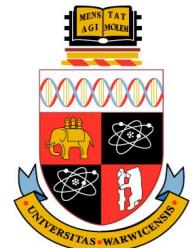




# Constraining the properties of the asteroid disintegrating around WD 1145+017



Dimitri Veras (University of Warwick)



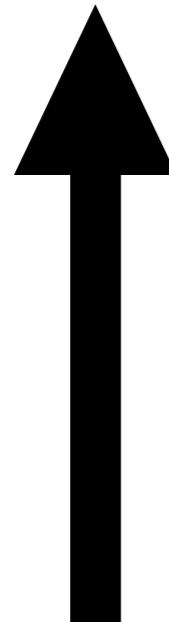
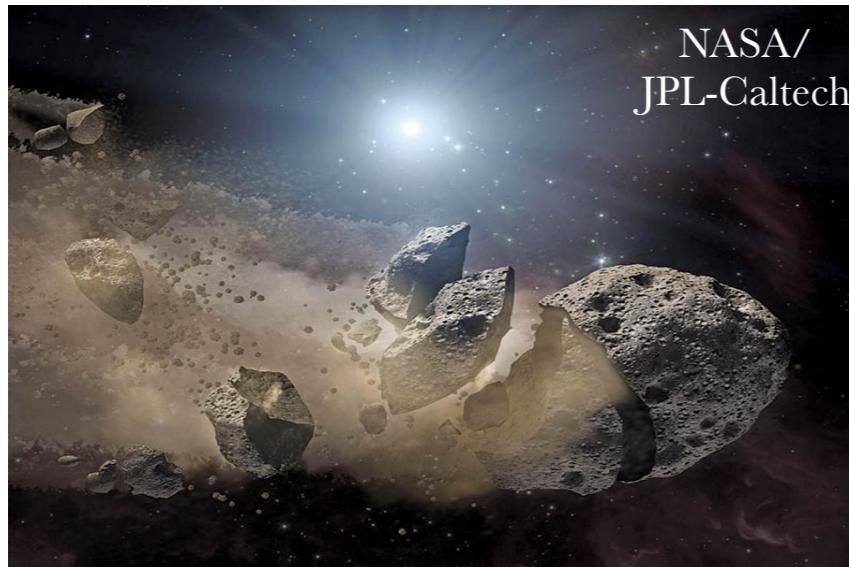
**break-up**



**form disc**



**pollute star**



**Missing  
Link** (before WD 1145+017)

# WD 1145+017 papers (so far)

## Observational

Vanderburg et al. (2015, *Nature*, 526, 546-549) | 2015

Alonso et al. (2016, *A&A*, 589, L6)

Gänsicke et al. (2016, *ApJL*, 818, L7)

Rappaport et al. (2016, *MNRAS*, 458, 3904)

Xu et al. (2016, *ApJL*, 816, L22)

Zhou et al. (2016, *MNRAS*, 463, 4422)

2016

Croll et al. (2017, *ApJ*, 836, 82)

Gary et al. (2017, *MNRAS*, 465, 3267)

Hallakoun et al. (2017, Submitted to *MNRAS*)

Redfield et al. (2017, Submitted to *ApJ*)

2017

## Theoretical

Gurri et al. (2017, *MNRAS*, 464, 321)

Veras et al. (2017, *MNRAS*, 465, 1008)

# WD 1145+017: other talks

Tuesday 11:10

Rik van Lieshout

Tuesday 12:00

Mark Wyatt

Thursday 16:40

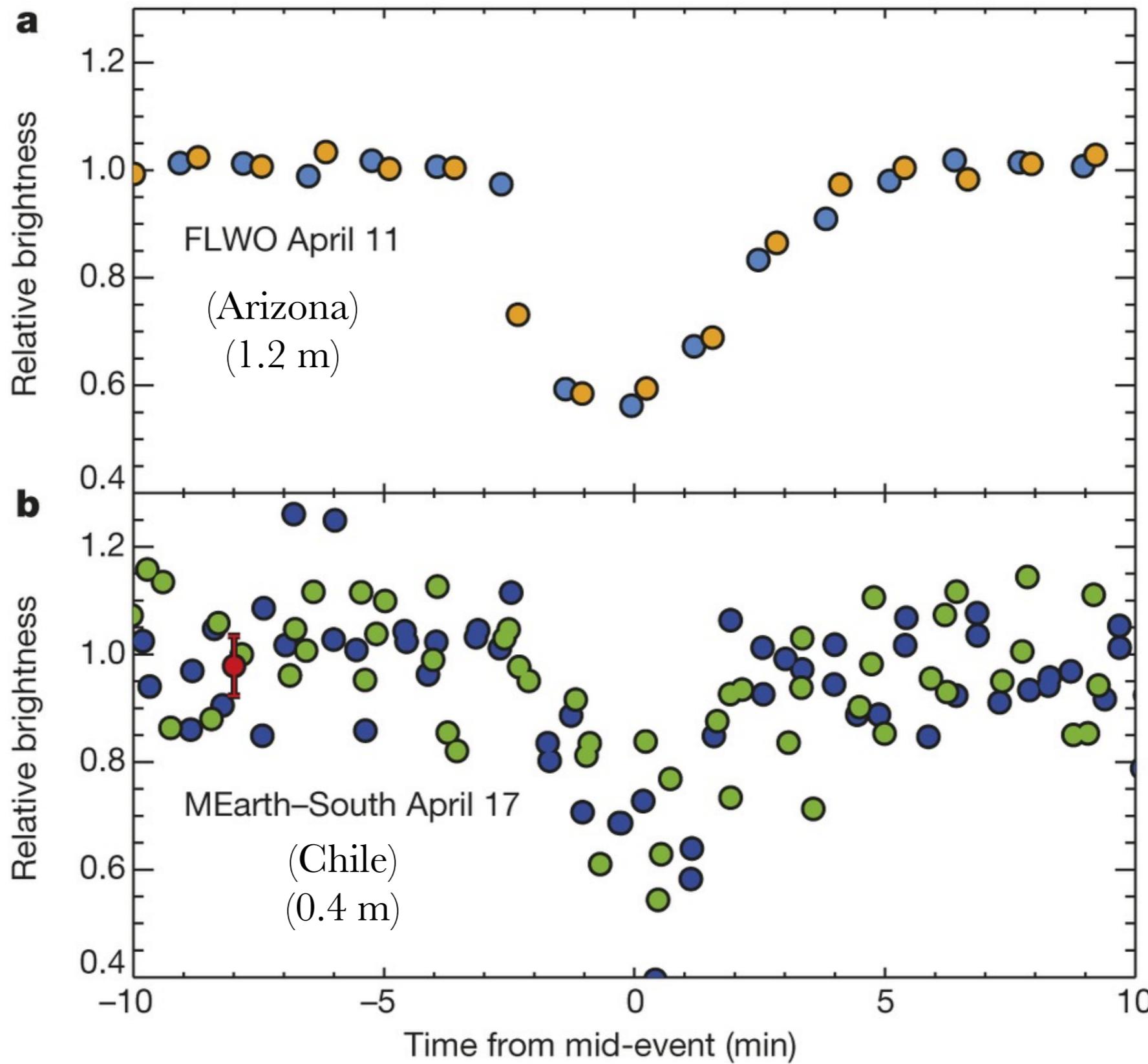
Na'ama Hallakoun

Friday 9:20

John Lewis

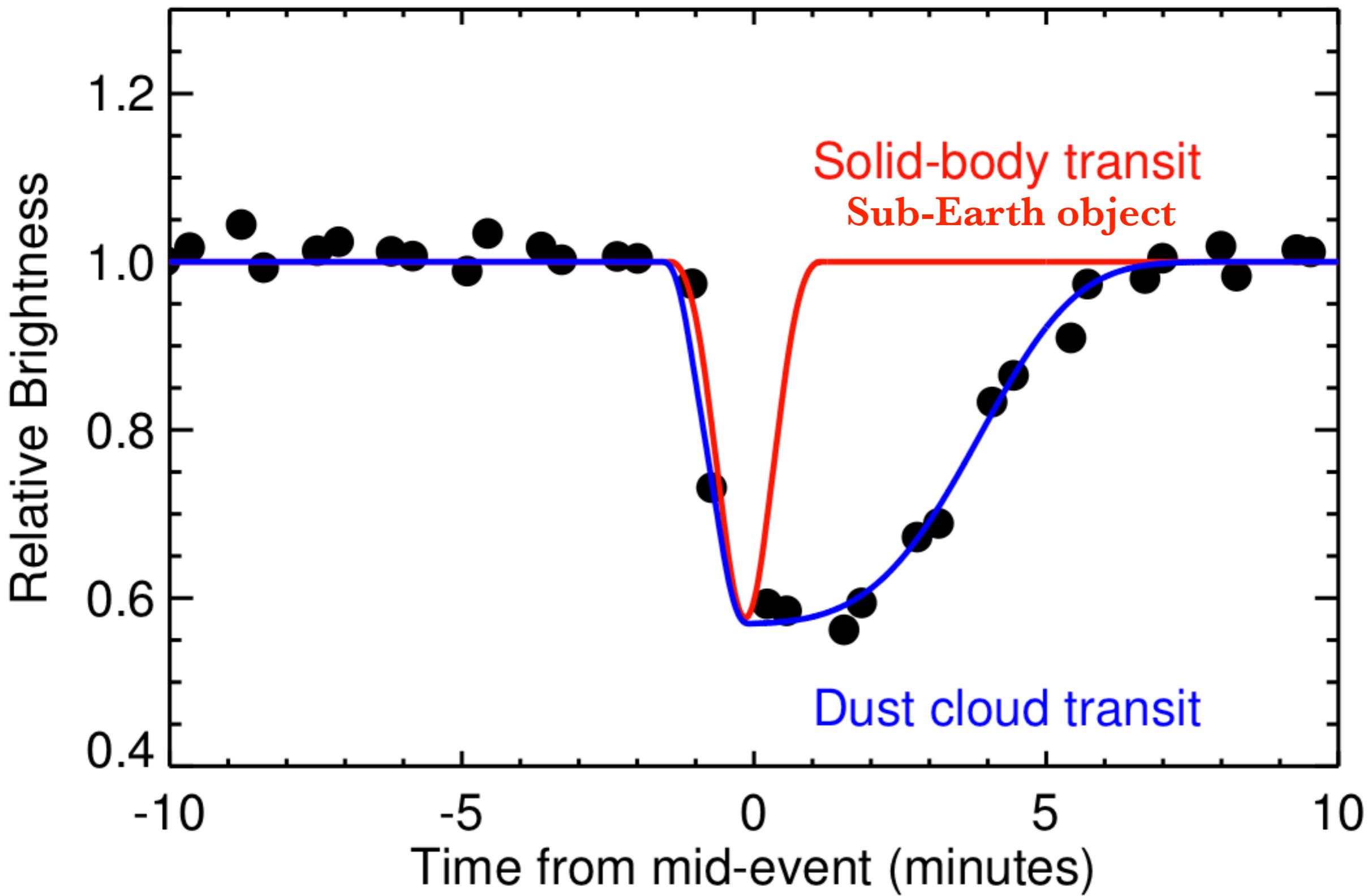
# Basic observational background

Vanderburg et al. (2015, Nature, 526, 546-549)



