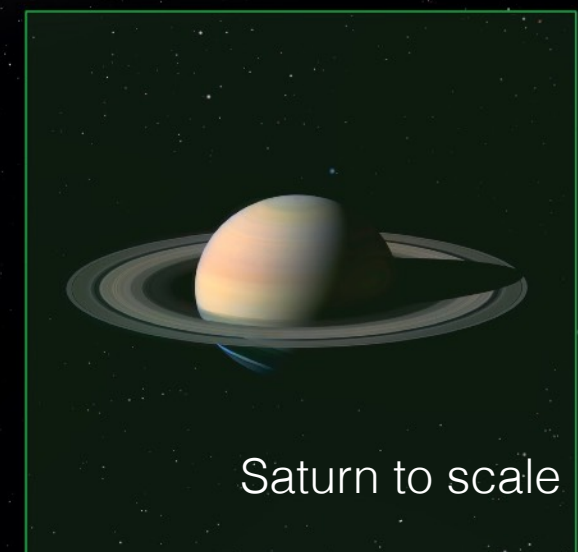
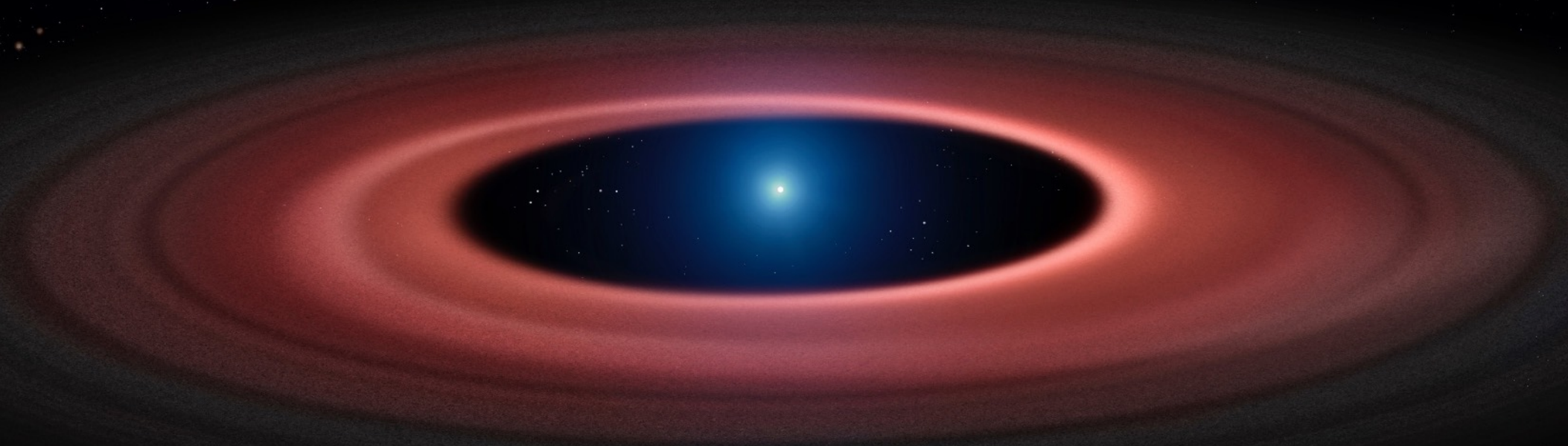


Gaseous debris discs around white dwarfs

Christopher J. Manser

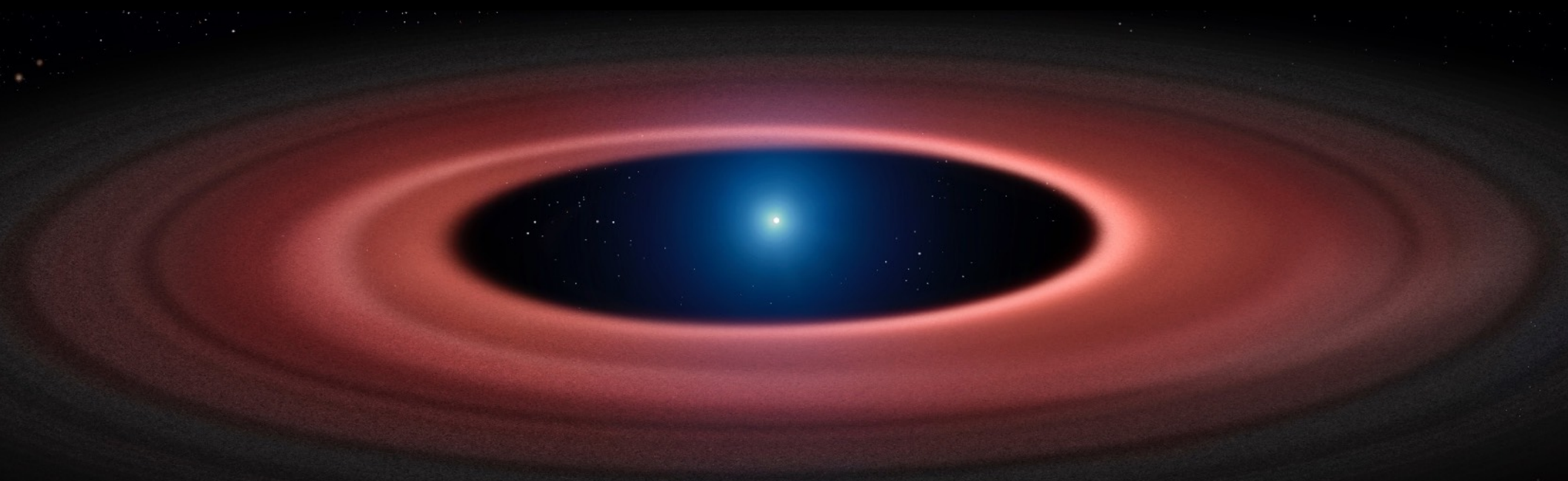
Collaborators: Boris Gänsicke, Tom Marsh, Detlev Koester,
Dimitri Veras, Nicola Pietro Gentile Fusillo



C.Manser@Warwick.ac.uk

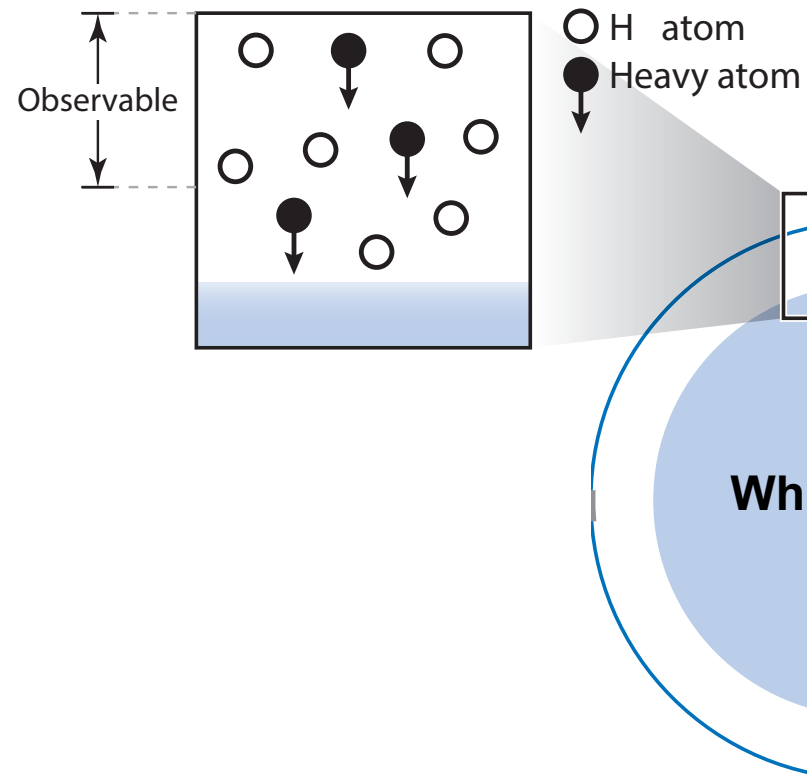
Talk Outline

- **One** - The gaseous debris disc around SDSS J1228+1040
- **Few** - Common variability of gaseous debris discs
- **Many** - Frequency of gaseous debris discs around white dwarfs



