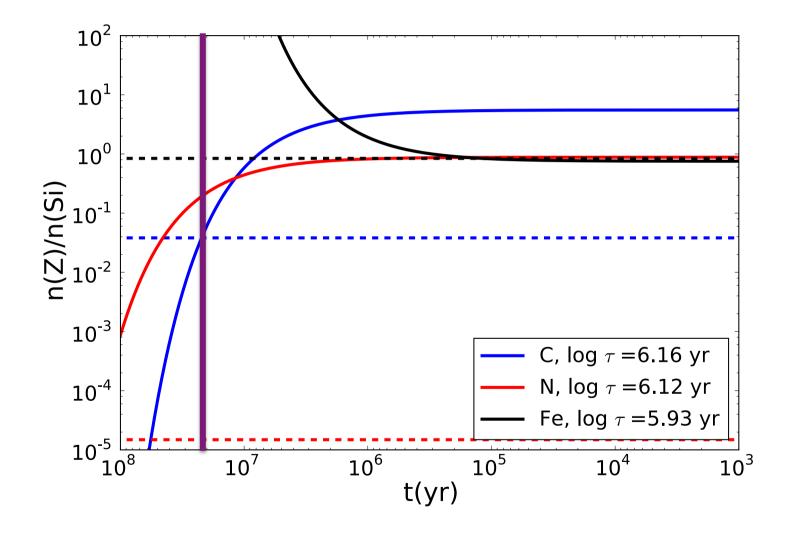


#### Future Work #2: Carbon Dredge-Up



From G. Fontaine

#### **Future Work #3: Accretion Stage**



For Si,  $\log \tau = 6.11$  yr

# Summary

- WD 1425+540 provides direct evidence for KBO analog around other stars
- Its chemical composition is very similar to that of comet Halley
- N is detected for the first time and it is enhanced
- A lot of future work...





2017, ApJL, 836, L7