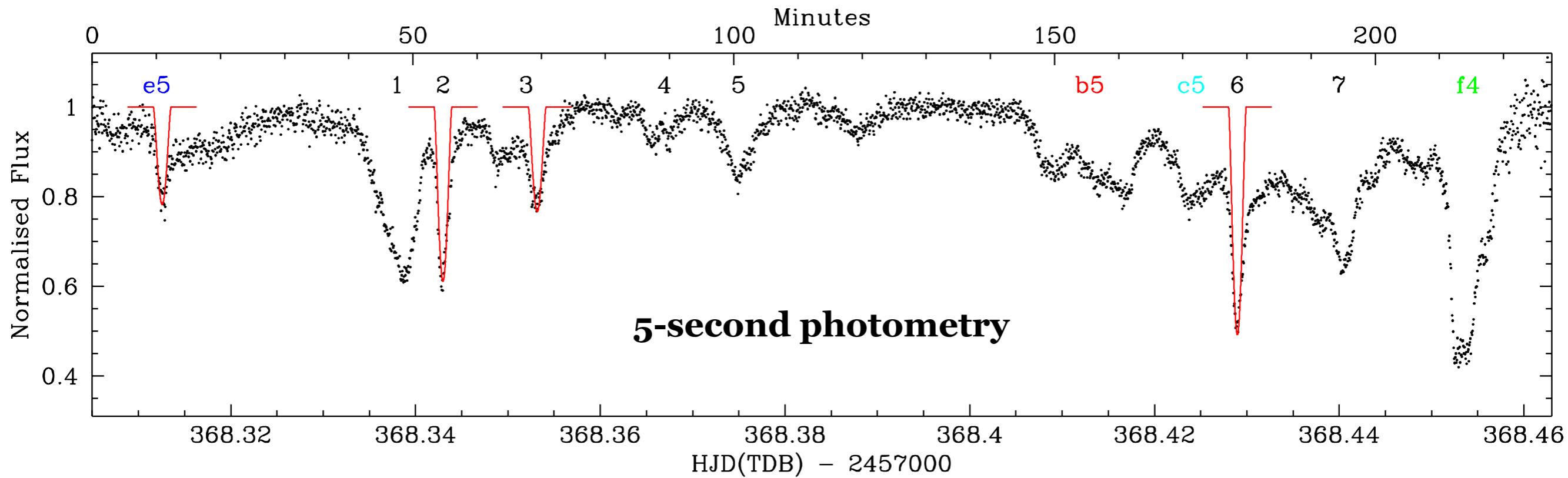
# Basic observational background

Vanderburg et al. (2015, Nature, 526, 546-549)



# Basic observational background

Gänsicke et al. (2016, ApJL, 818, L7)

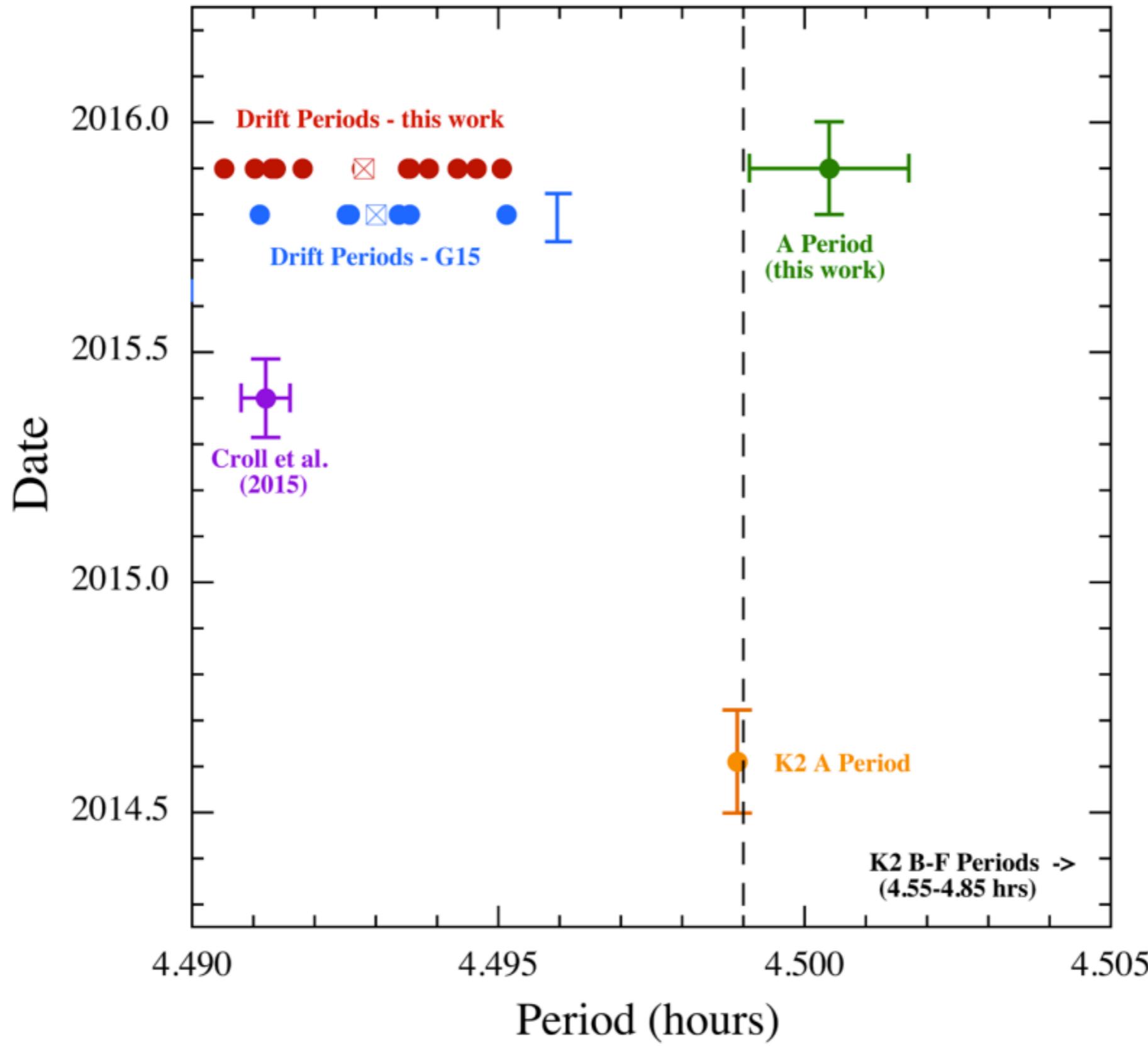


(Thailand)  
(2.4 m)

(Spain)  
(1.0 m)

# Basic observational background

Rappaport et al. (2016, MNRAS, 458, 3904)



# Explaining observables

Veras et al. (2017, MNRAS, 465, 1008)