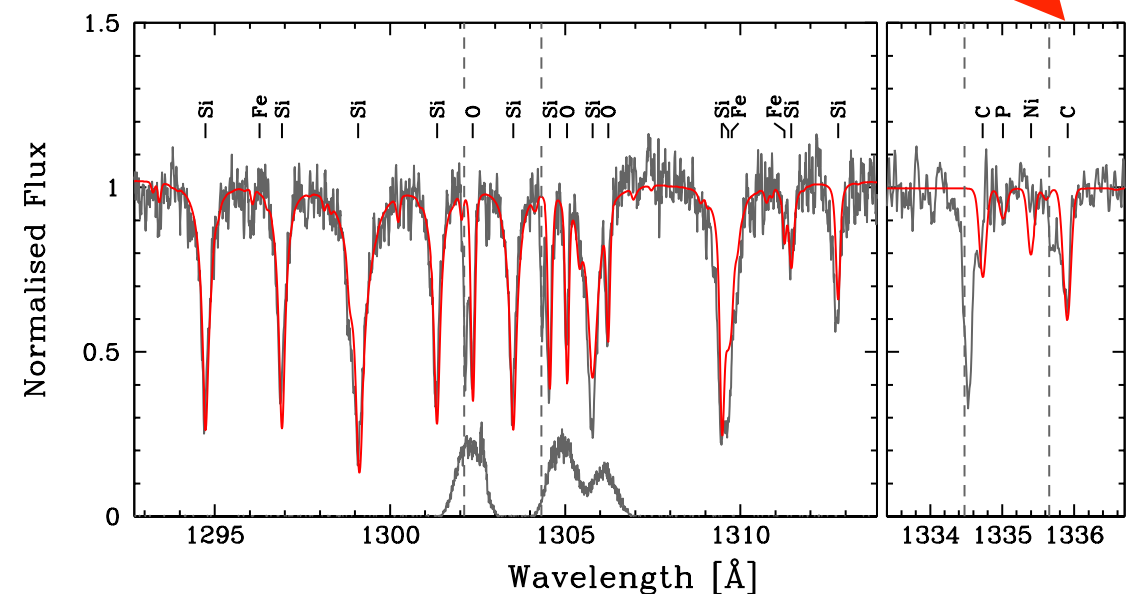
Remnant Planetary Systems

SDSS J1228+1040



Jura & Young, 2014, Annu.
Rev. Earth Planet. Sci., 42, 45

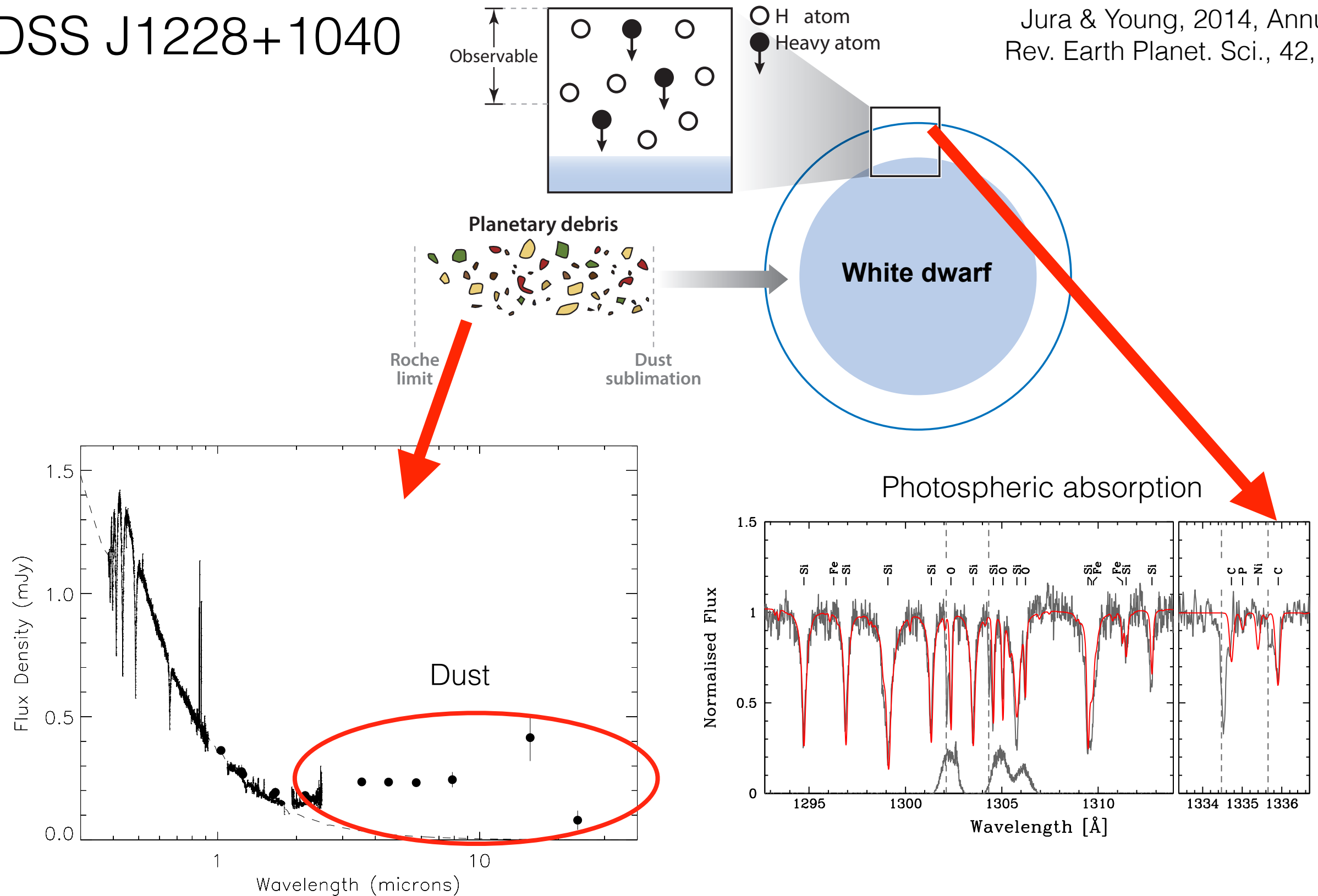
Photospheric absorption



Remnant Planetary Systems

SDSS J1228+1040

Jura & Young, 2014, Annu. Rev. Earth Planet. Sci., 42, 45



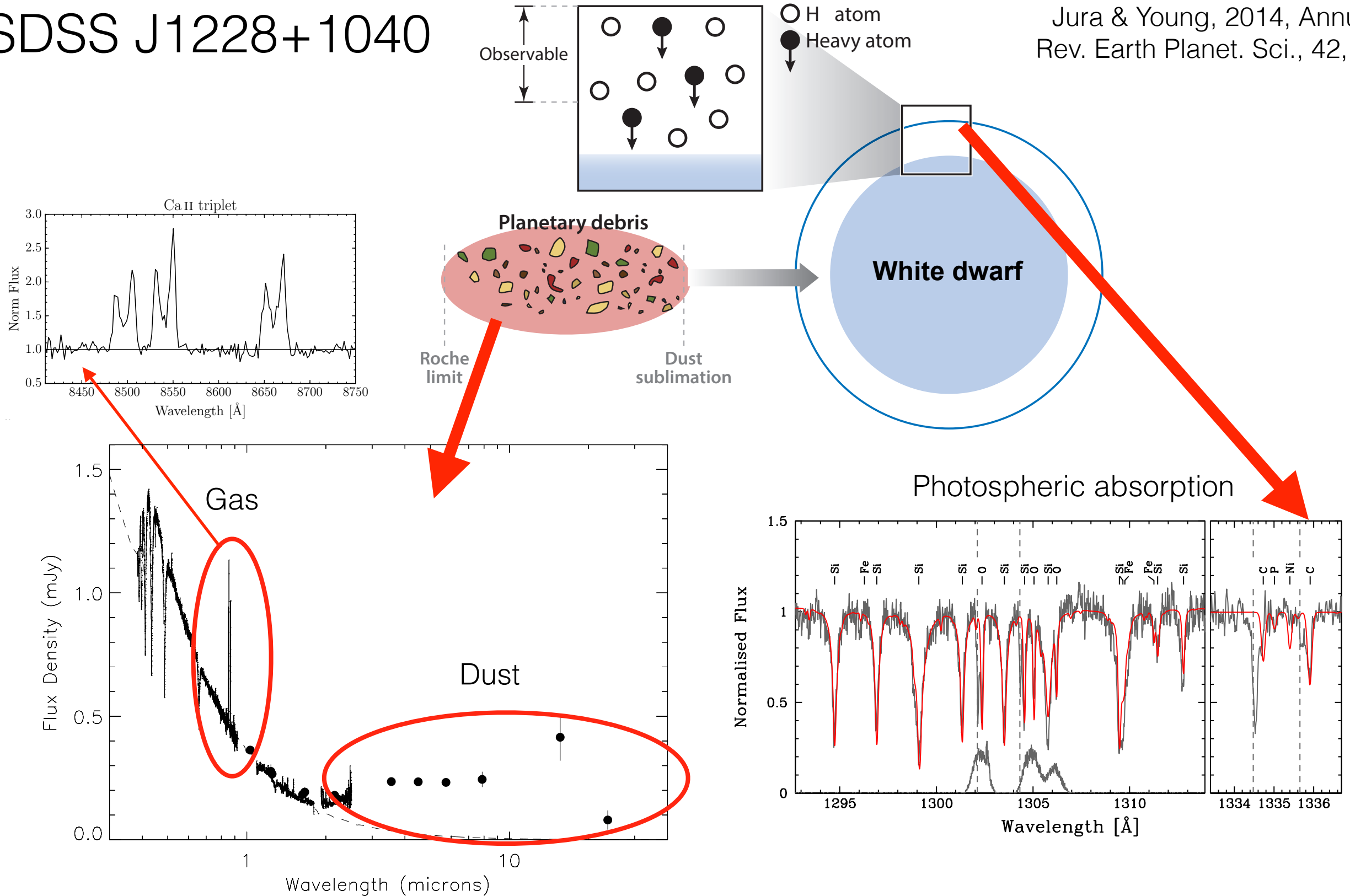
Brinkworth et. al. 2009, ApJ, 696, 1402

Gänsicke et. al. 2012, MNRAS, 424, 333

Remnant Planetary Systems

SDSS J1228+1040

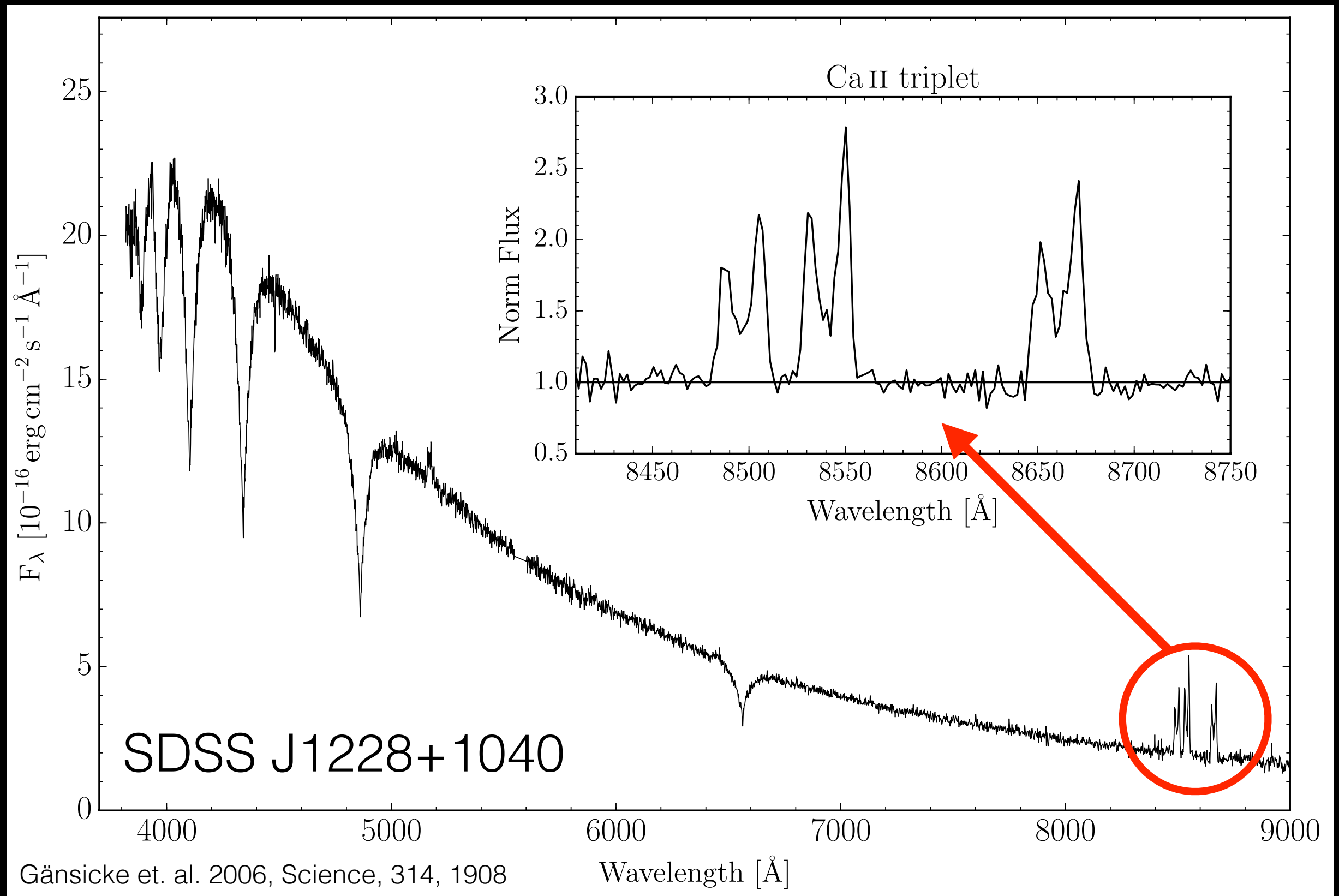
Jura & Young, 2014, Annu.
Rev. Earth Planet. Sci., 42, 45



Brinkworth et. al. 2009, ApJ, 696, 1402

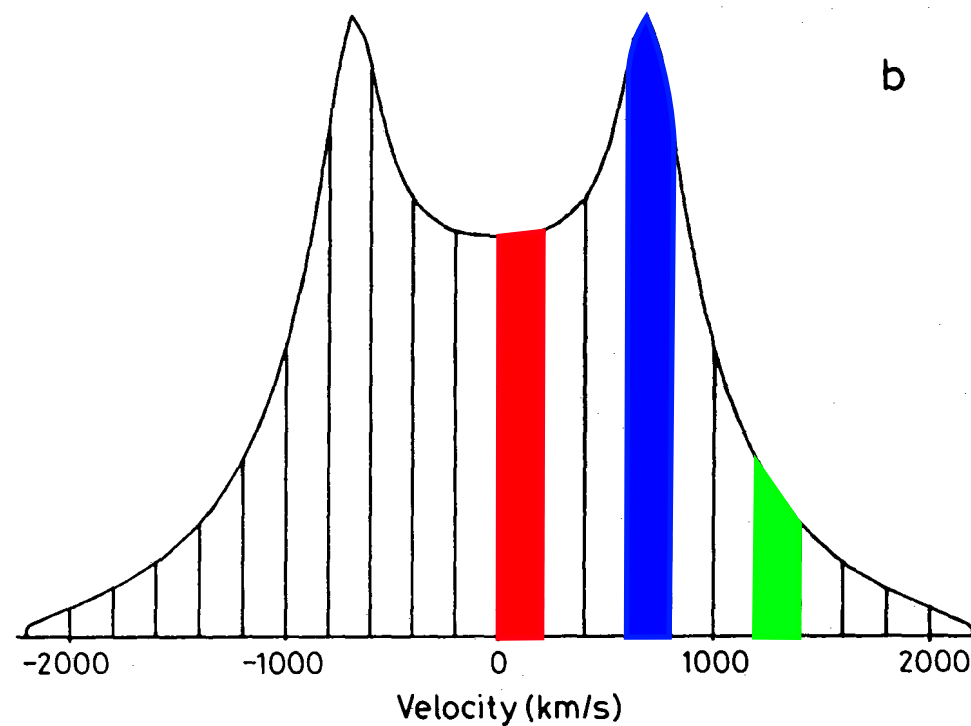
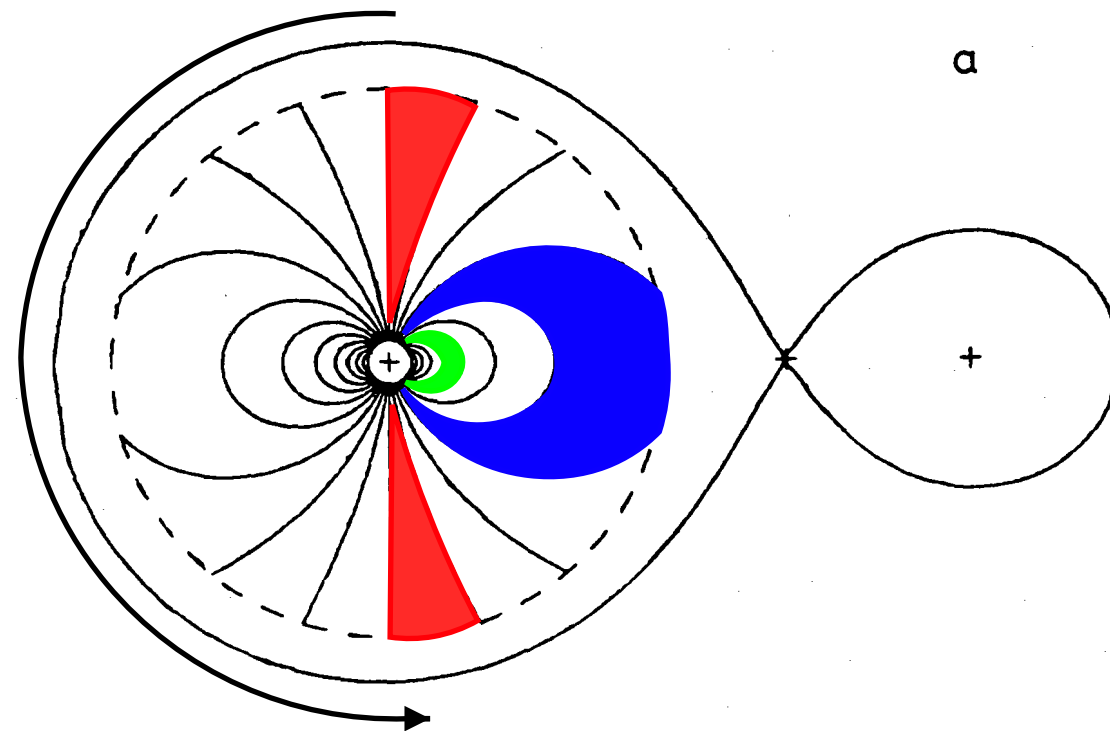
Gänsicke et. al. 2012, MNRAS, 424, 333

The gaseous component of the debris disc



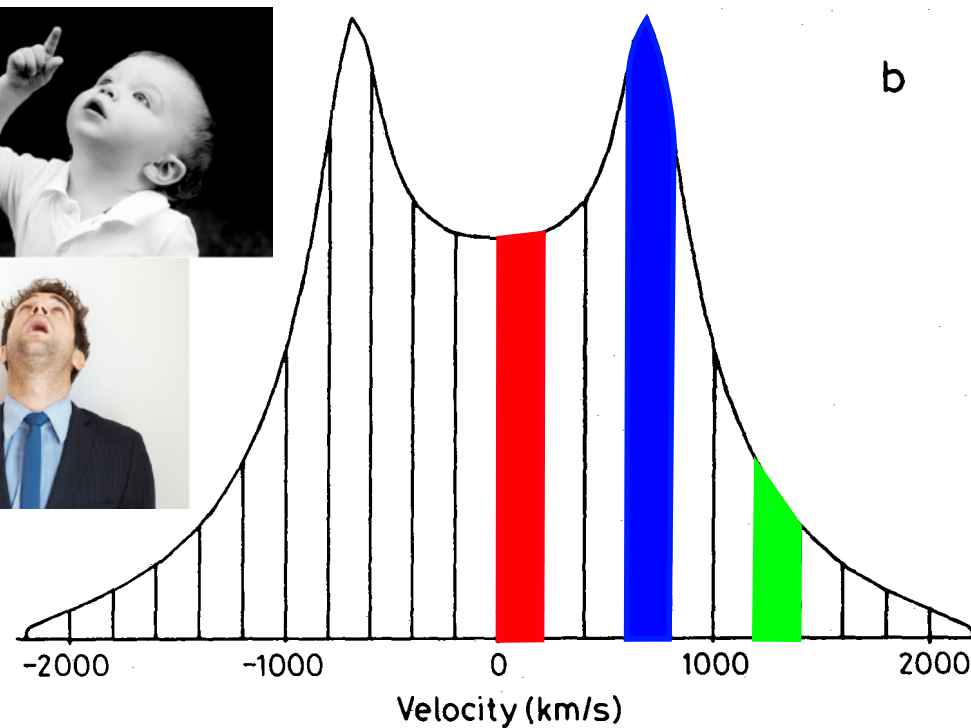
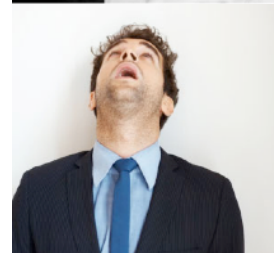
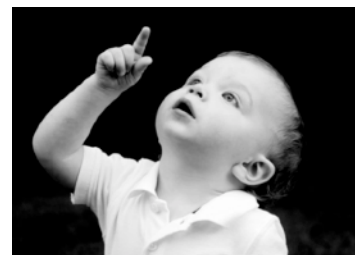
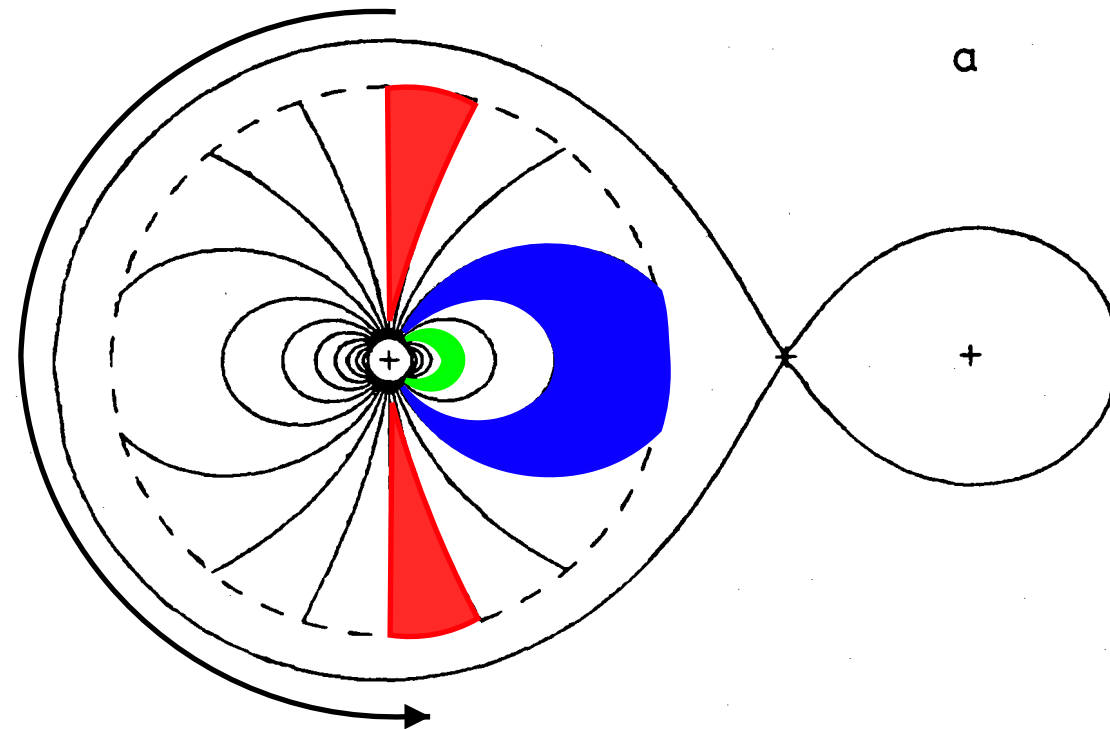
Accretion disc in a binary