## Collaborators

Philip J. Carter

Zoë M. Leinhardt

Boris T. Gänsicke

## Constraints

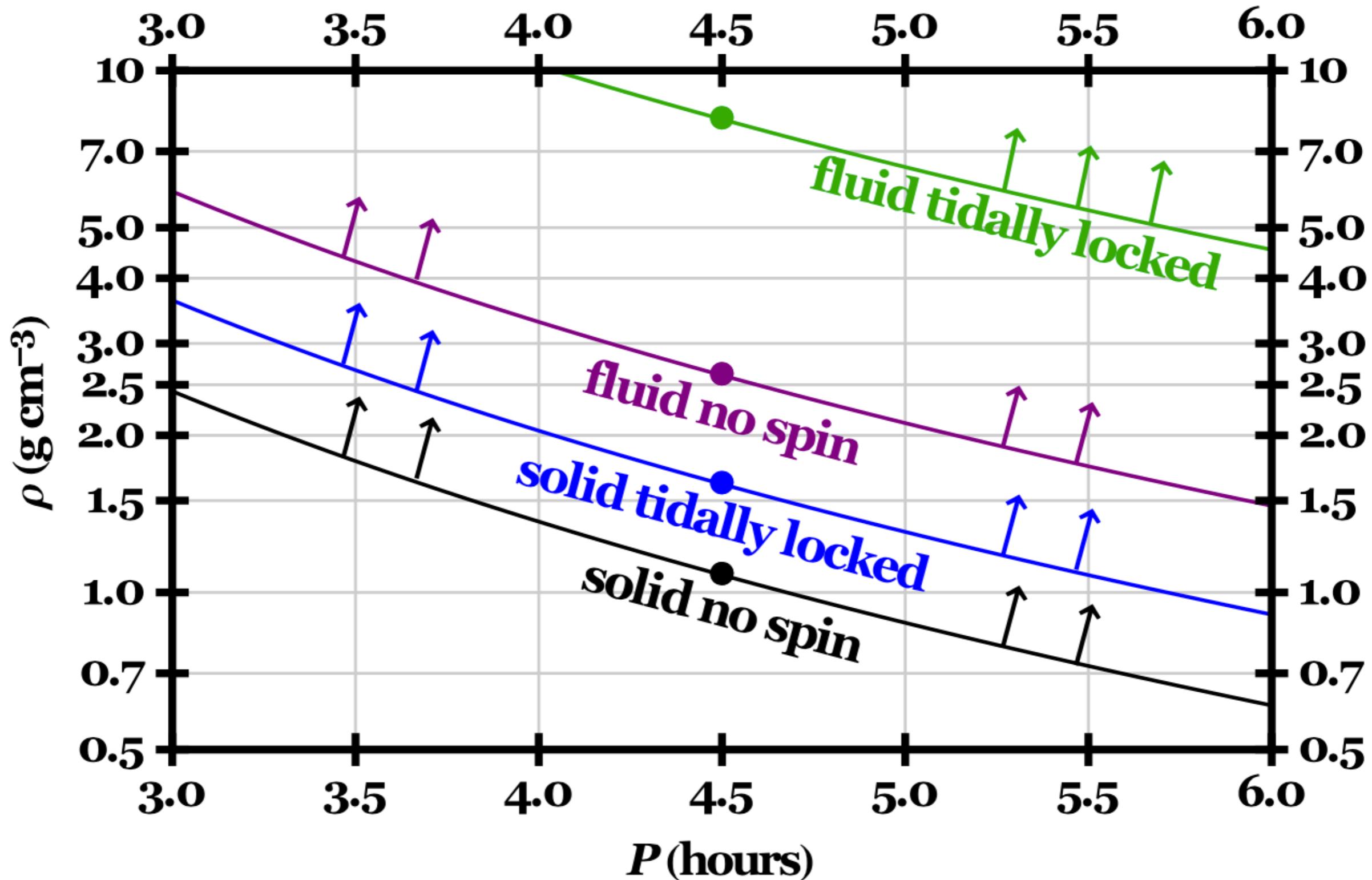
Disruption persists for 2 years

Disruption intermittent

Orbital period of  $\sim$ 4.499 hours

# Stable to disruption

Veras et al. (2017, MNRAS, 465, 1008)



# Rubble pile representations

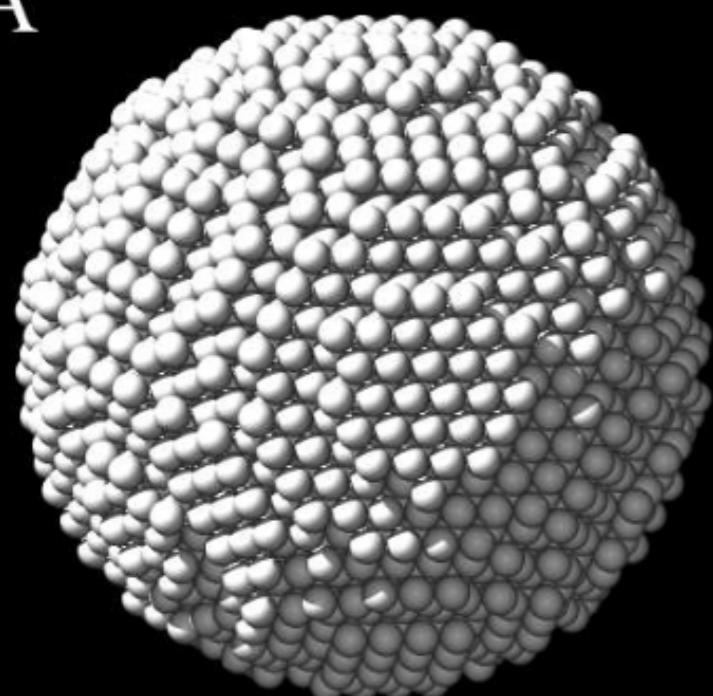
Veras et al. (2017, MNRAS, 465, 1008)

**Homogeneous**  
Hexagonal  
Closest  
Packing

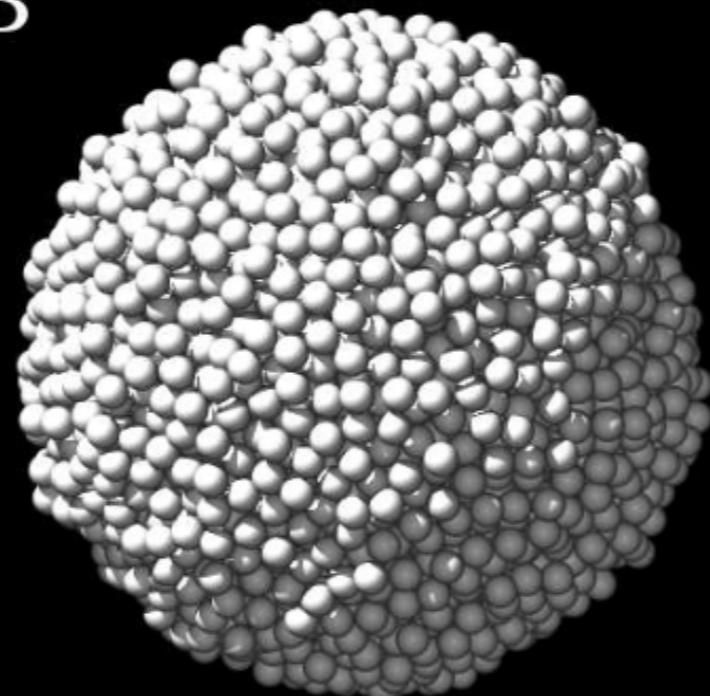
**Homogeneous**  
Random  
Packing

**Differentiated**  
Core  
+  
Mantle

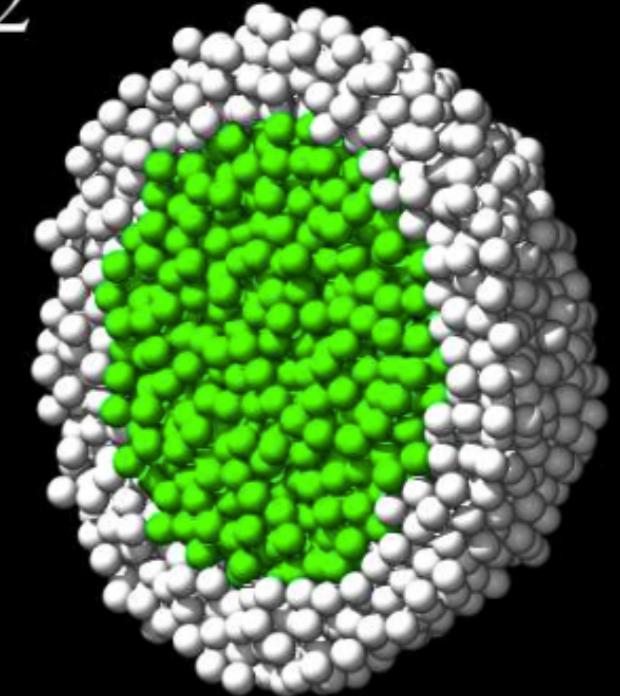
A



B



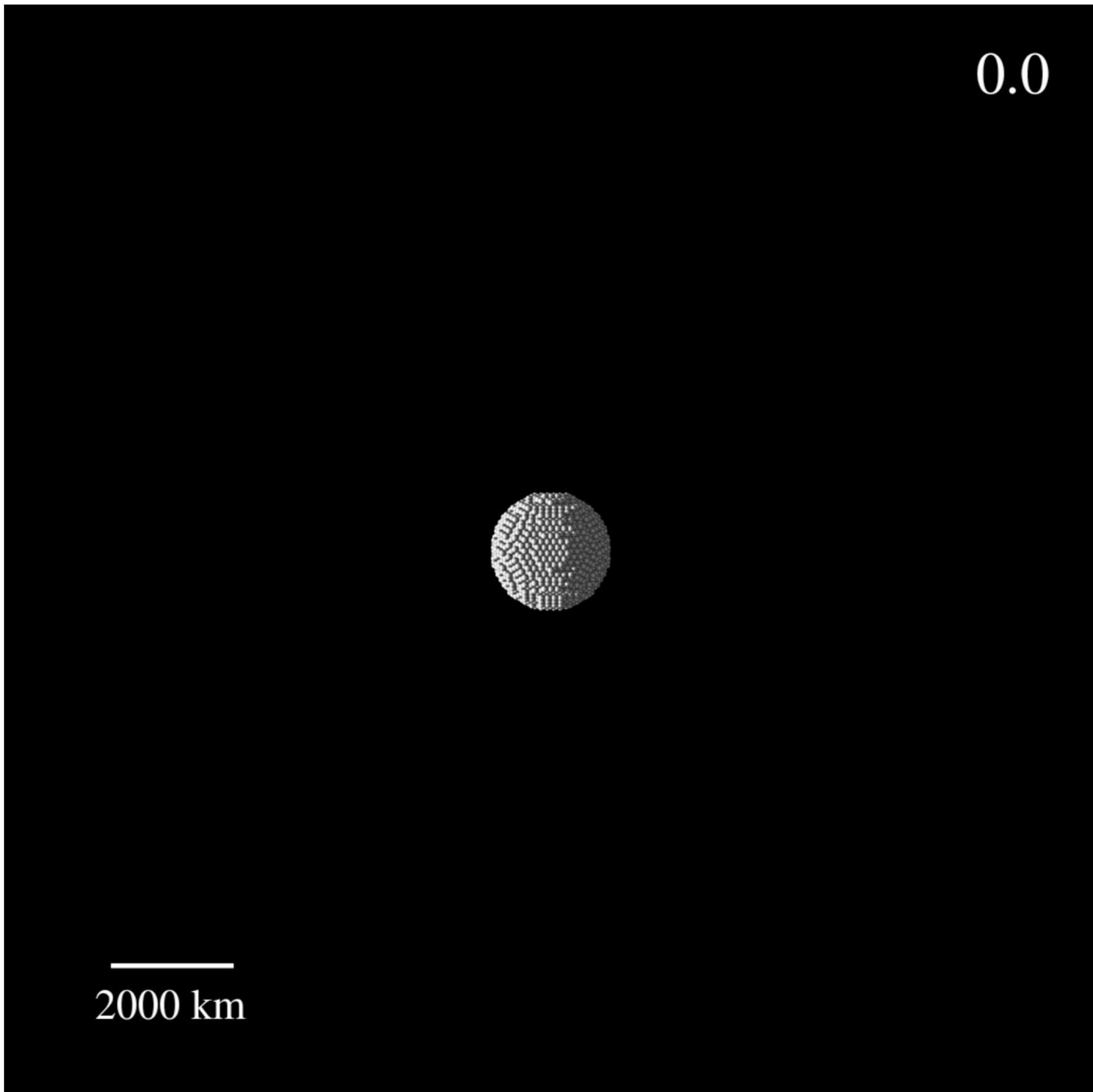
B2



1000 km

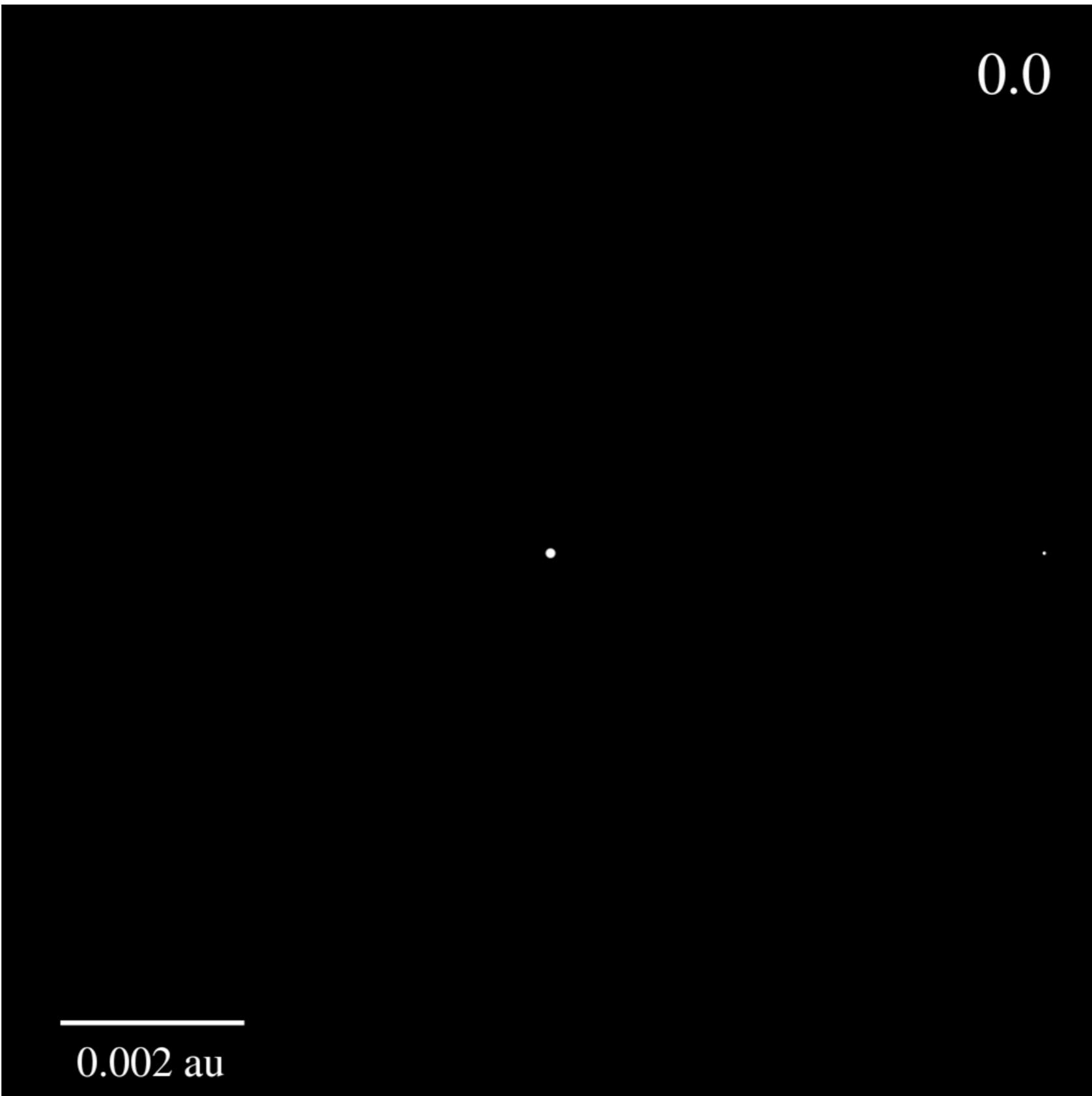
# Total disruption

Veras et al. (2017, MNRAS, 465, 1008)



# Ring / disc formation

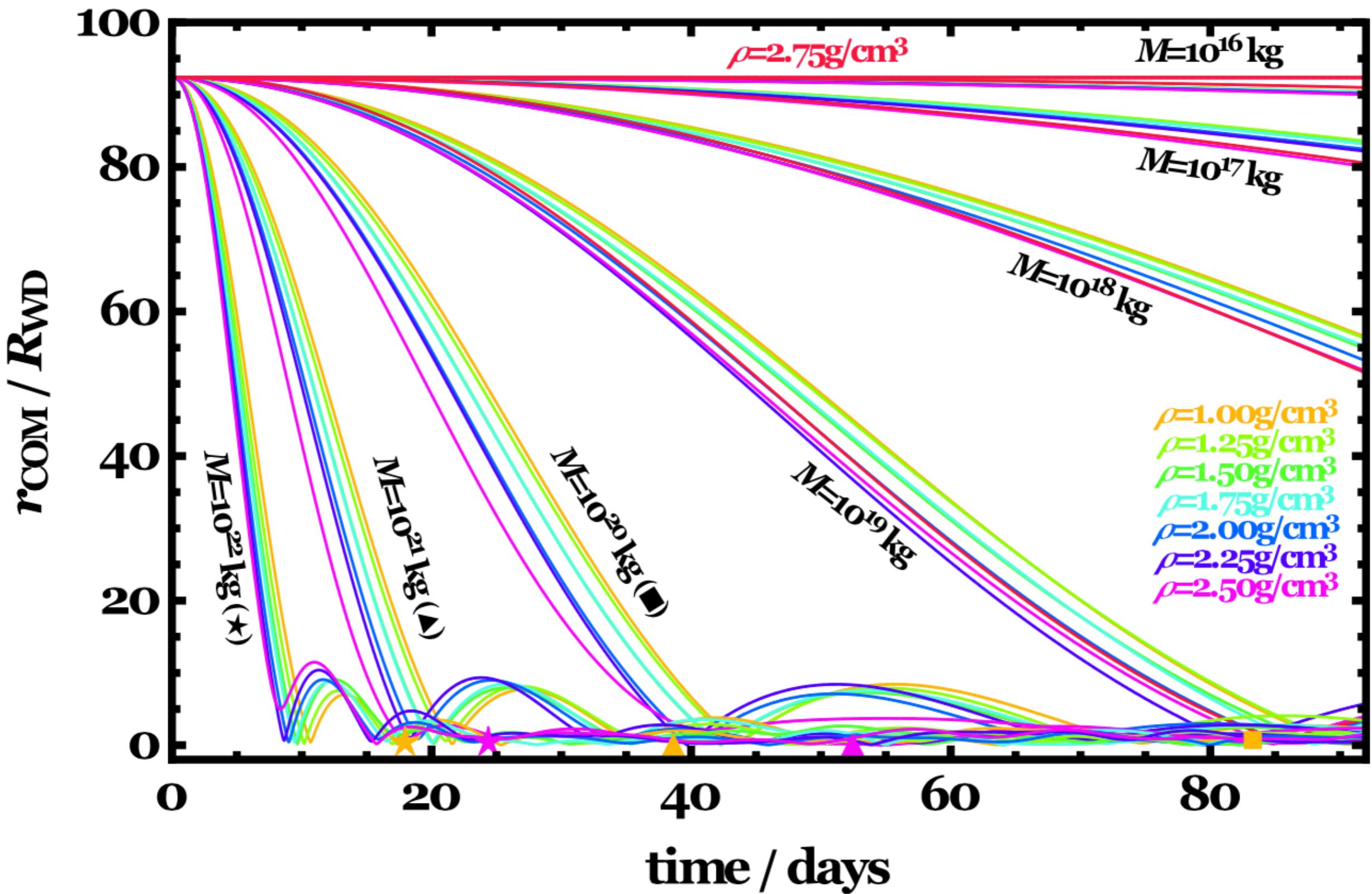
Veras et al. (2017, MNRAS, 465, 1008)



# Ring / disc formation timescales

Veras et al. (2017, MNRAS, 465, 1008)