K. Horne and T. Marsh

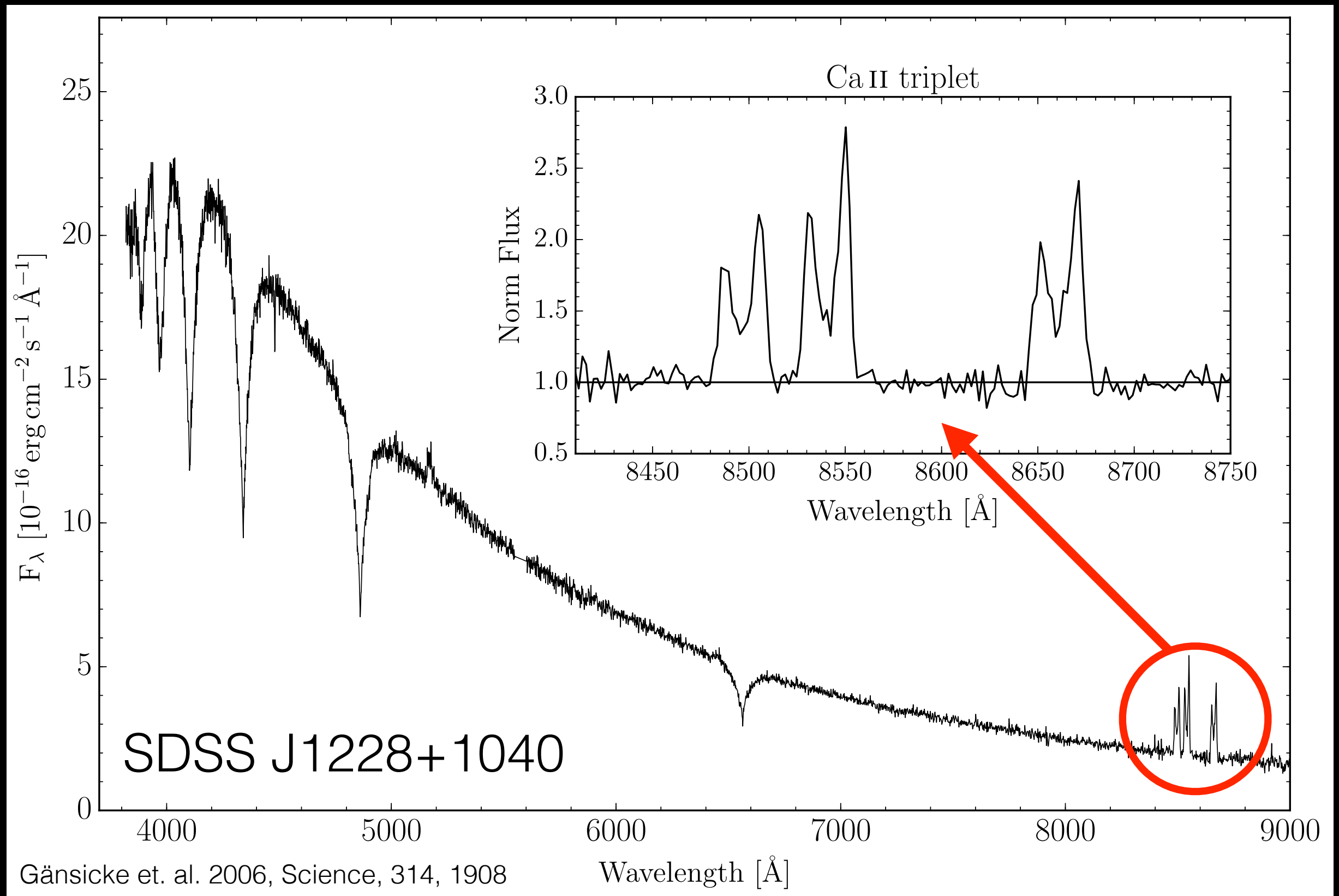


Accretion disc in a binary

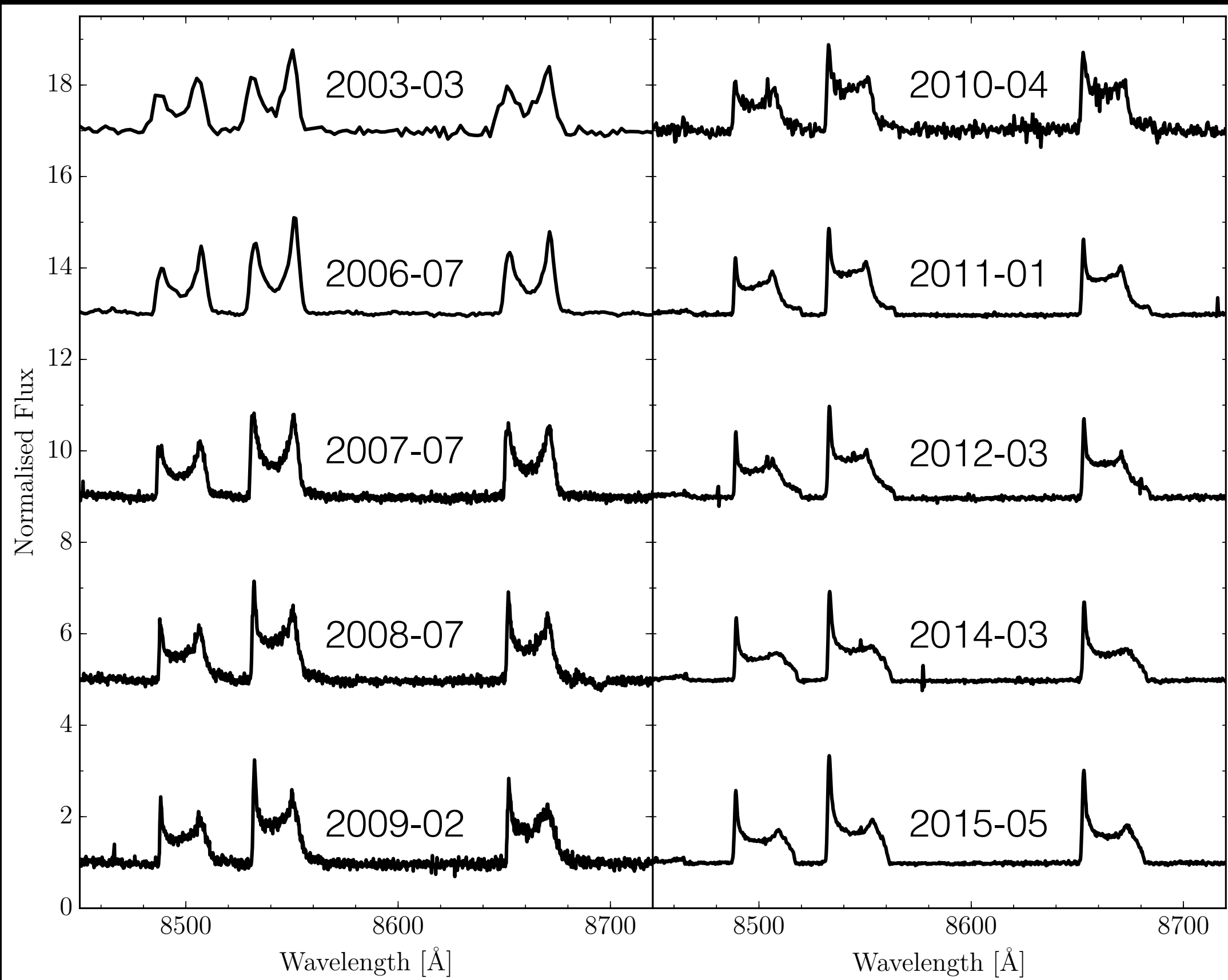
K. Horne and T. Marsh



The gaseous component of the debris disc

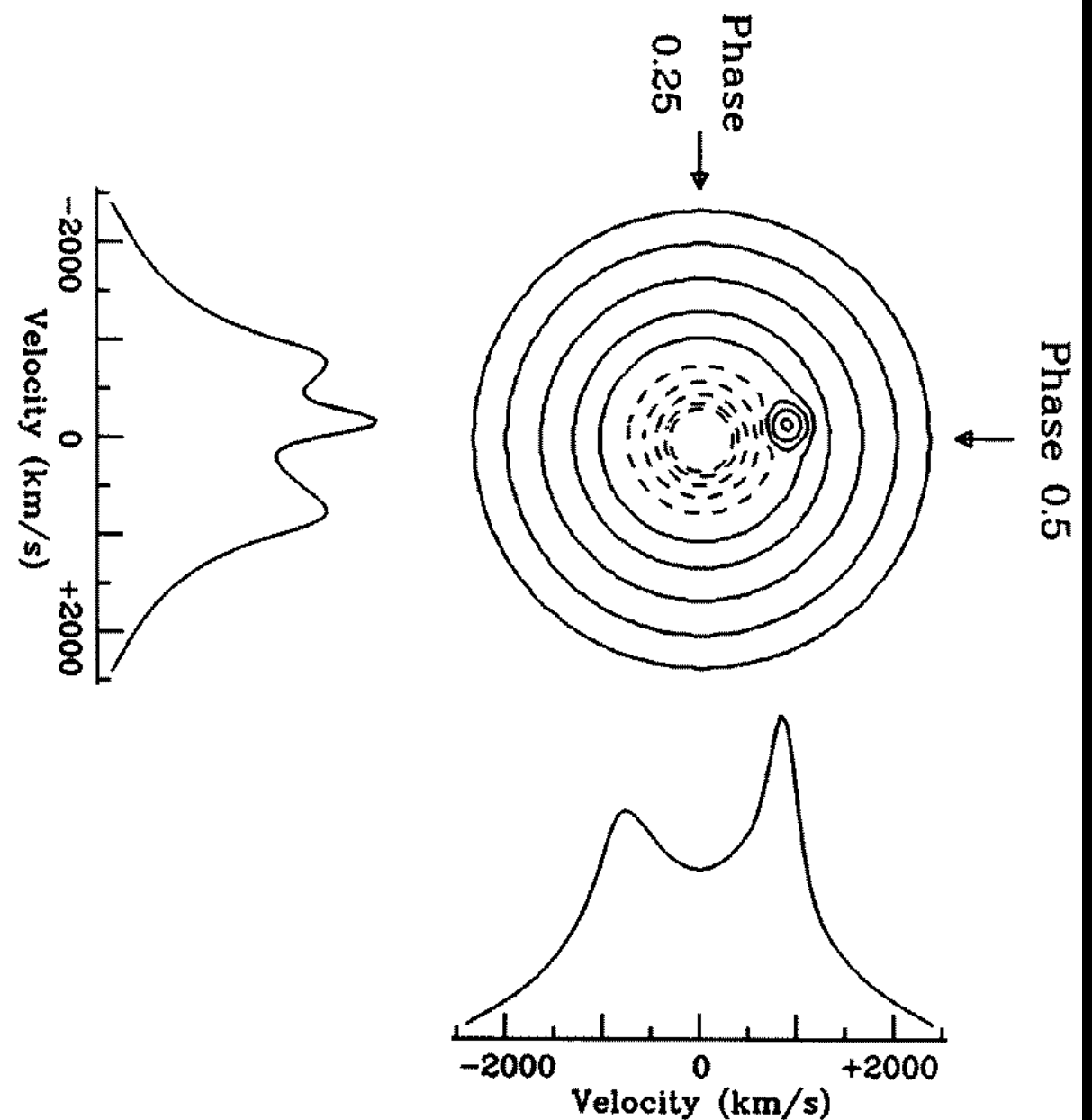
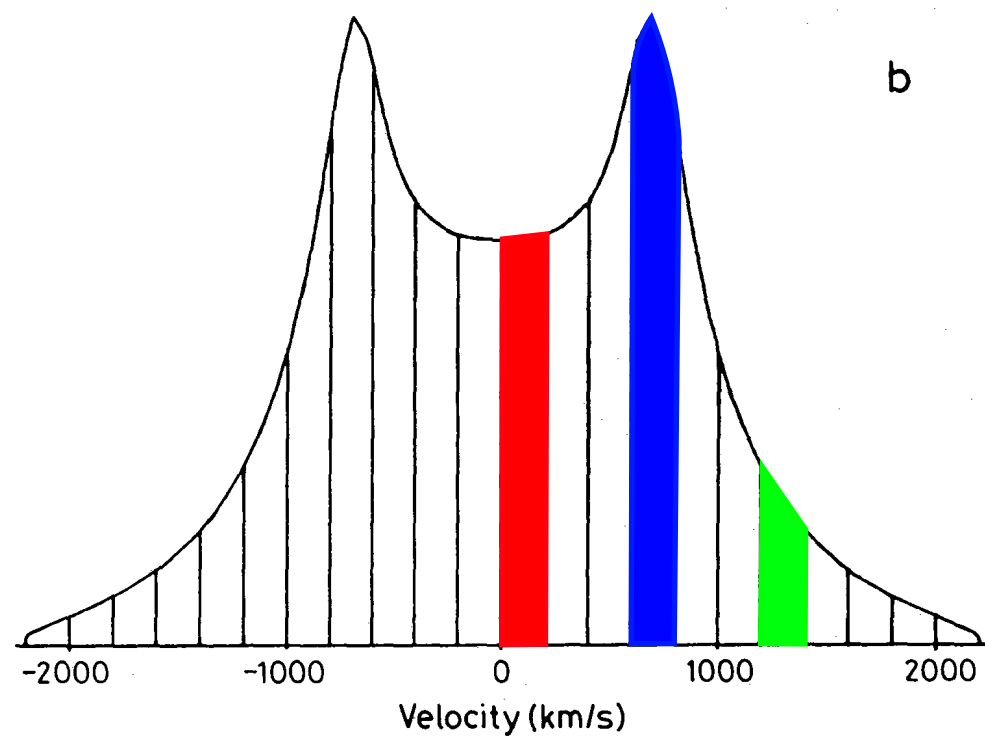
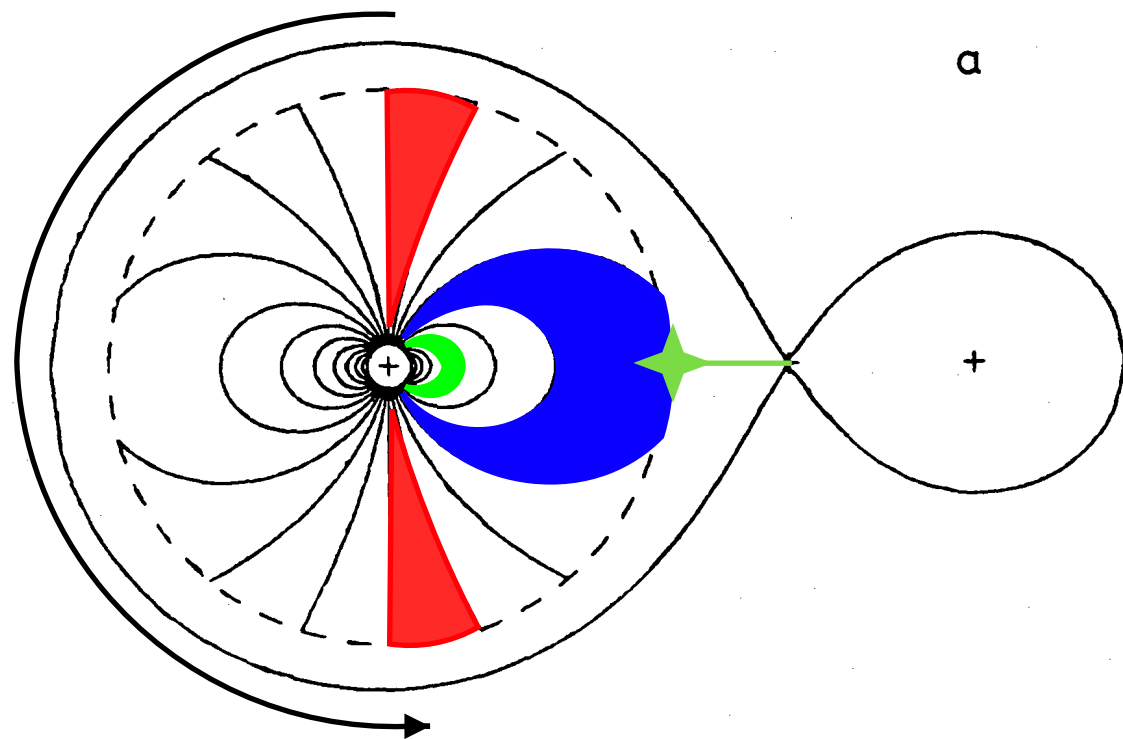