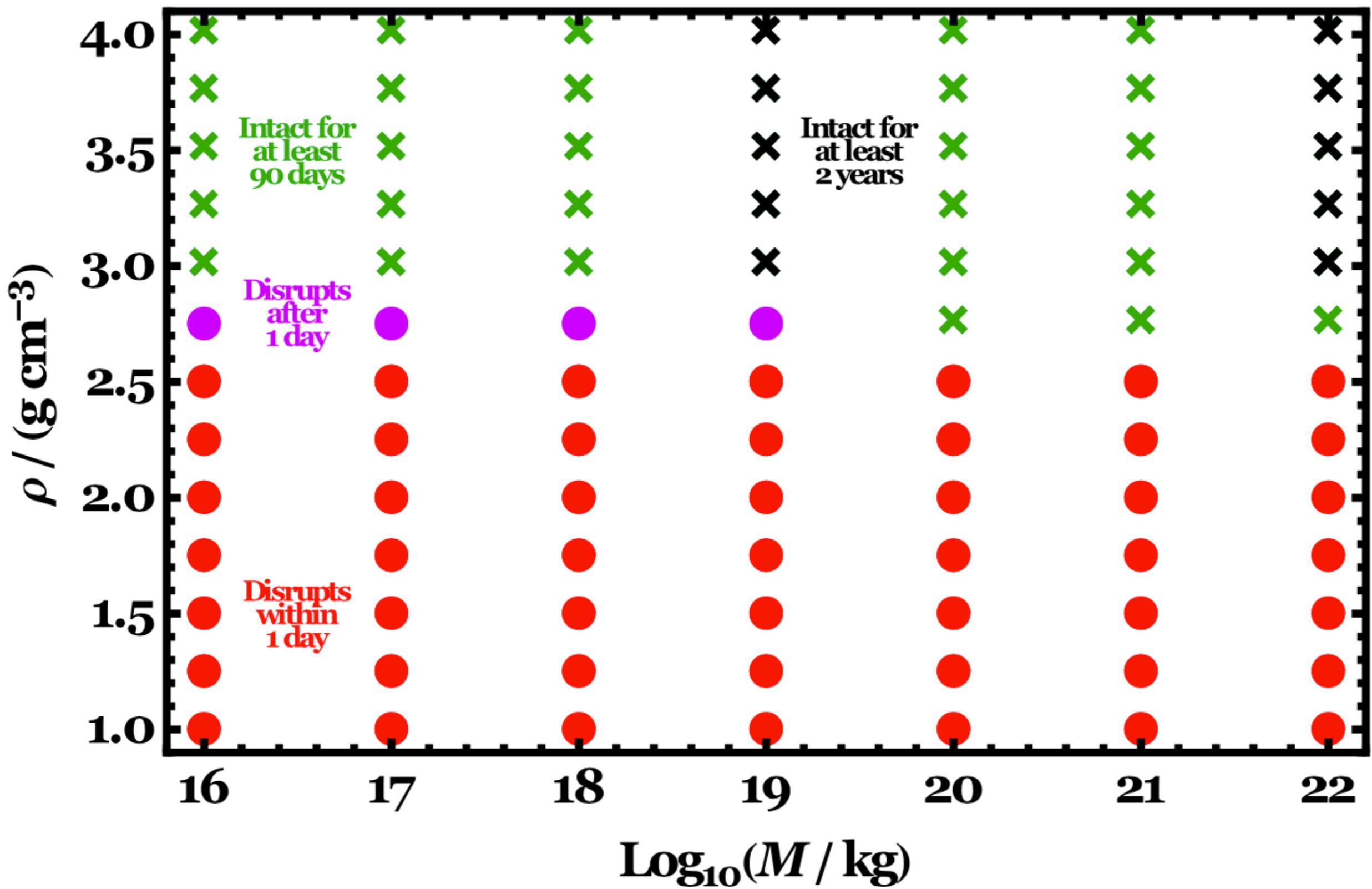
filling out uniform rings



# Disruption timescales

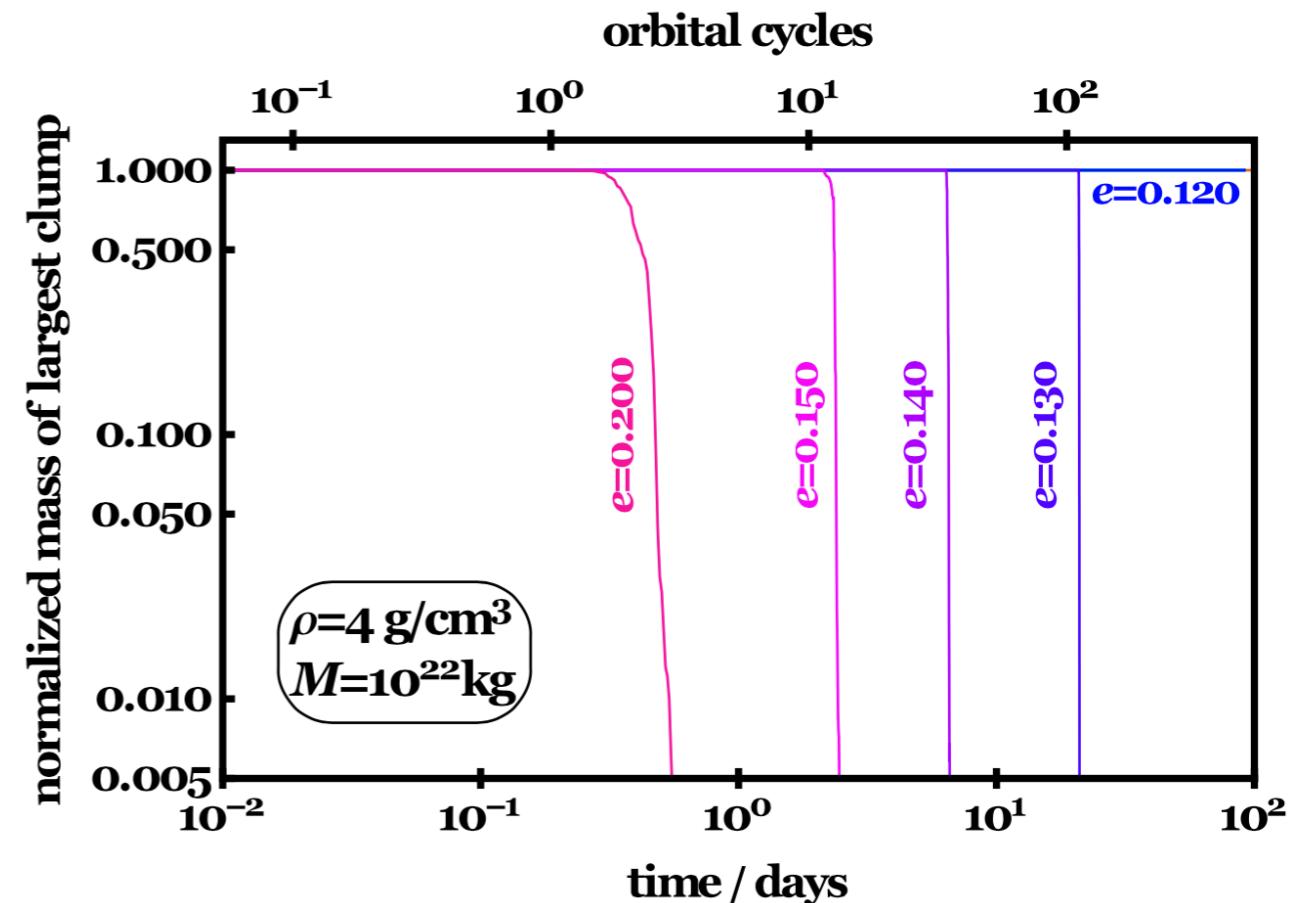
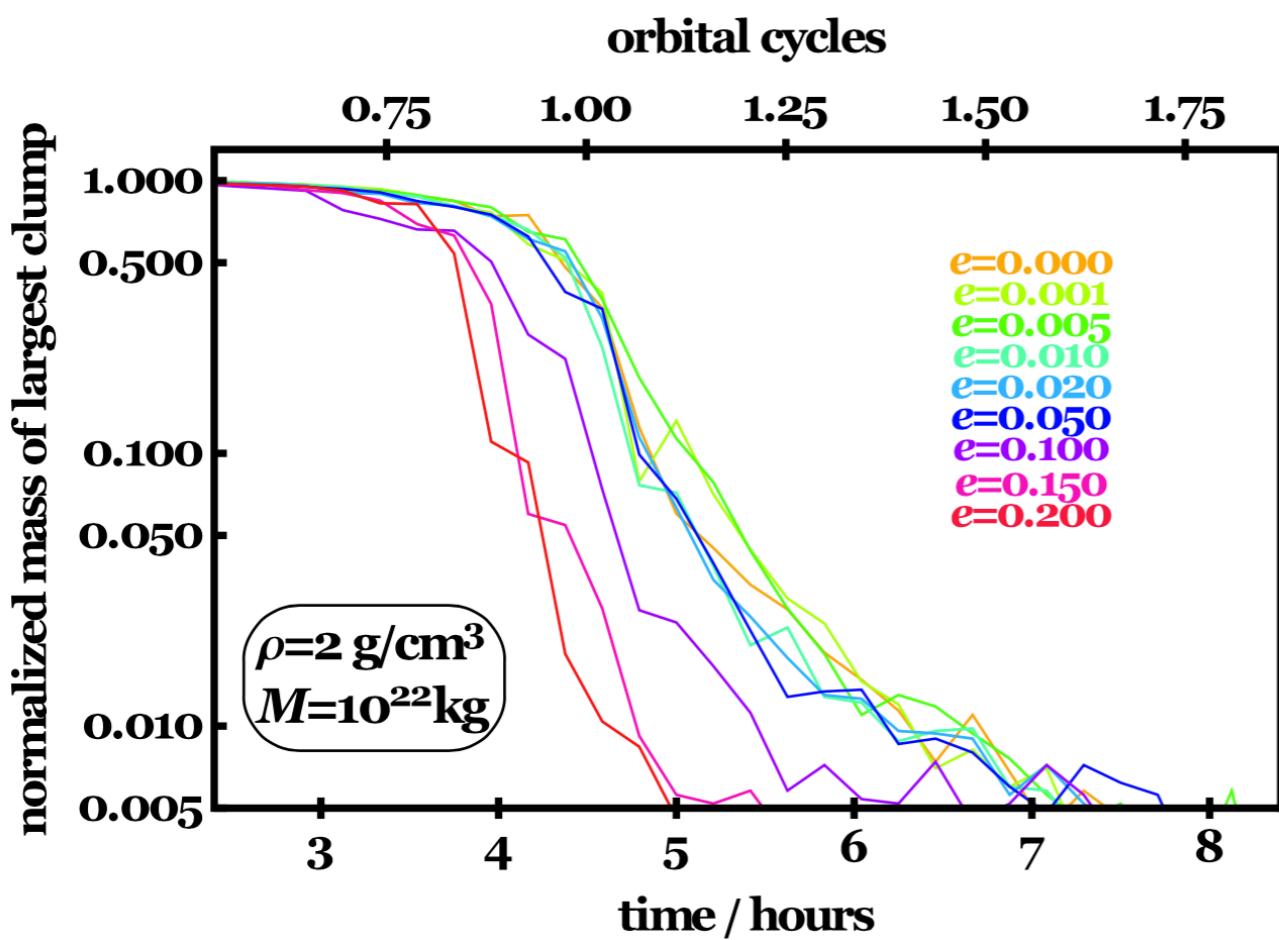
Veras et al. (2017, MNRAS, 465, 1008)

## Disruption times



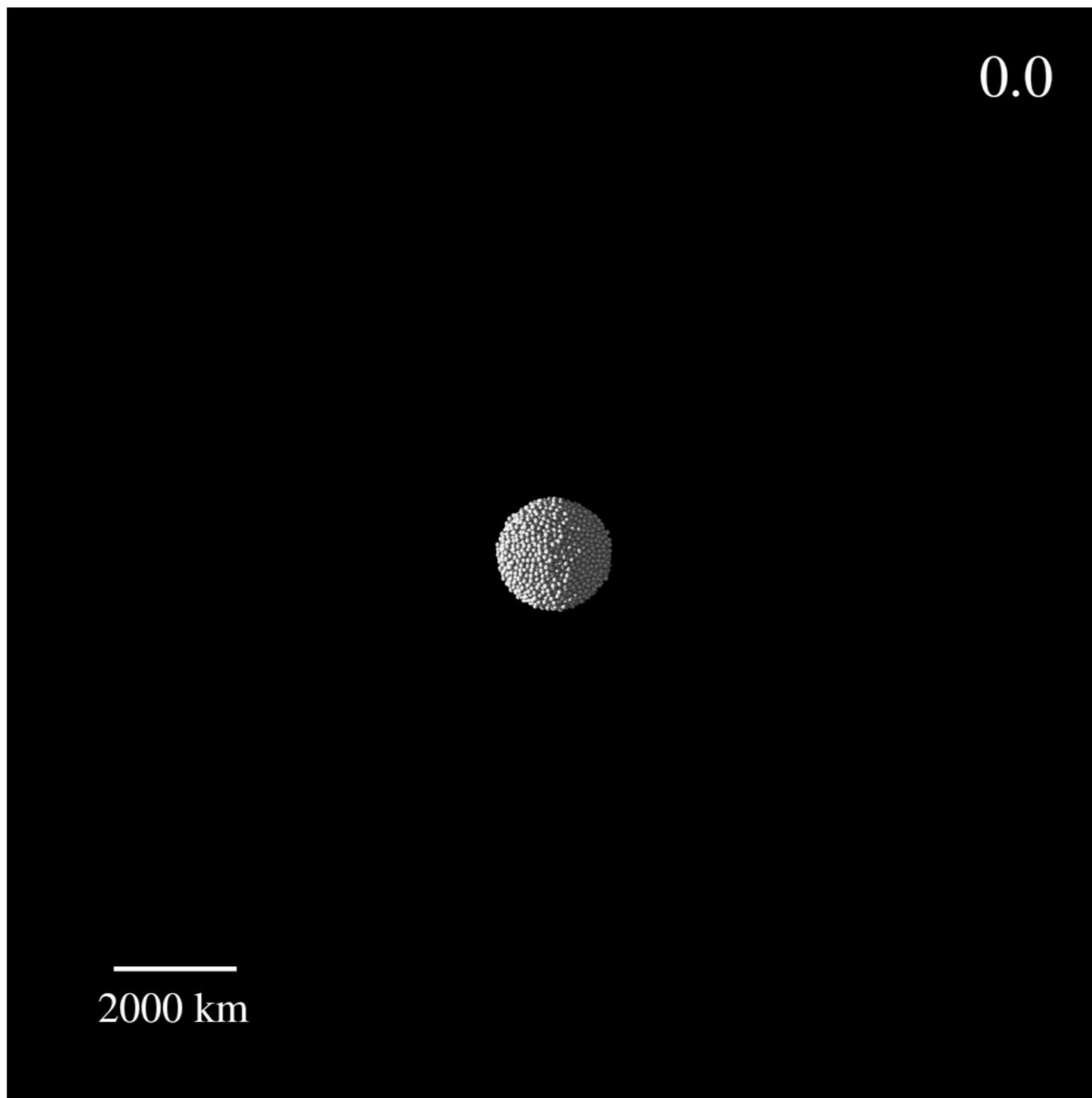
# Disruption timescales

Veras et al. (2017, MNRAS, 465, 1008)



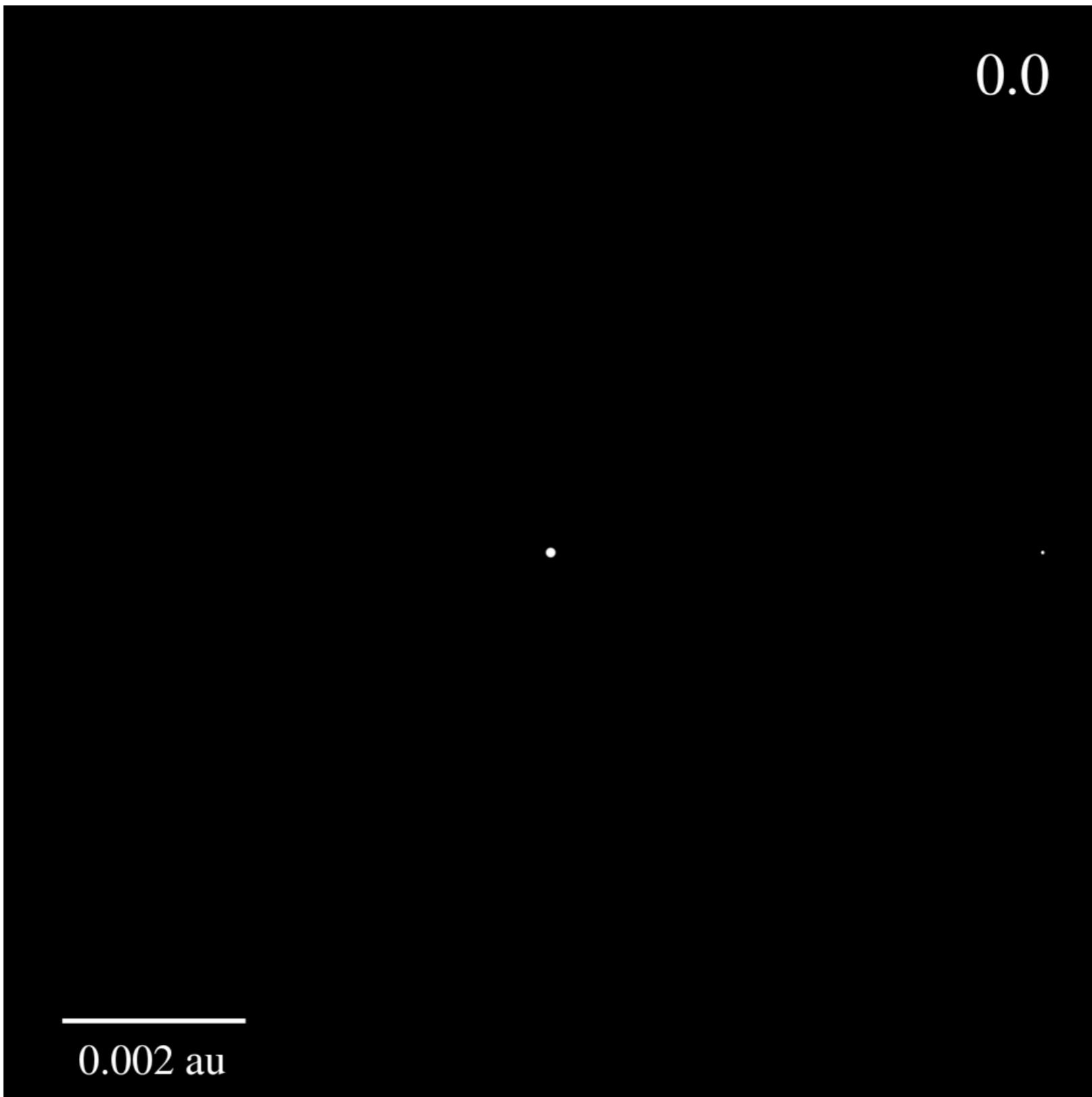
# Differentiated: Total disruption

Veras et al. (2017, MNRAS, 465, 1008)



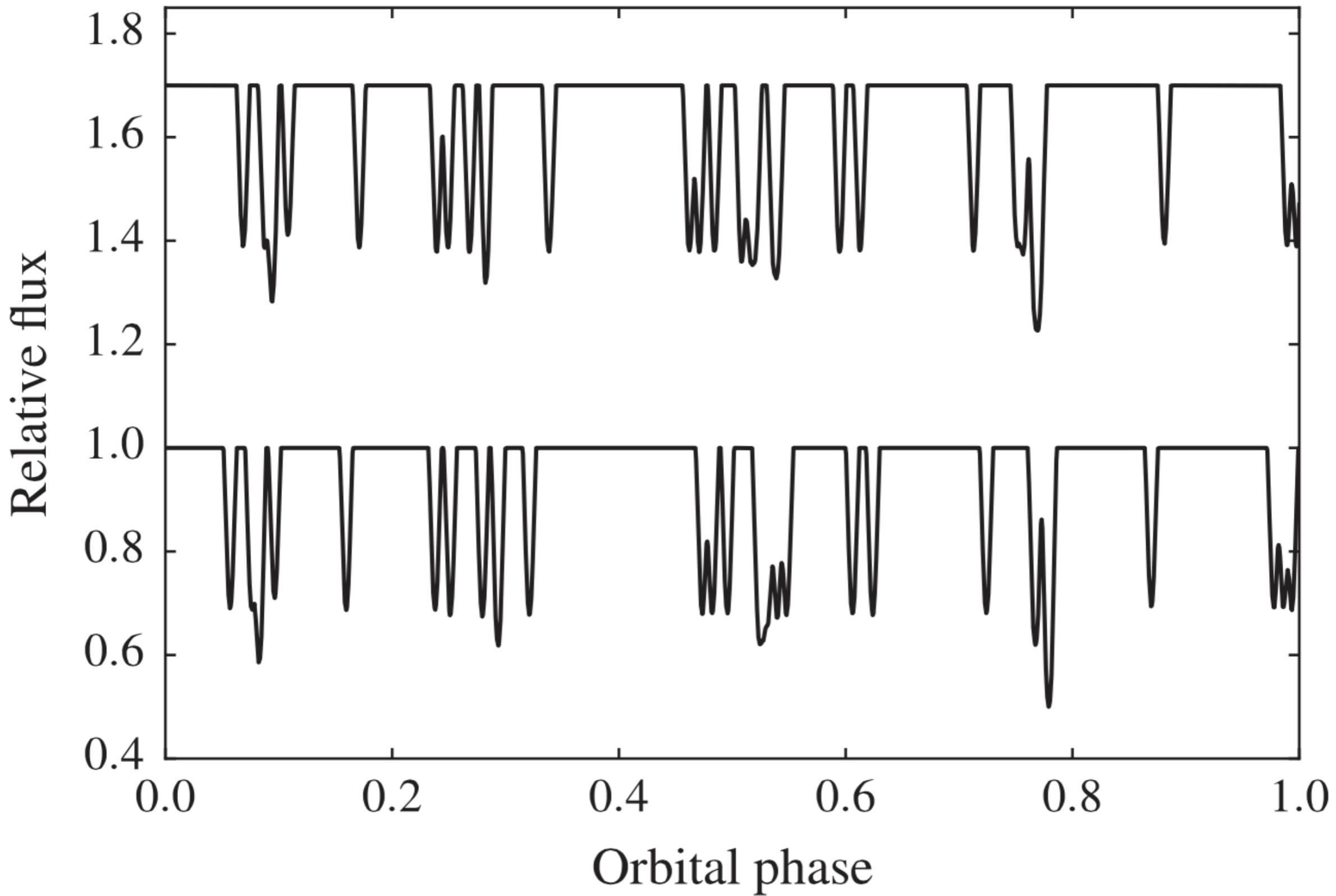
# Differentiated: Partial disruption

Veras et al. (2017, MNRAS, 465, 1008)