10 out of 18

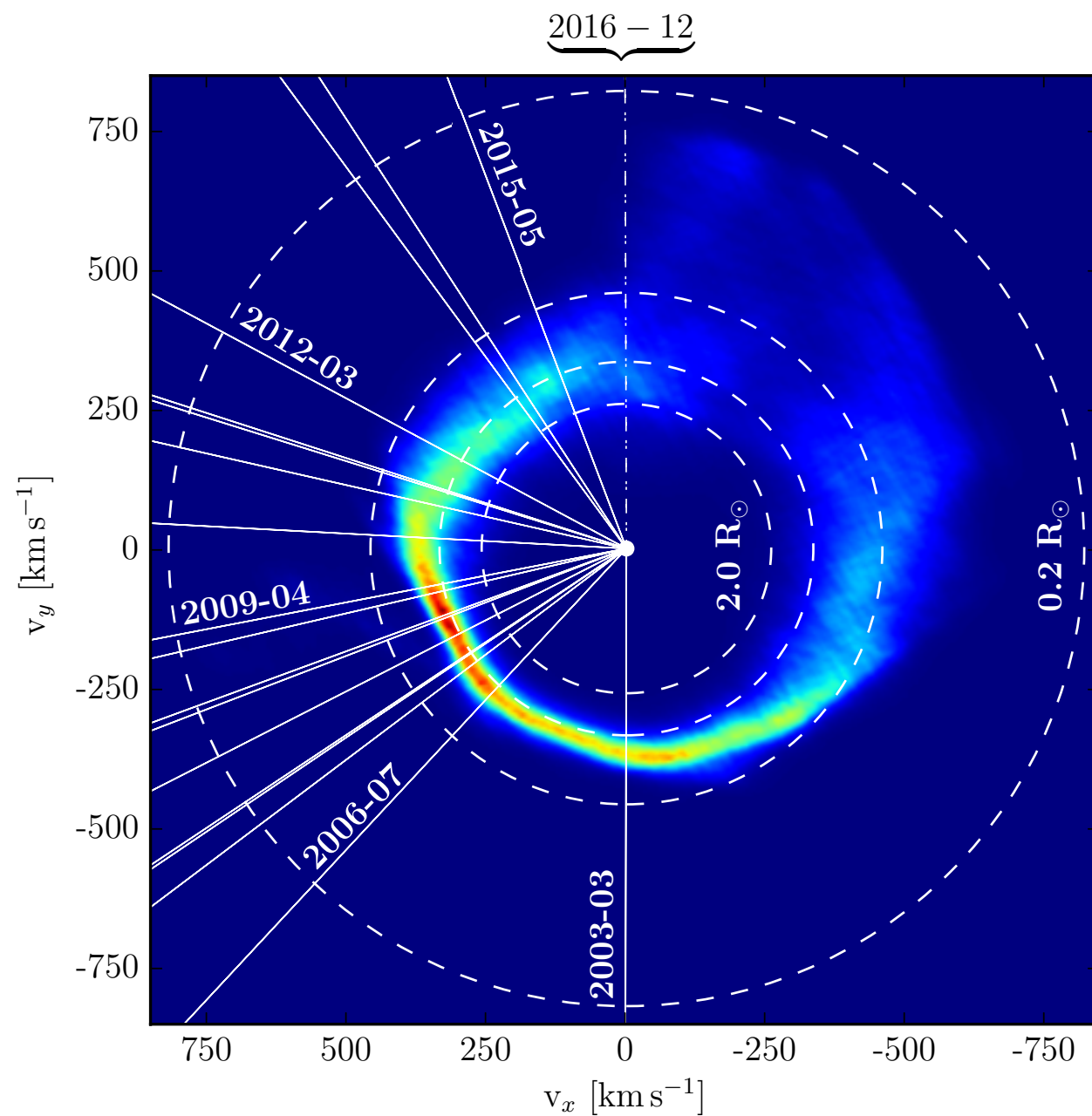


Doppler Tomography

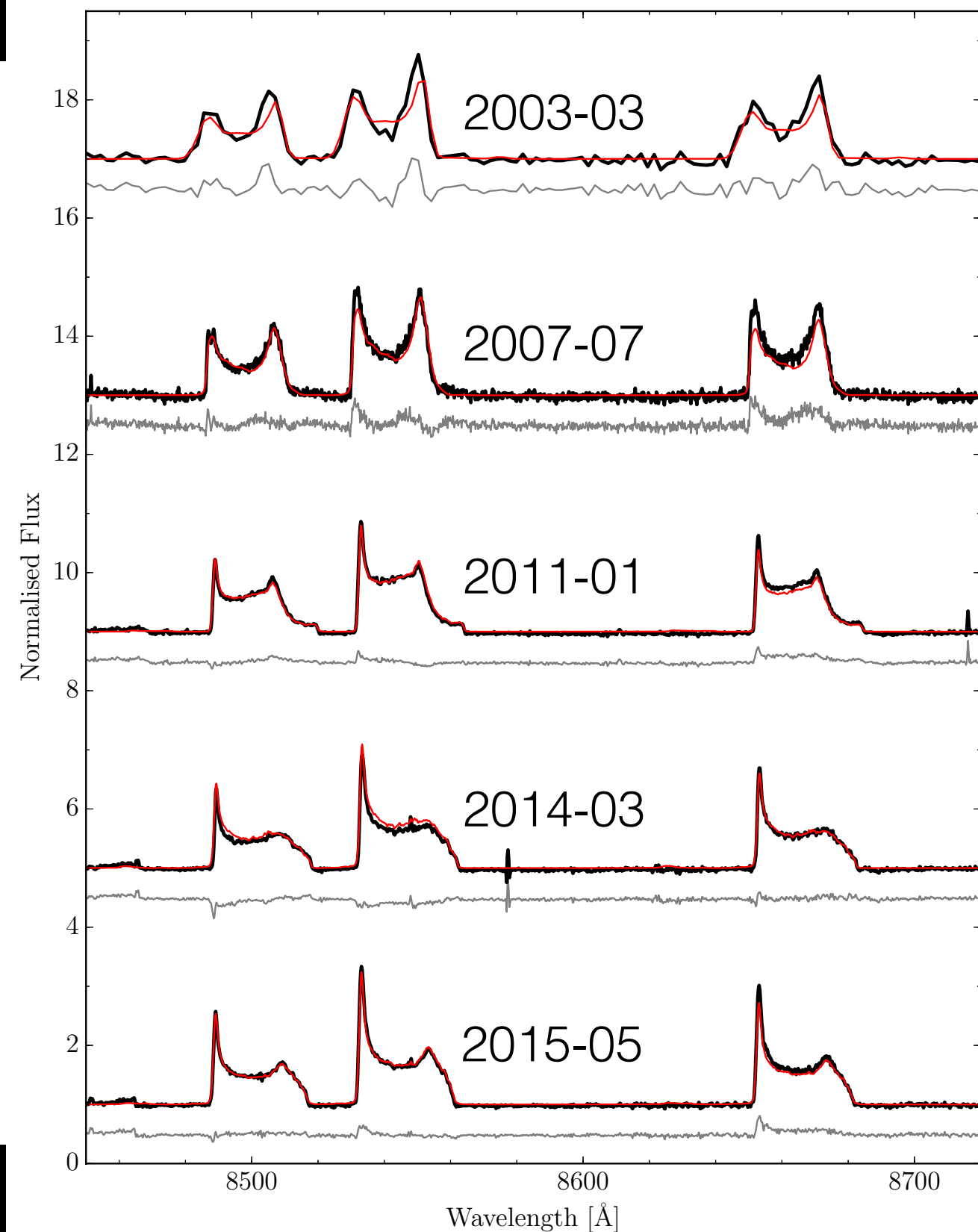
K. Horne and T. Marsh



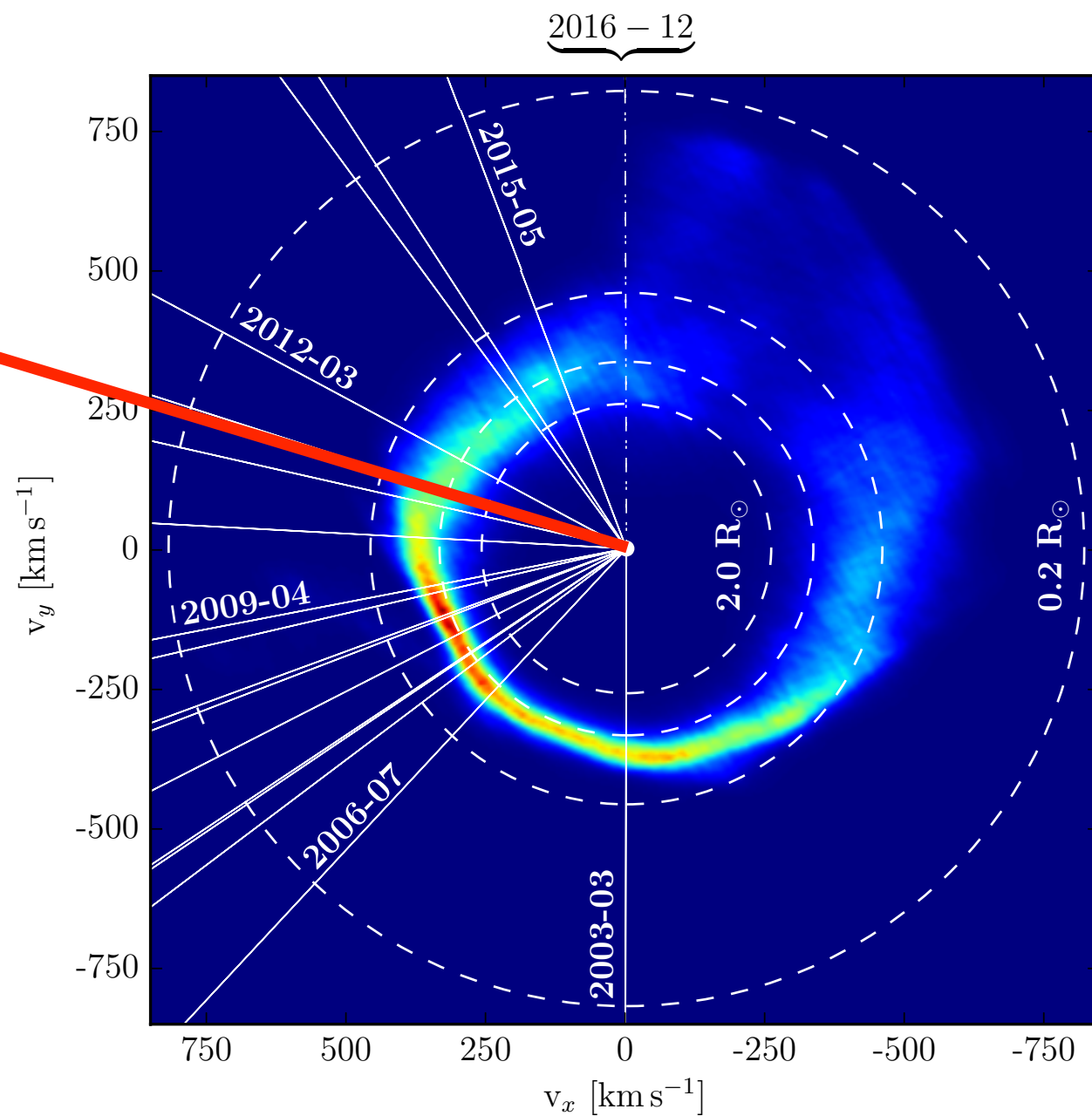
Doppler Tomography



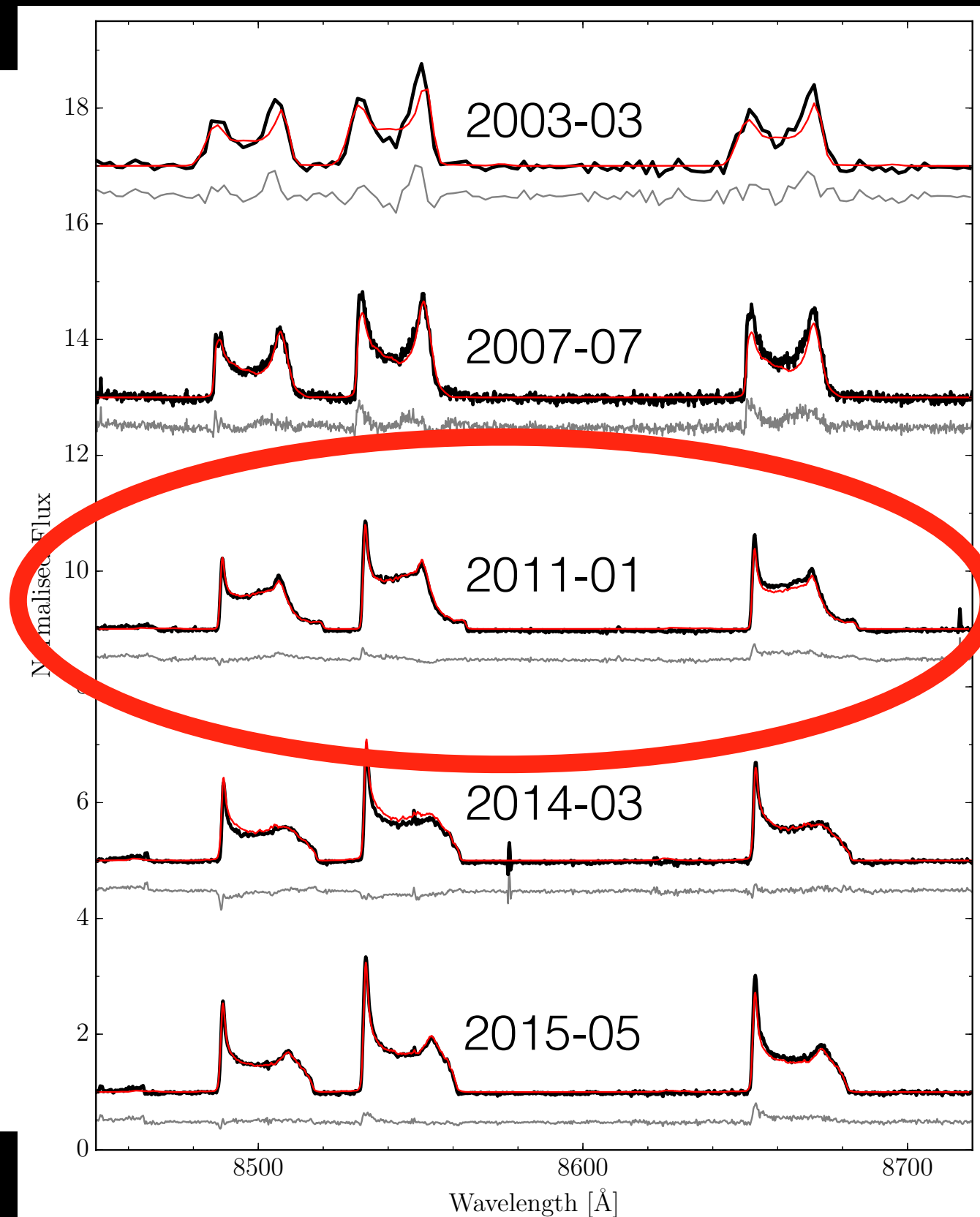
Manser et. al. 2016, MNRAS, 455, 4467



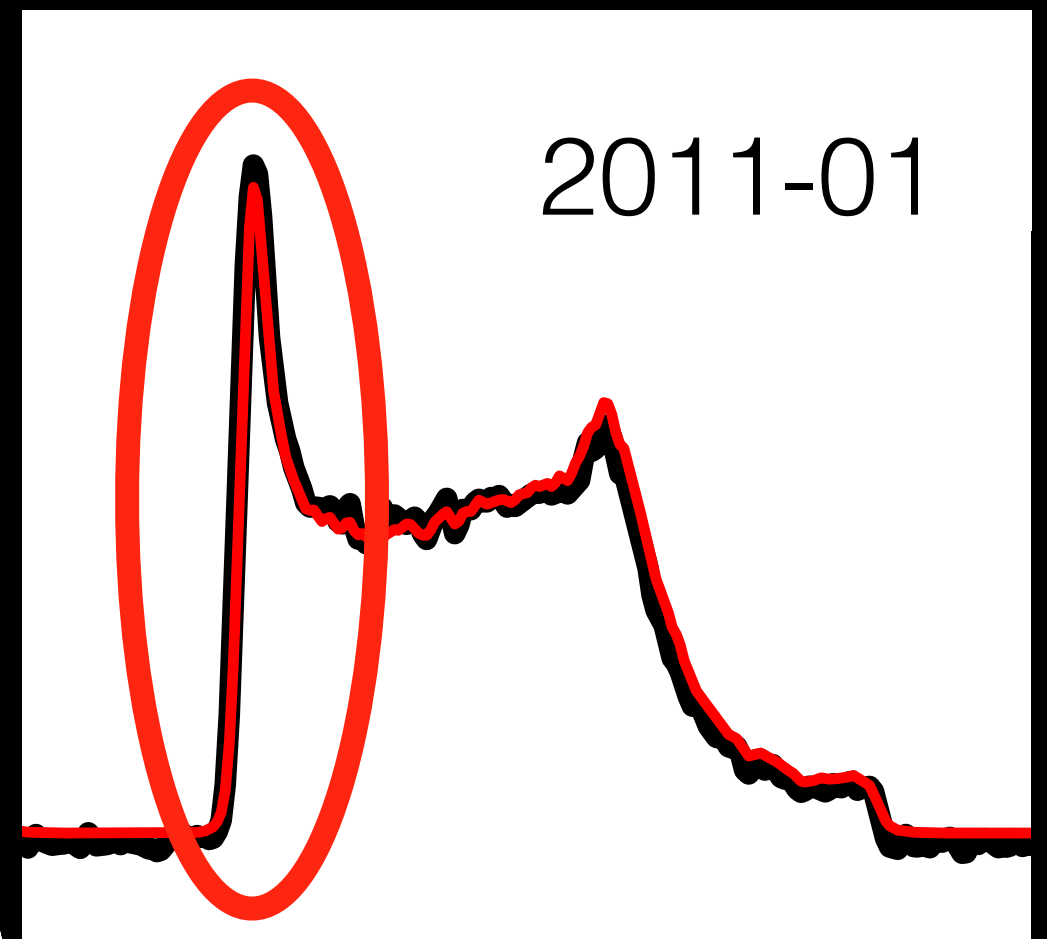
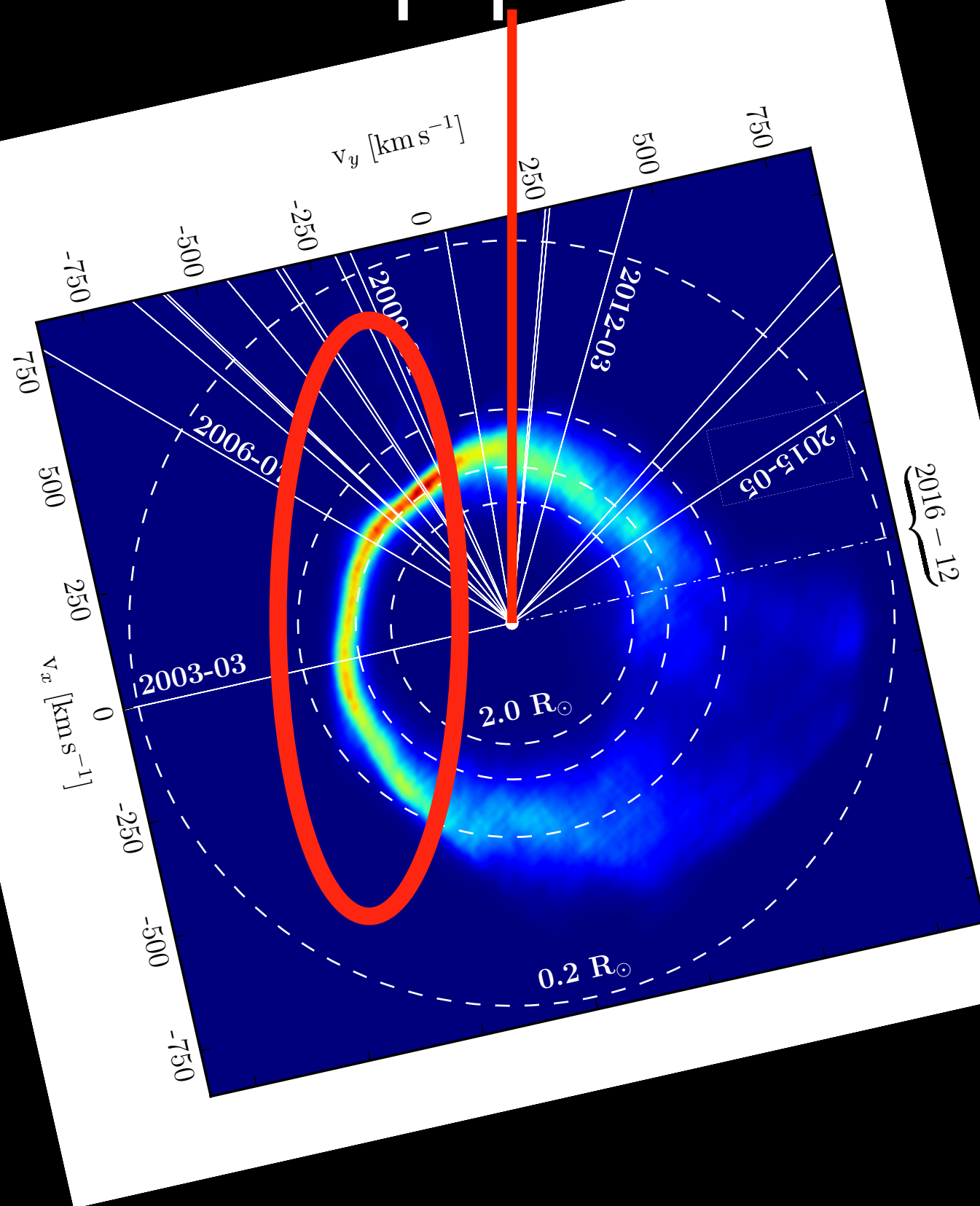
Doppler Tomography



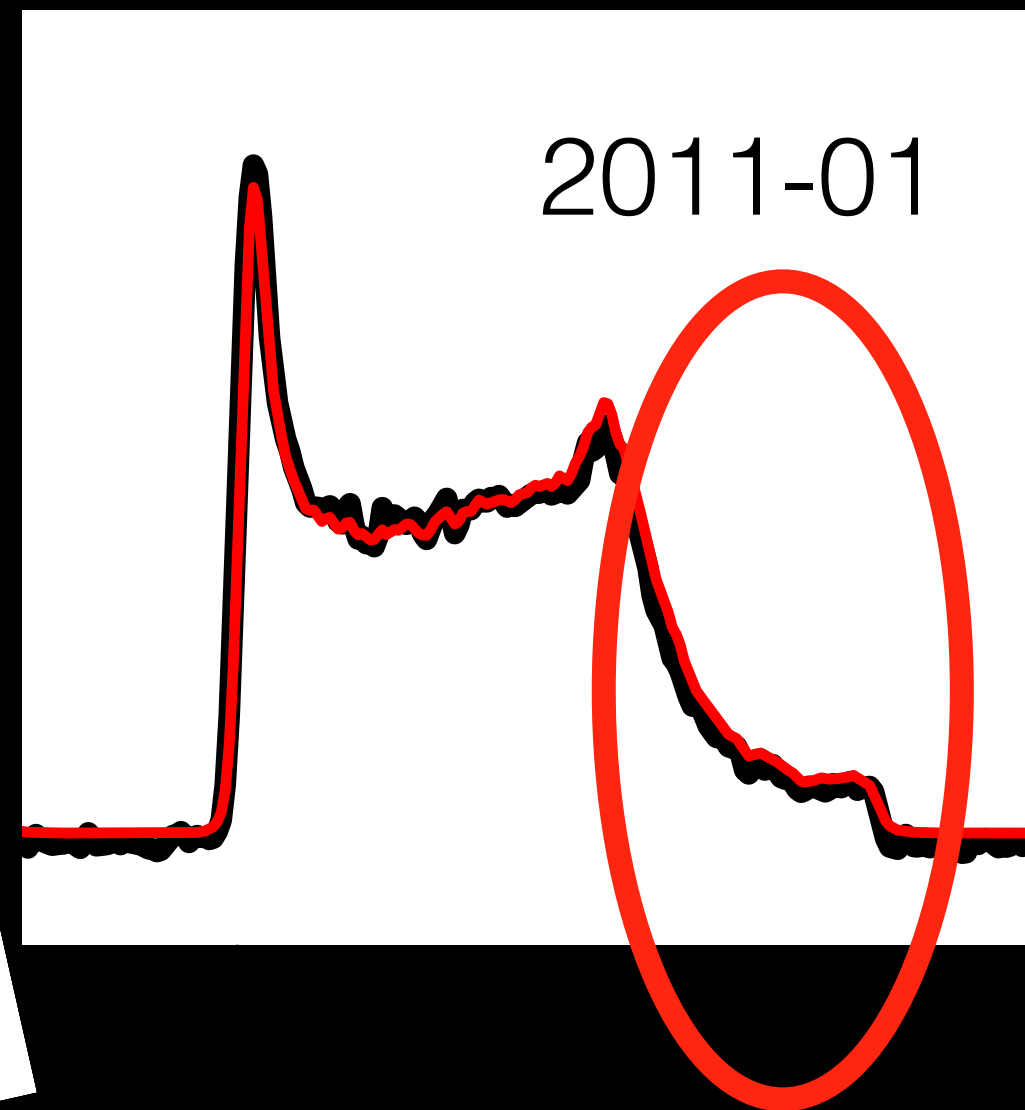
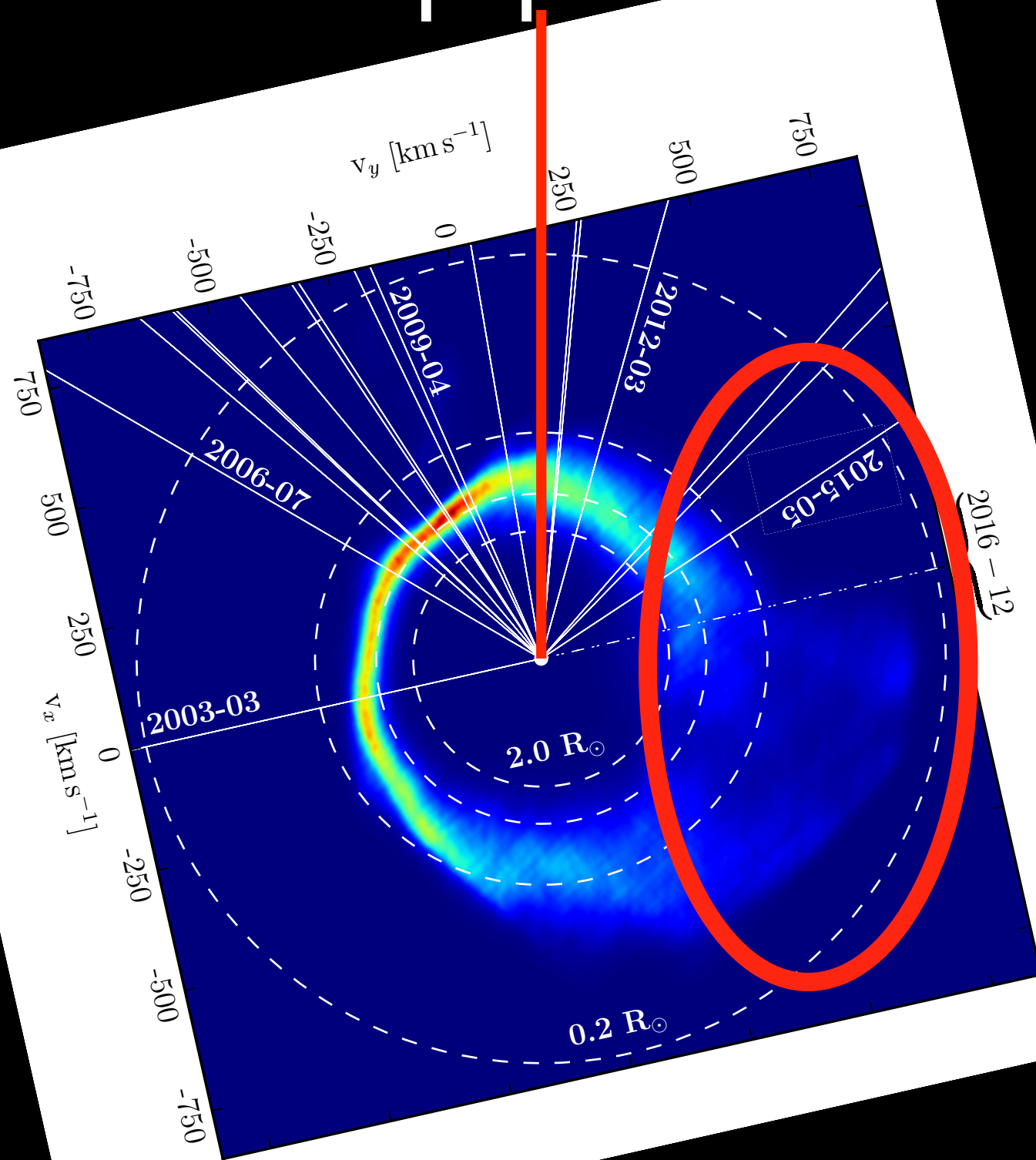
Manser et. al. 2016, MNRAS, 455, 4467