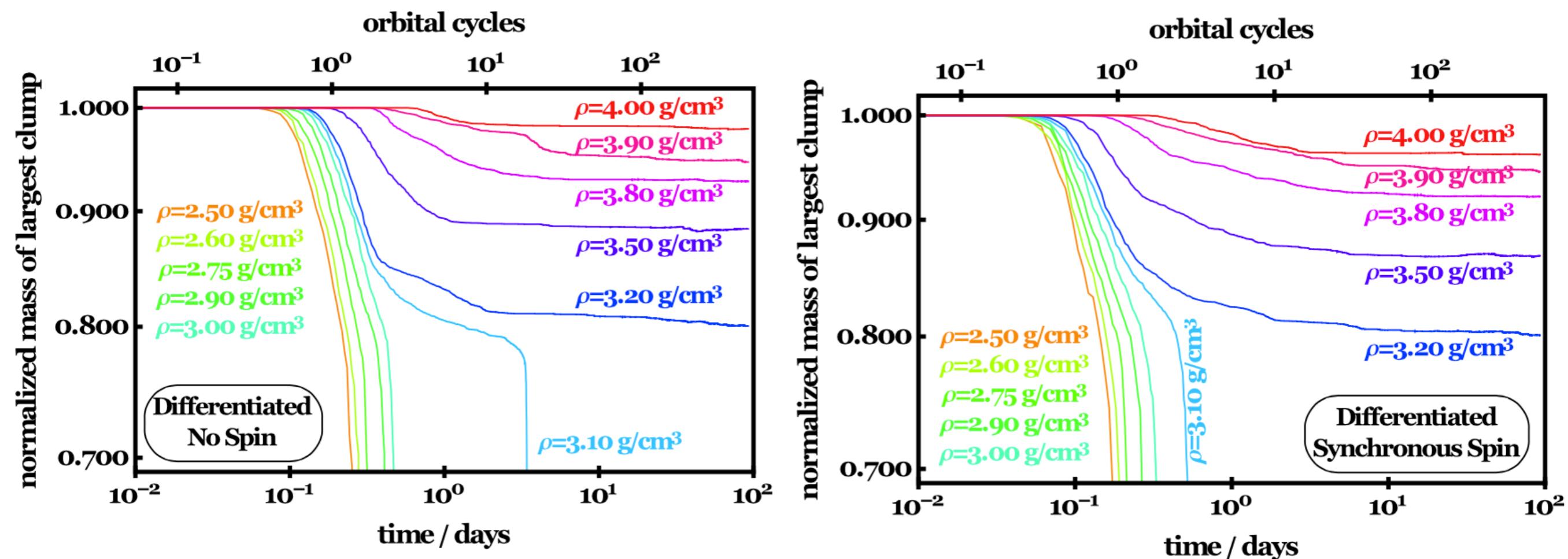
# Resulting transit curves

Veras et al. (2017, MNRAS, 465, 1008)



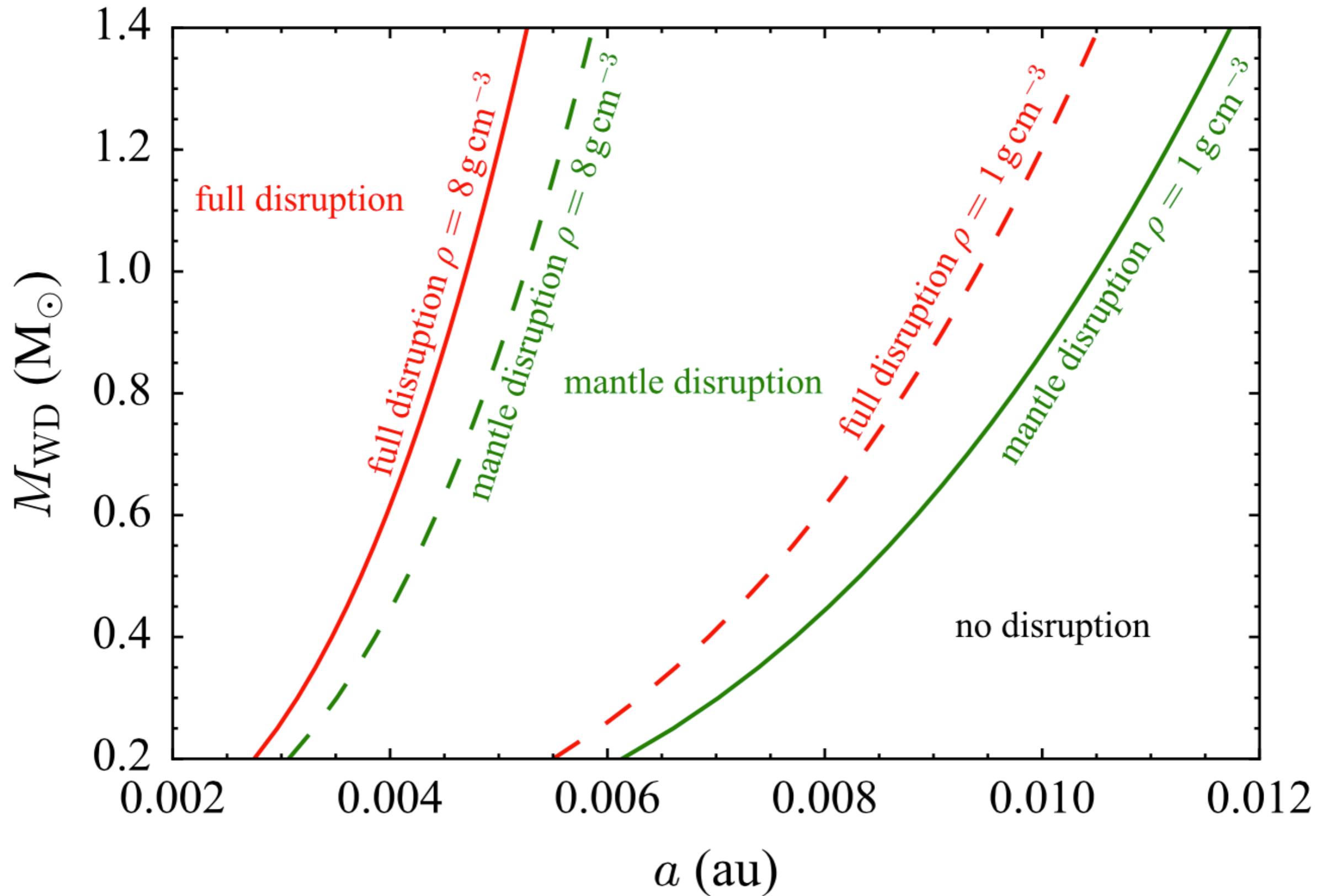
# Partial disruption characteristics

Veras et al. (2017, MNRAS, 465, 1008)