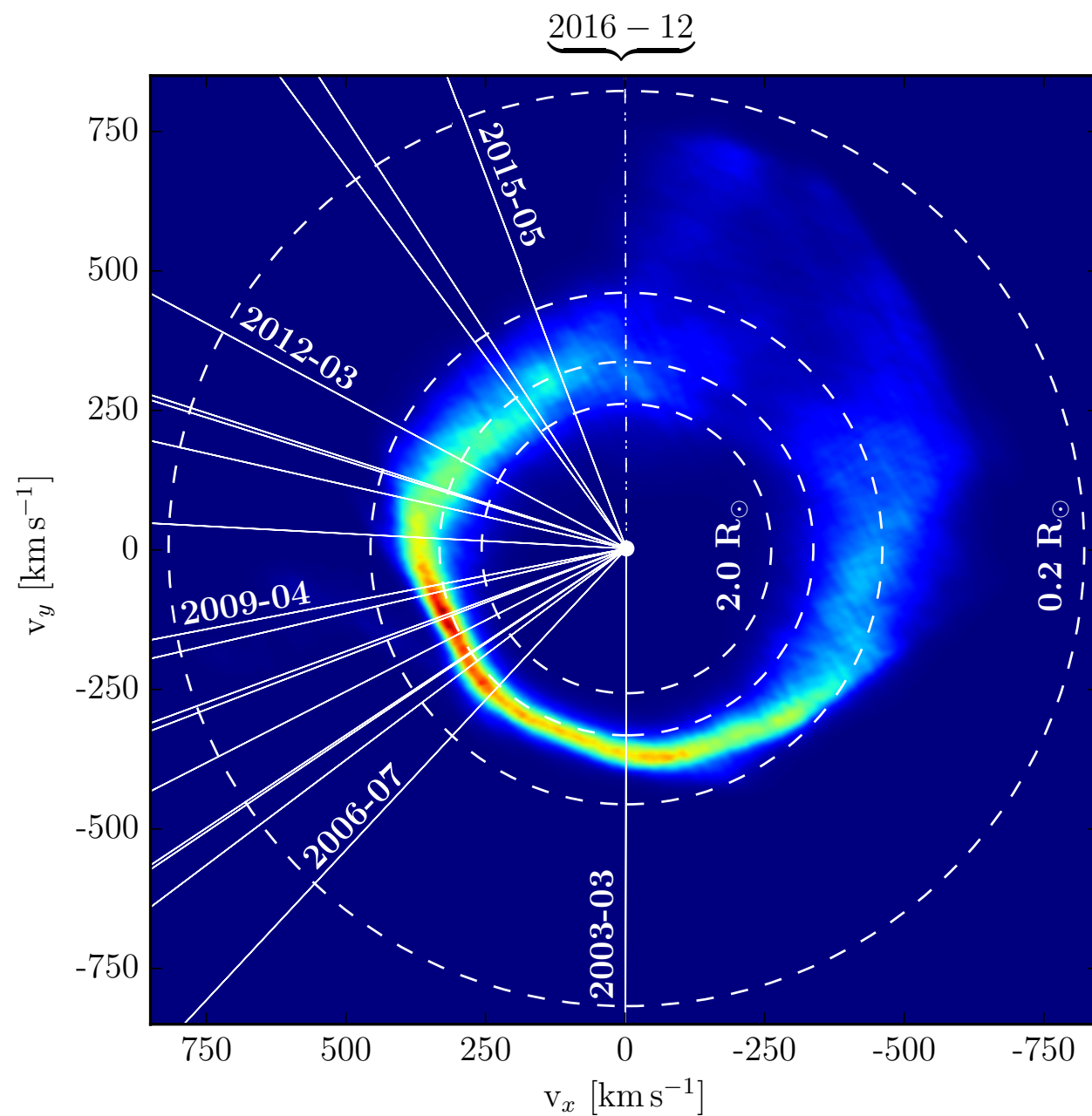
Doppler Tomography



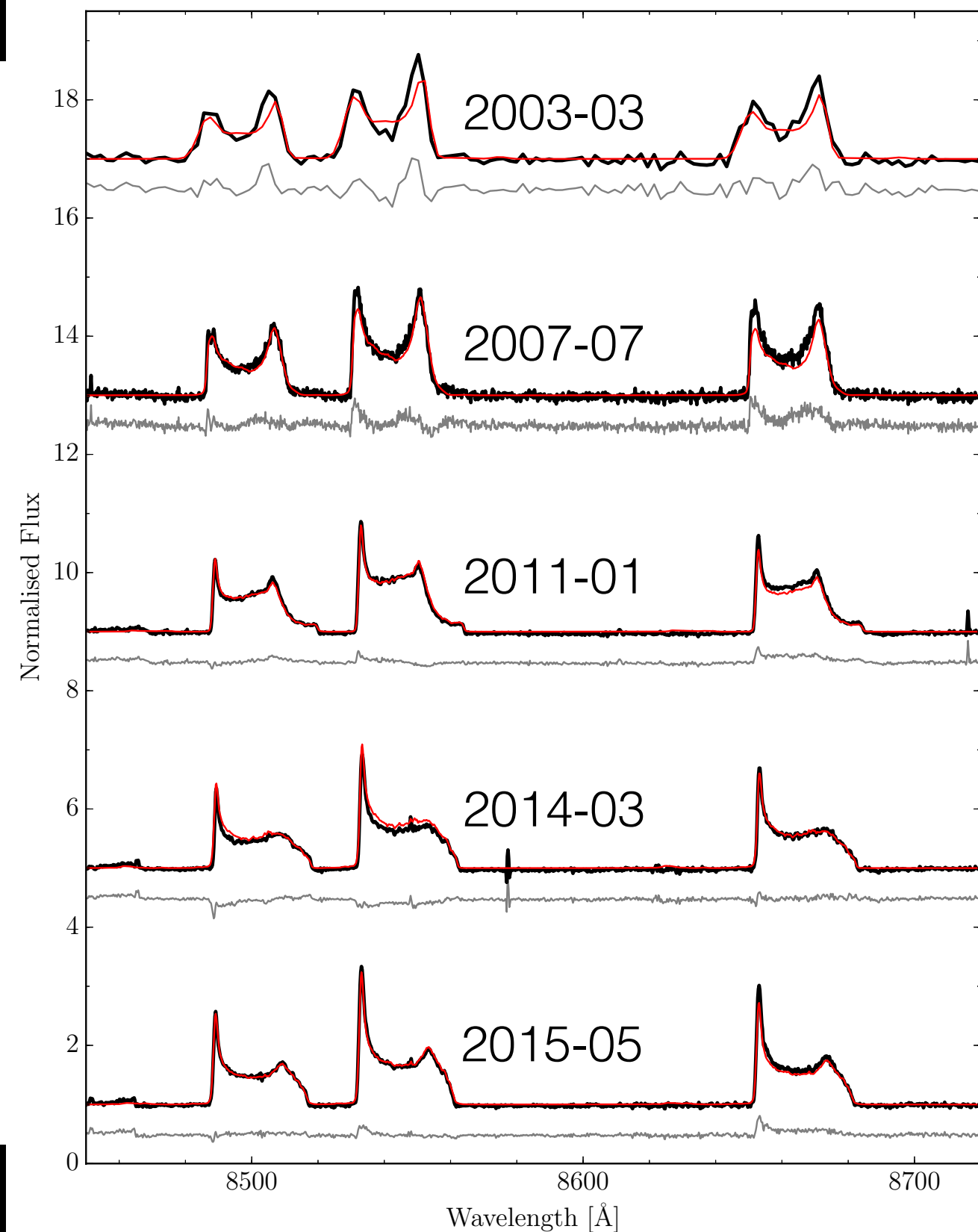
Doppler Tomography



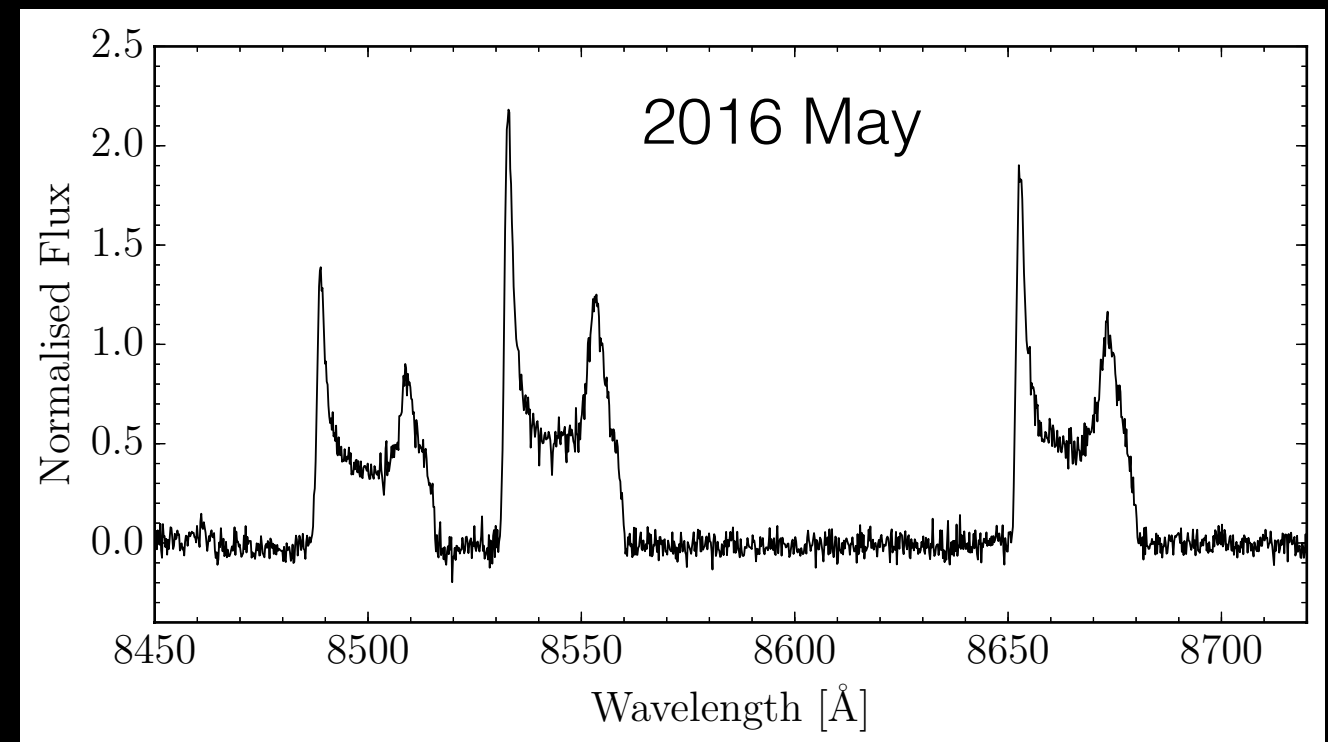
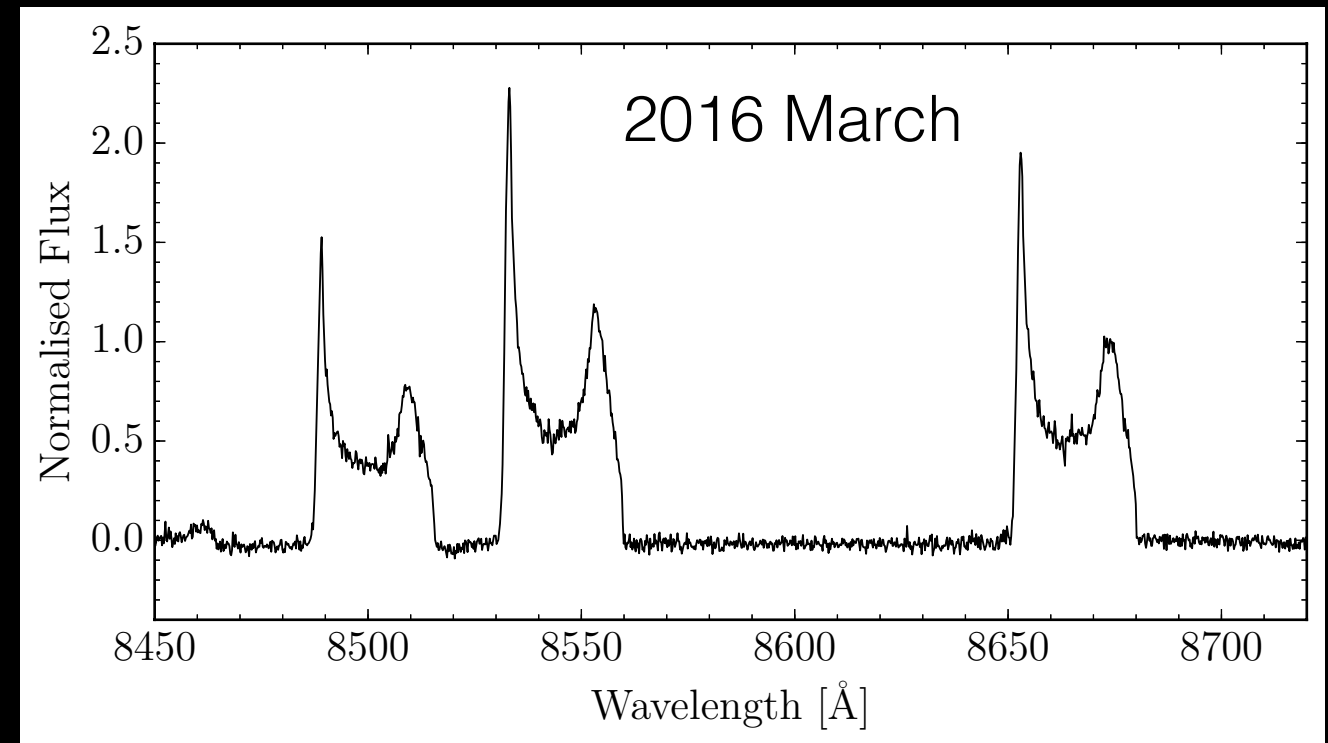
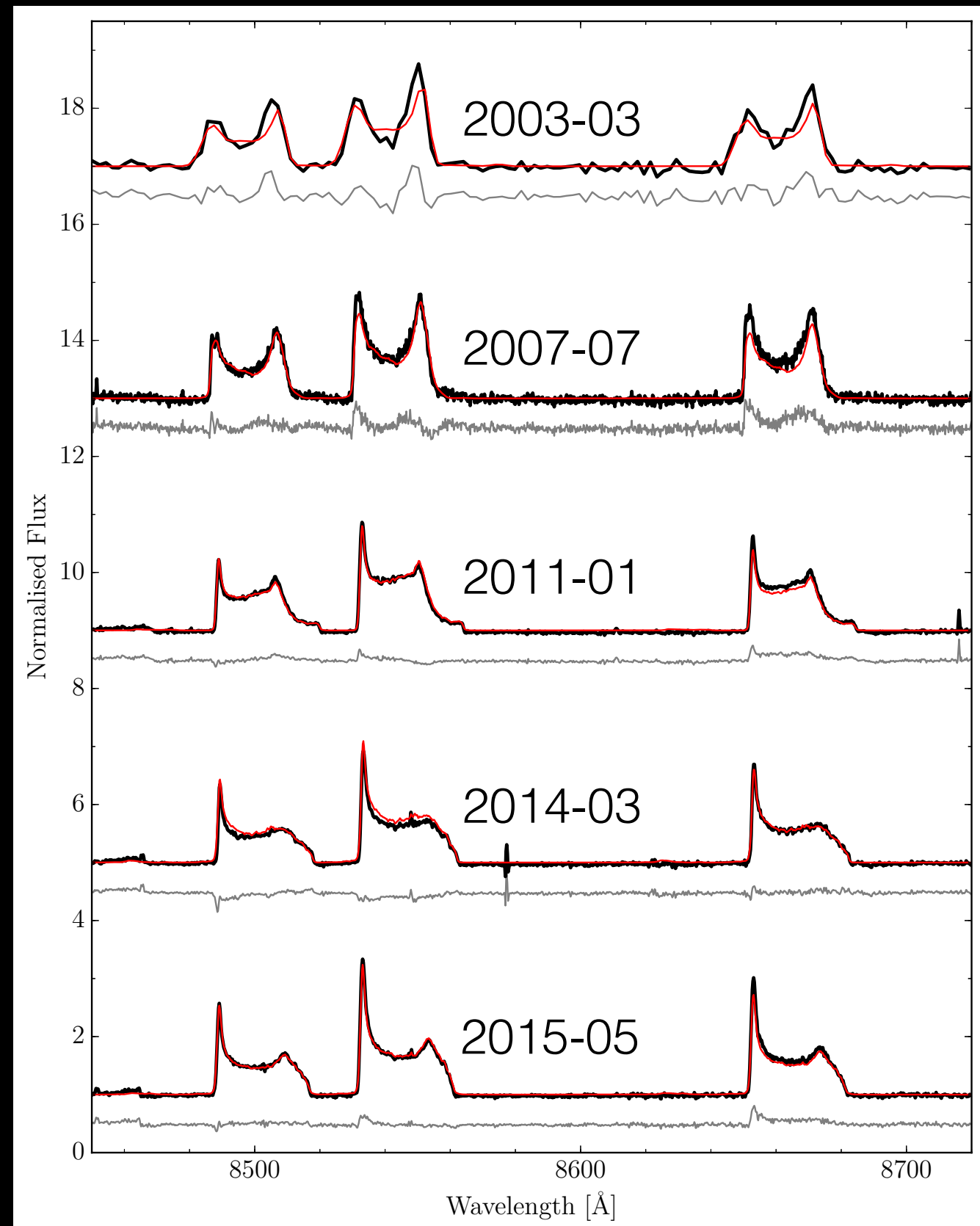
Doppler Tomography



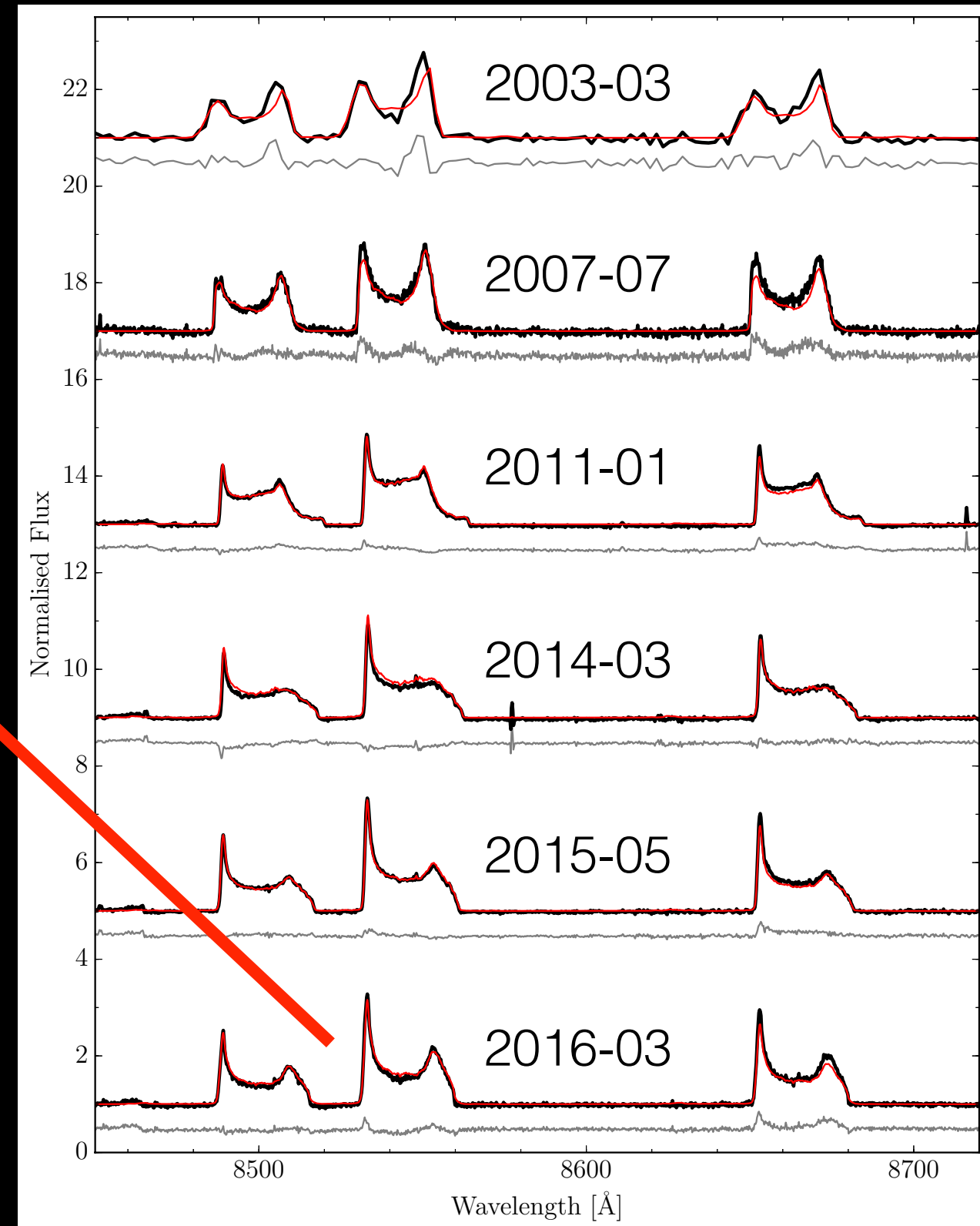
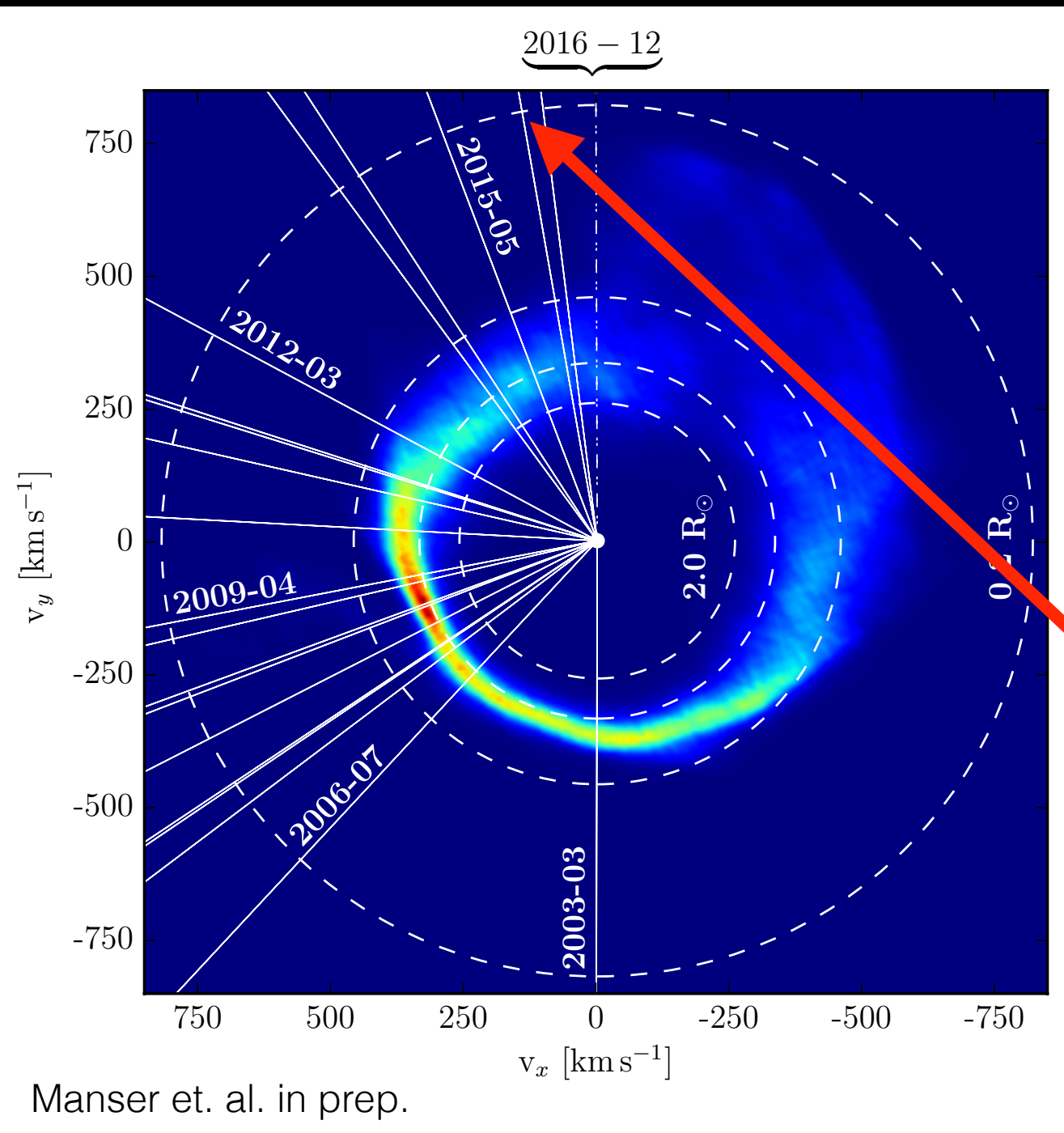
Manser et. al. 2016, MNRAS, 455, 4467