# Partial disruption characteristics

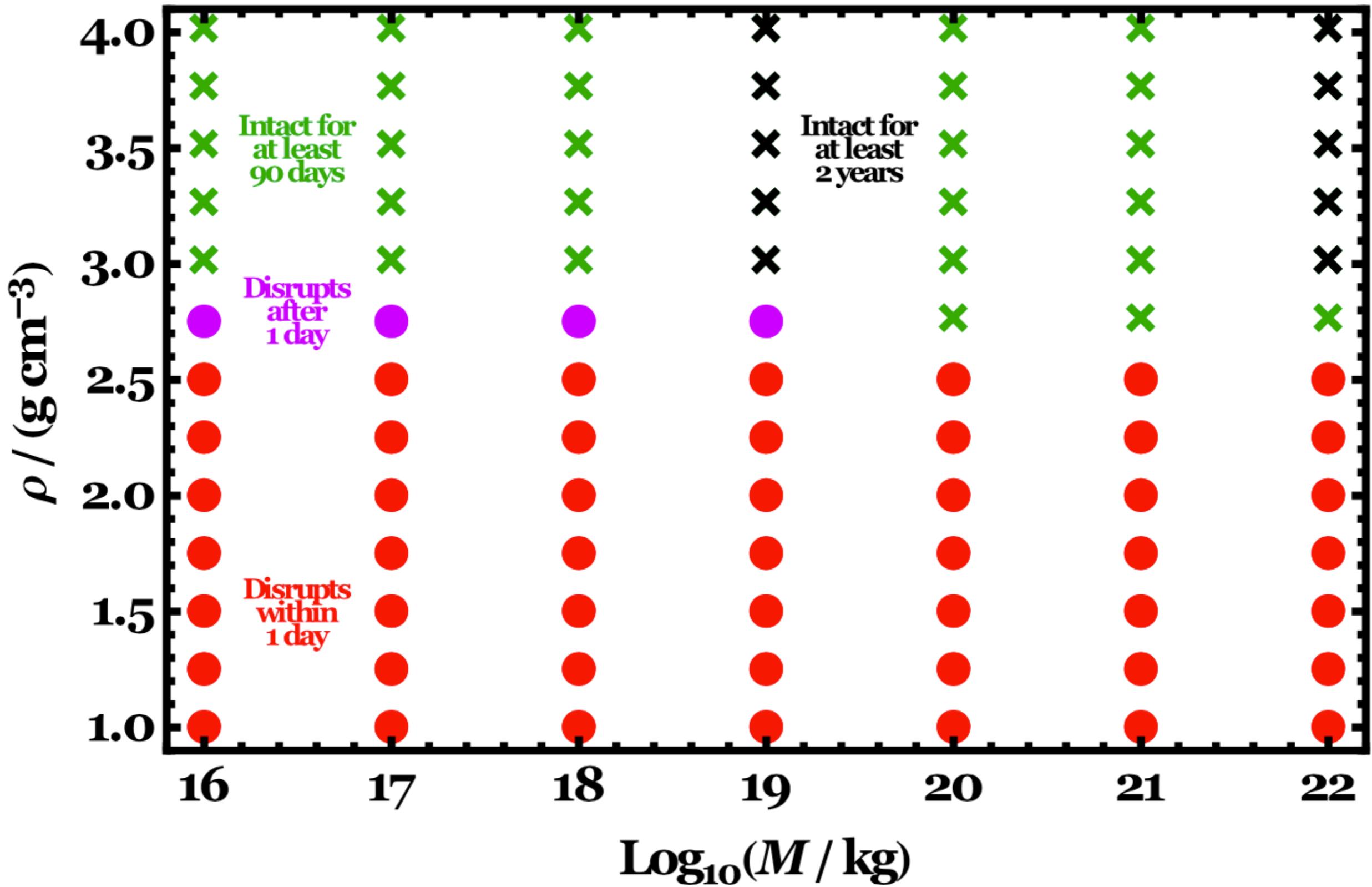
Veras et al. (2017, MNRAS, 465, 1008)



# Poor constraint on mass

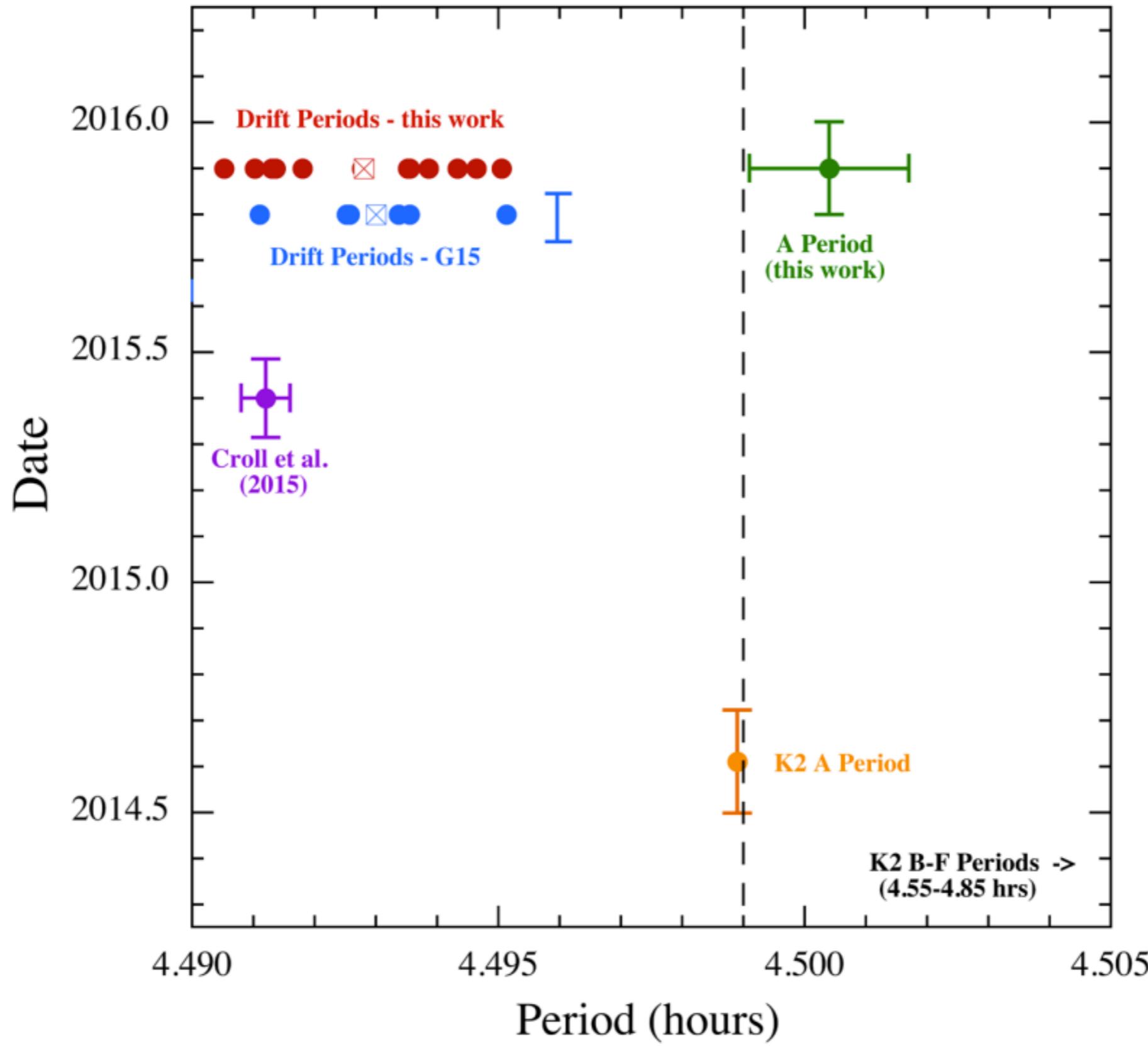
Veras et al. (2017, MNRAS, 465, 1008)

## Disruption times



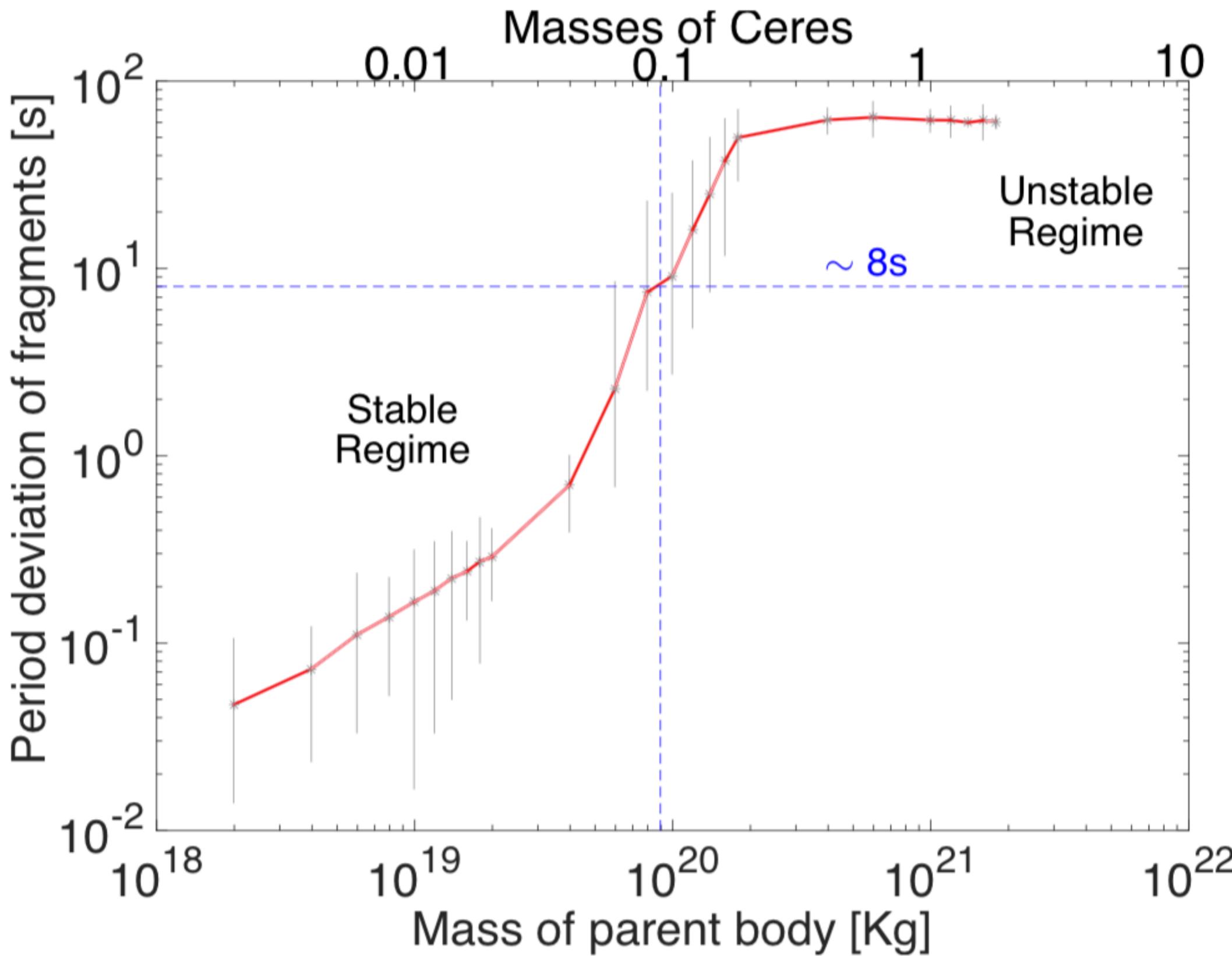
# Use other repeating features

Rappaport et al. (2016, MNRAS, 458, 3904)



# Constraining mass

Gurri et al. (2017, MNRAS, 464, 321)



# Conclusions

Asteroid disintegrating around WD 1145+017:

- Differentiated, not homogeneous
- Eccentricity < 0.01
- Bulk density 3-4 g/cm<sup>3</sup>
- Mass no greater than 10<sup>20</sup> kg