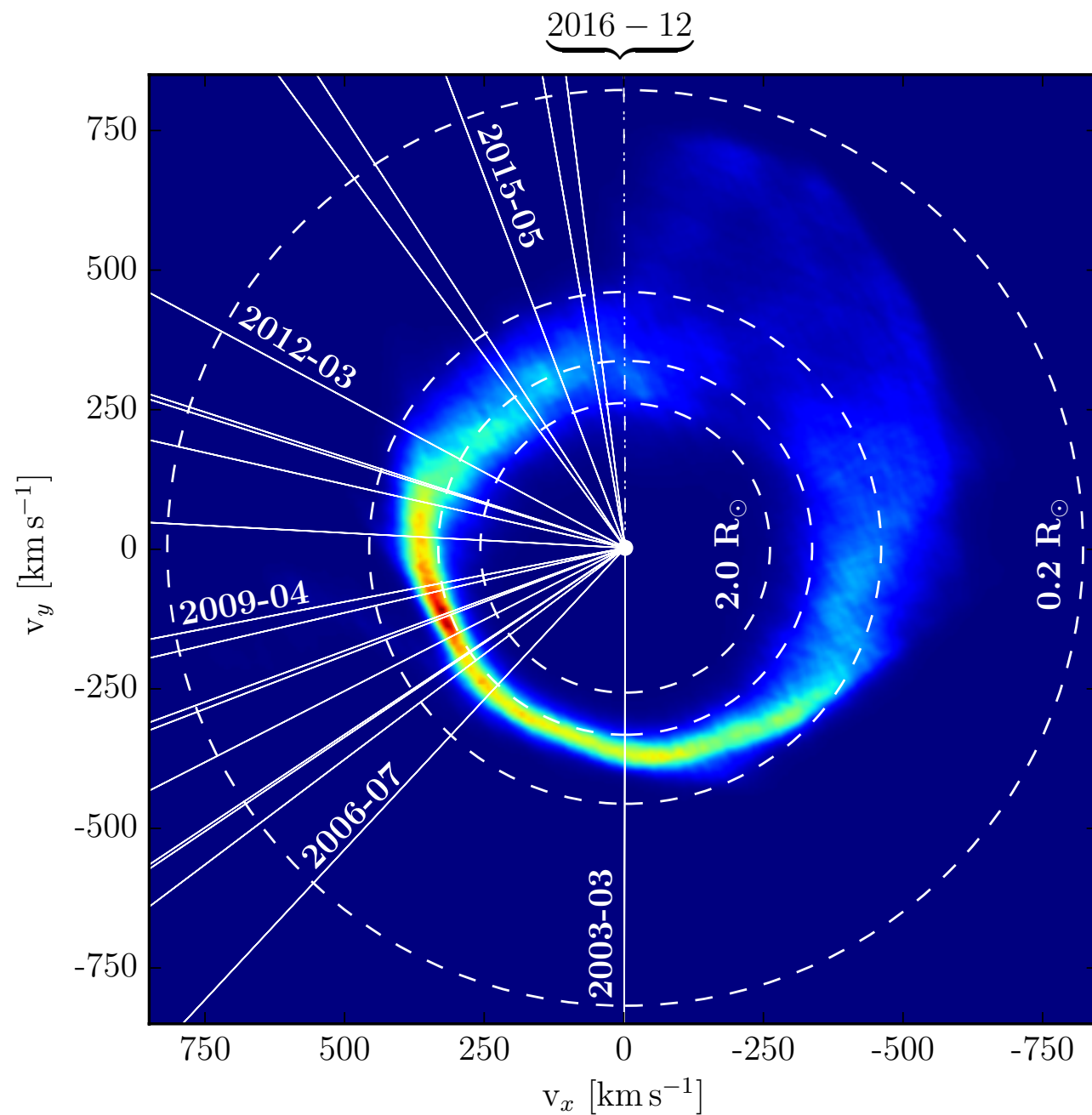
New observations in March and May



A Whole New Map

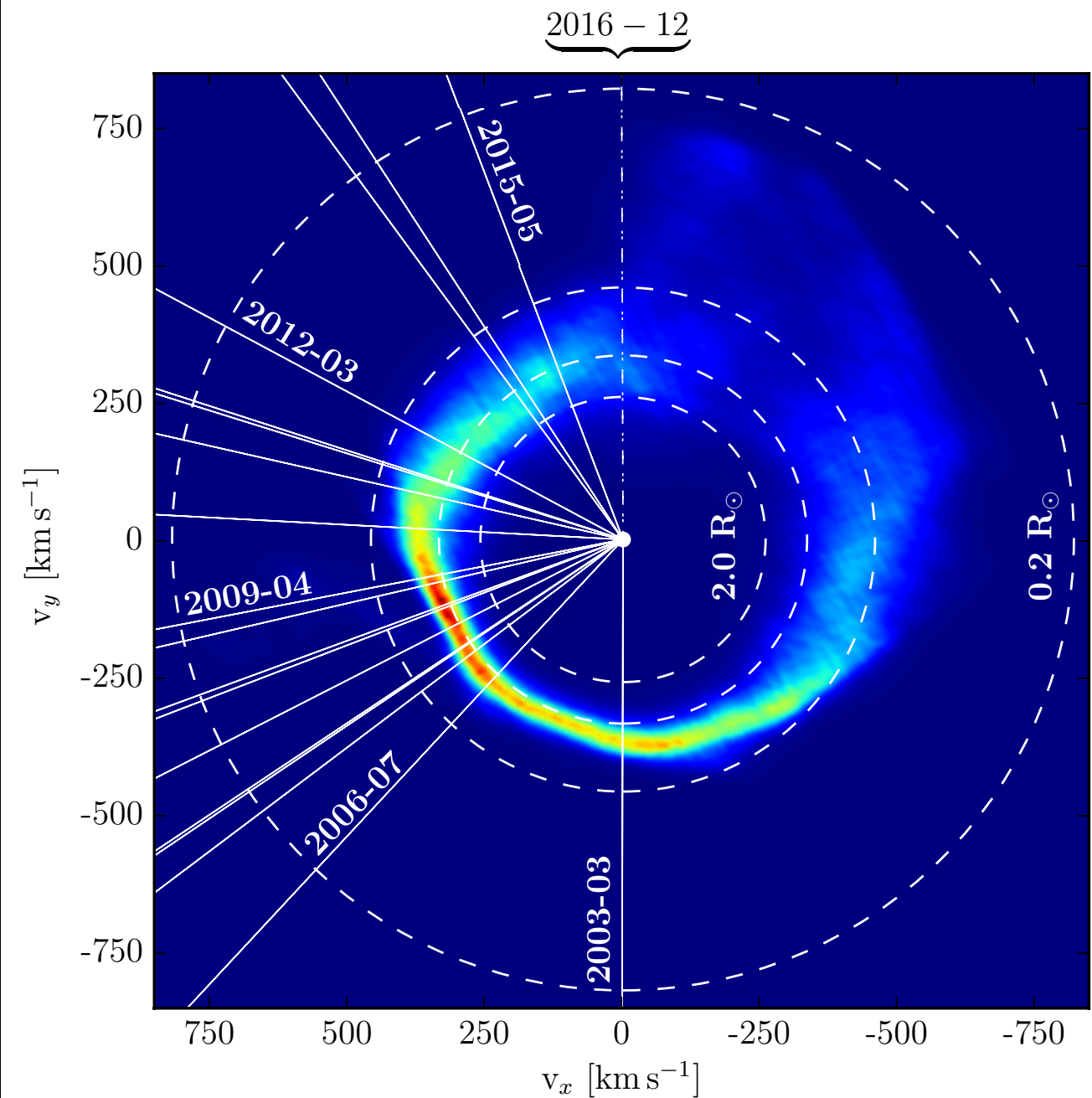


Some more comparing



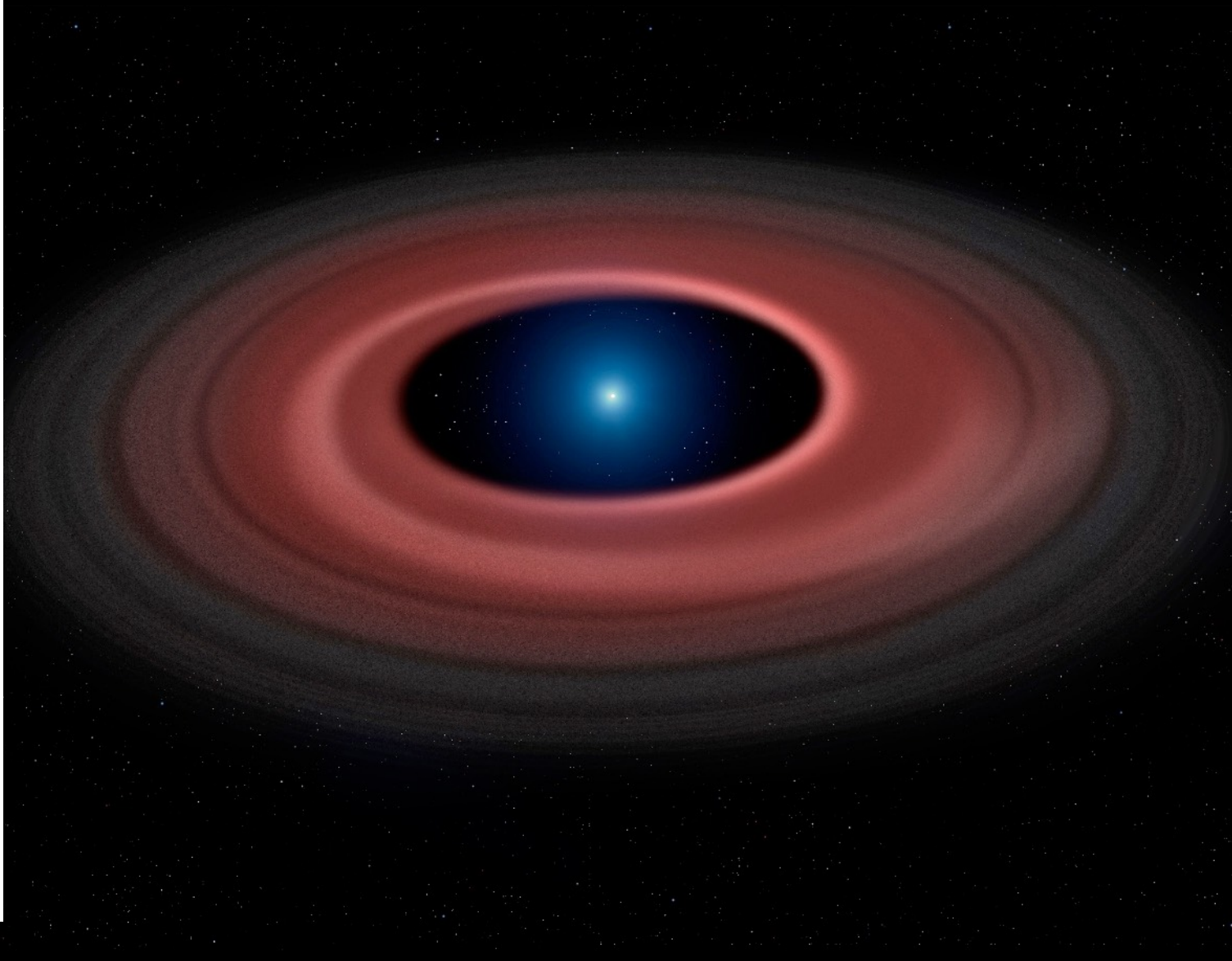
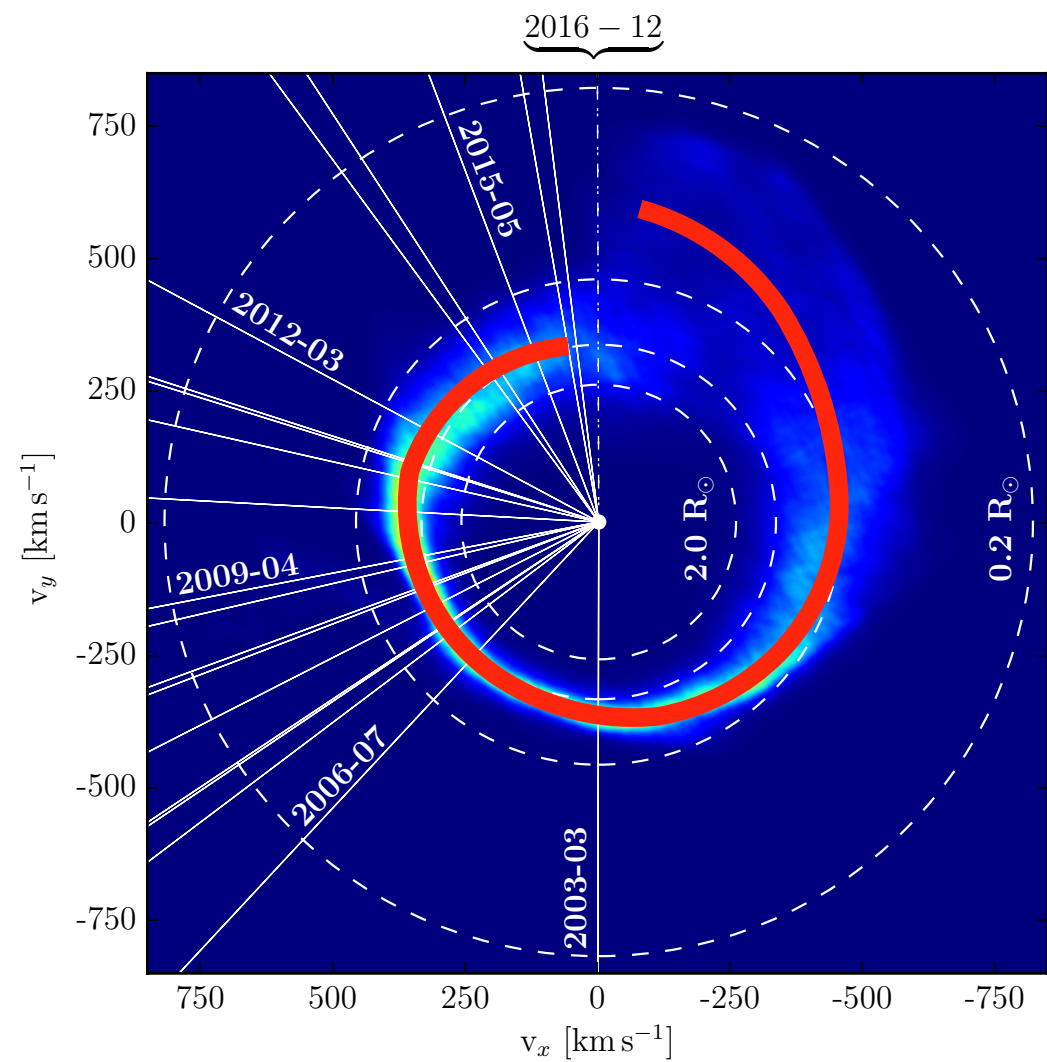
Manser et. al. in prep.

New (20 epochs)

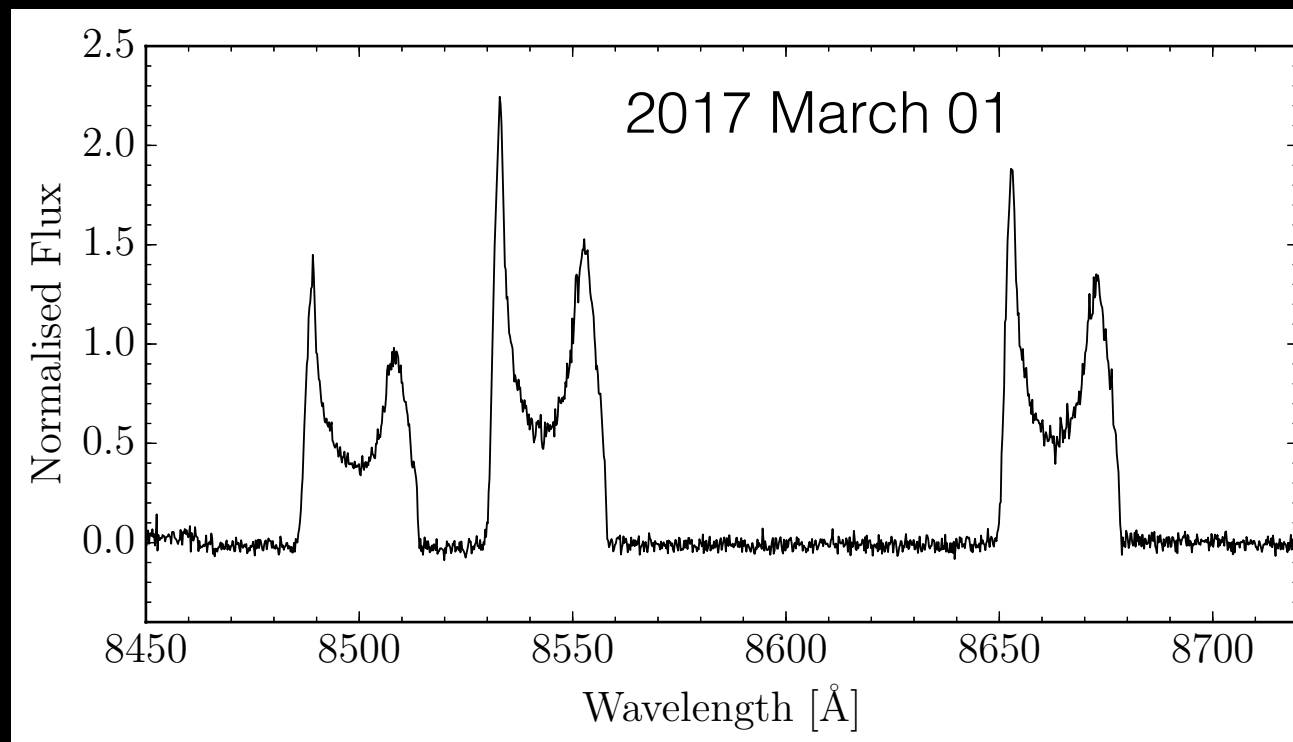
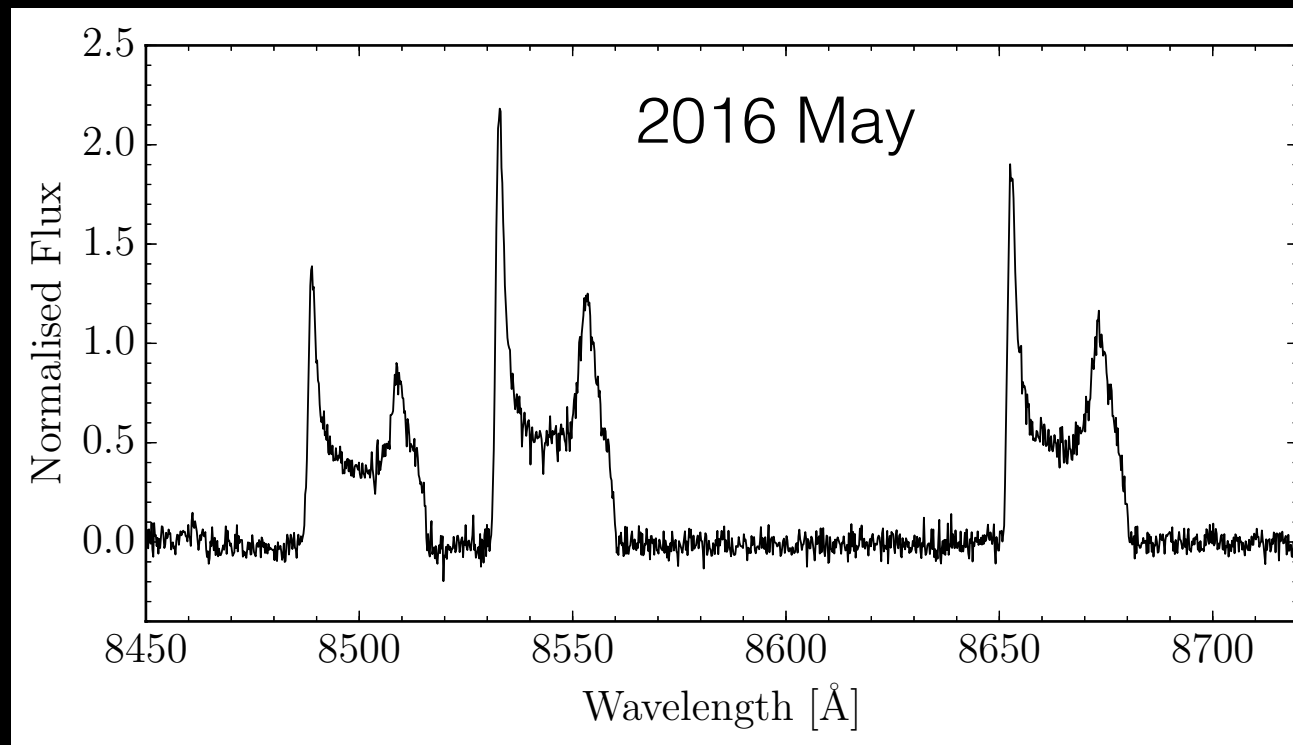


Old (18 epochs)

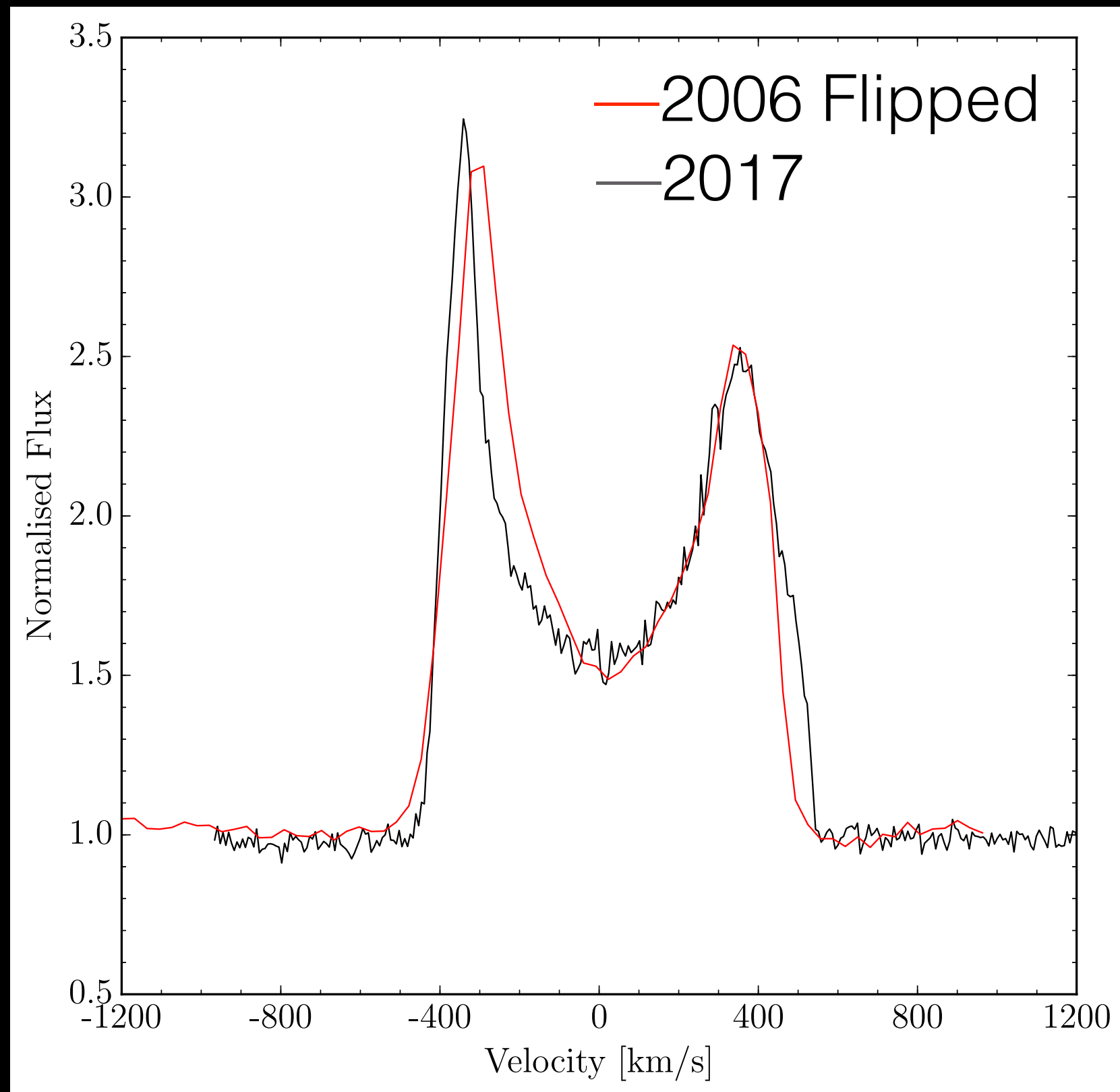
Spiral?



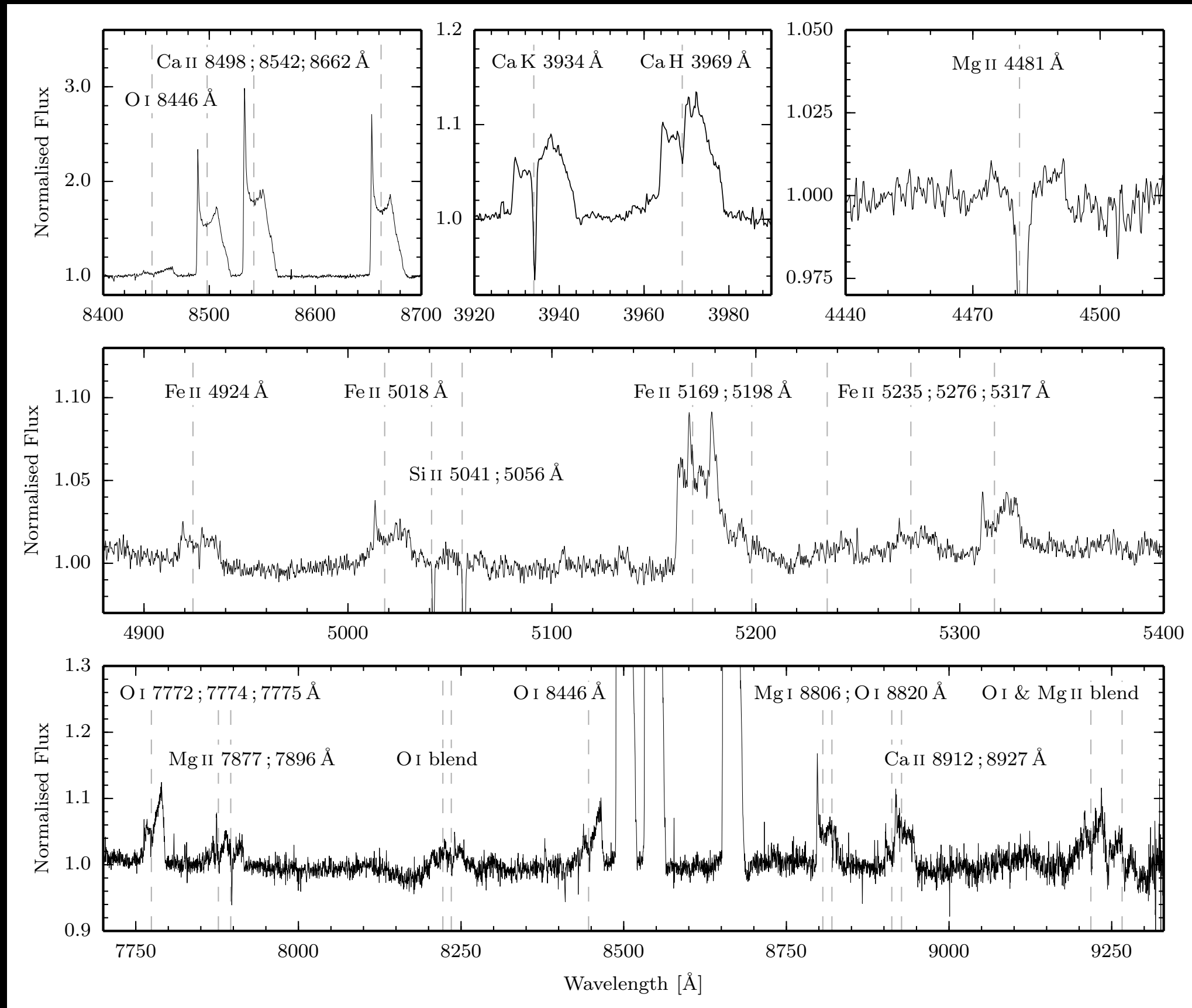
Even newer data!



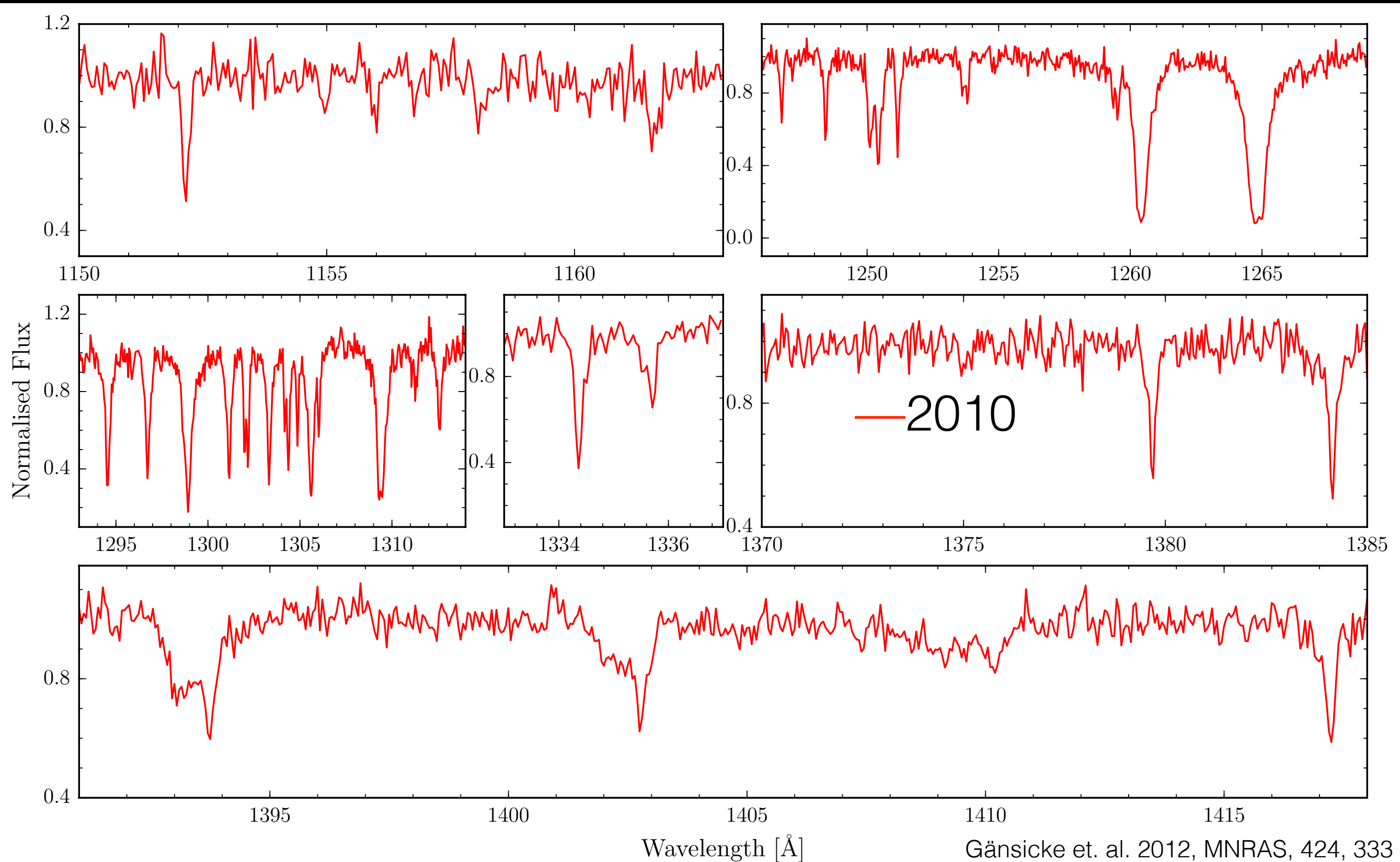
Reached half way?



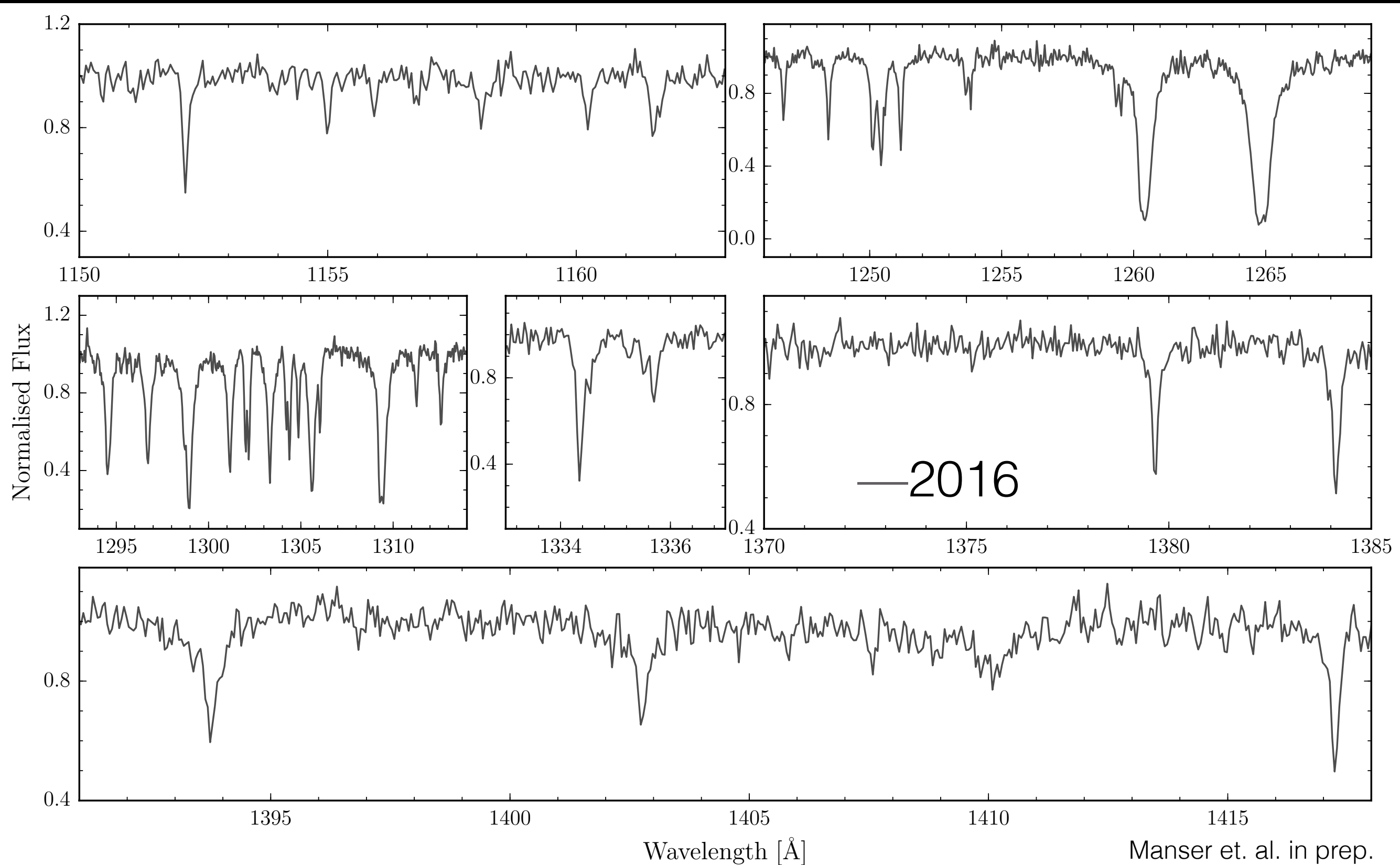
Coadded X-Shooter spectrum



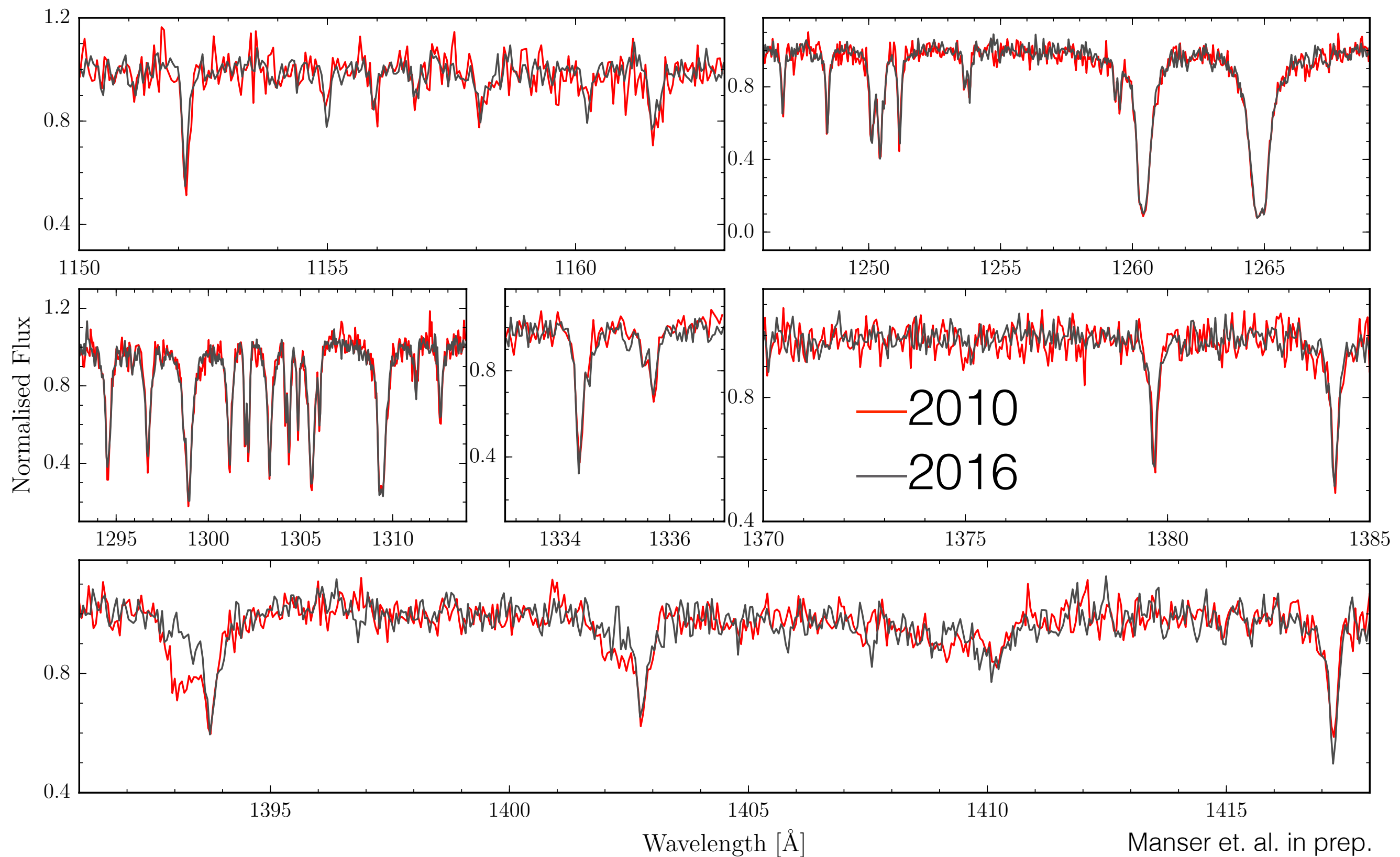
April 2010 Hubble Spectrum



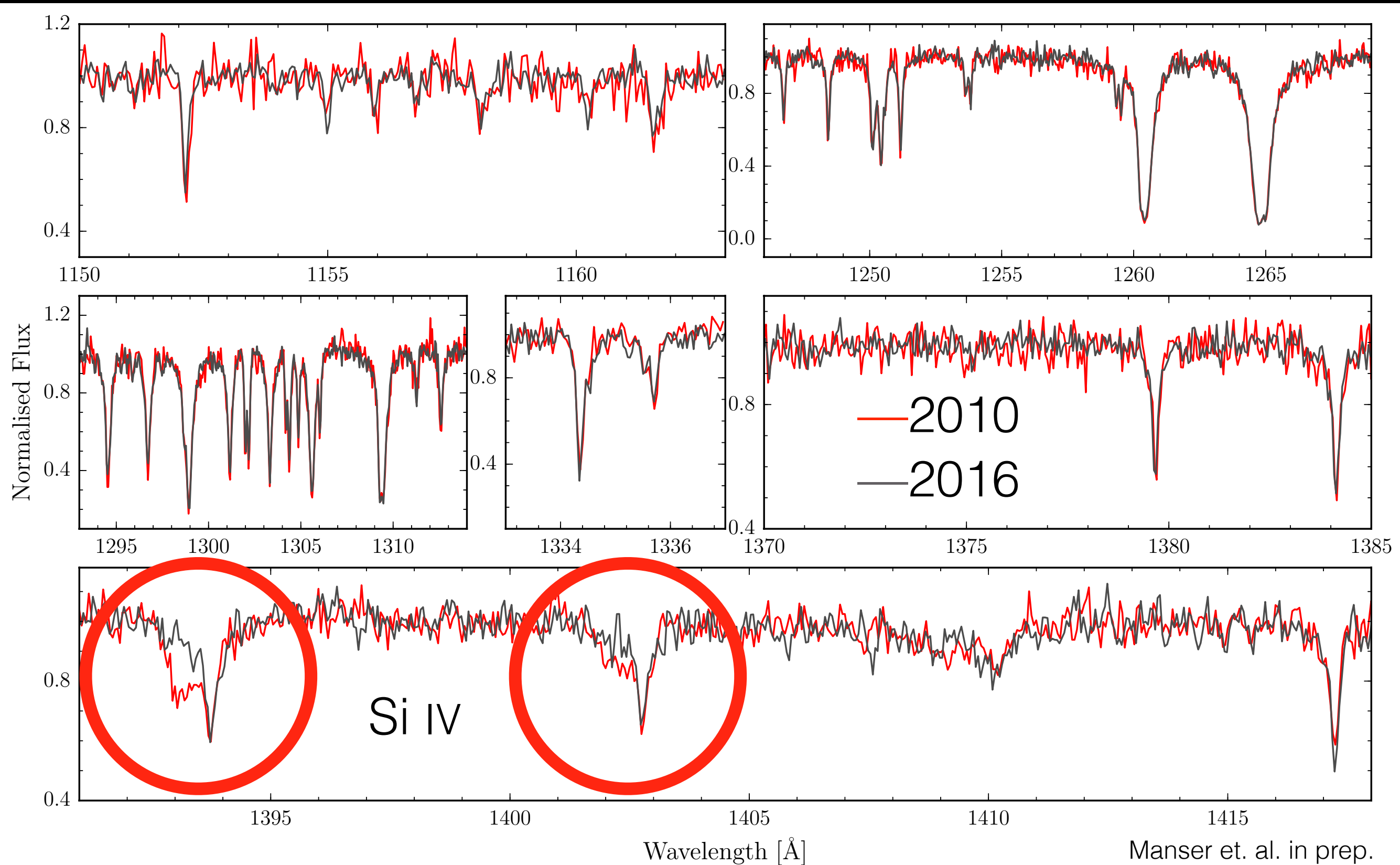
March 2016 Hubble Spectrum



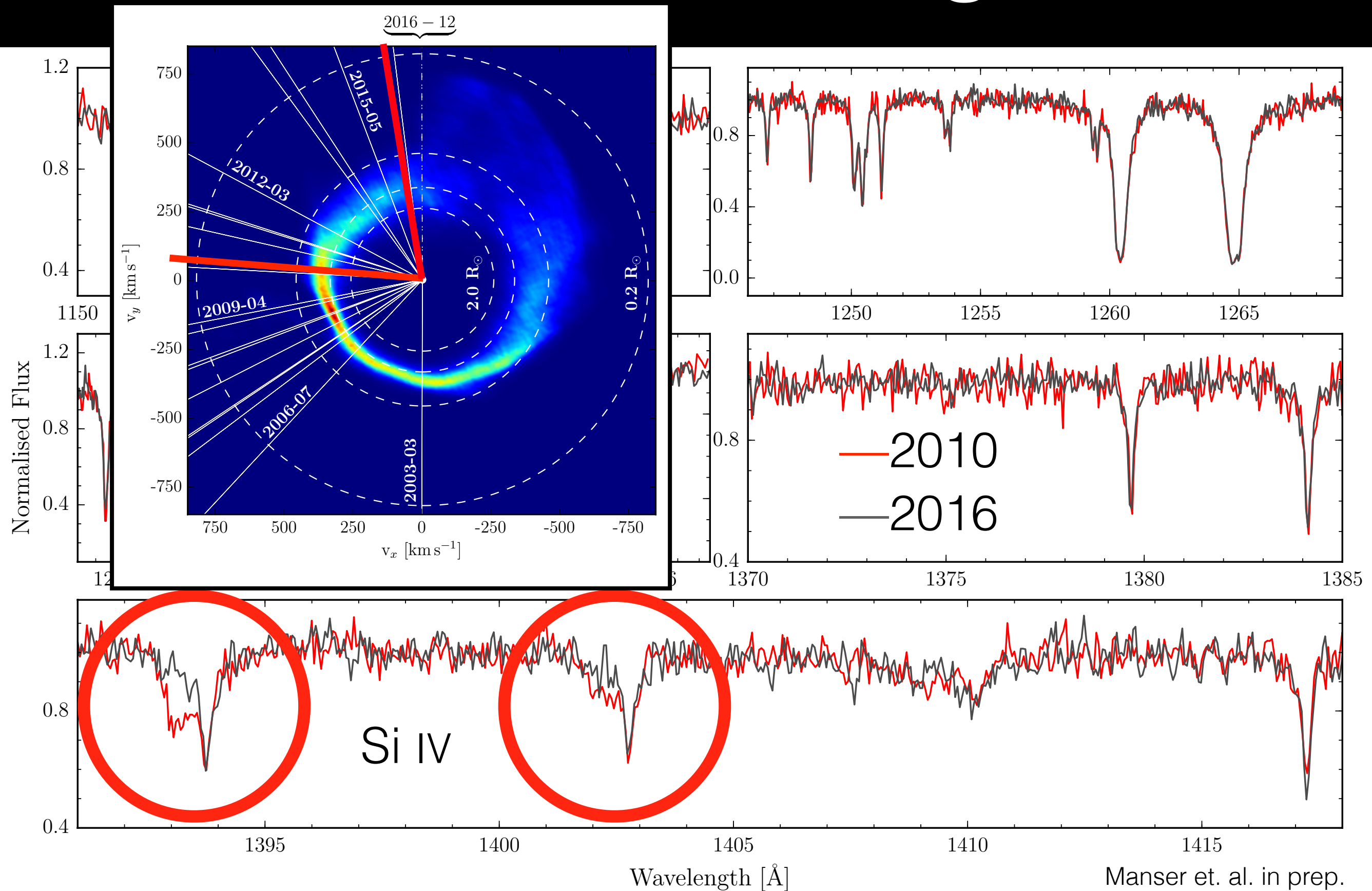
Comparing the two...



Circumstellar gas

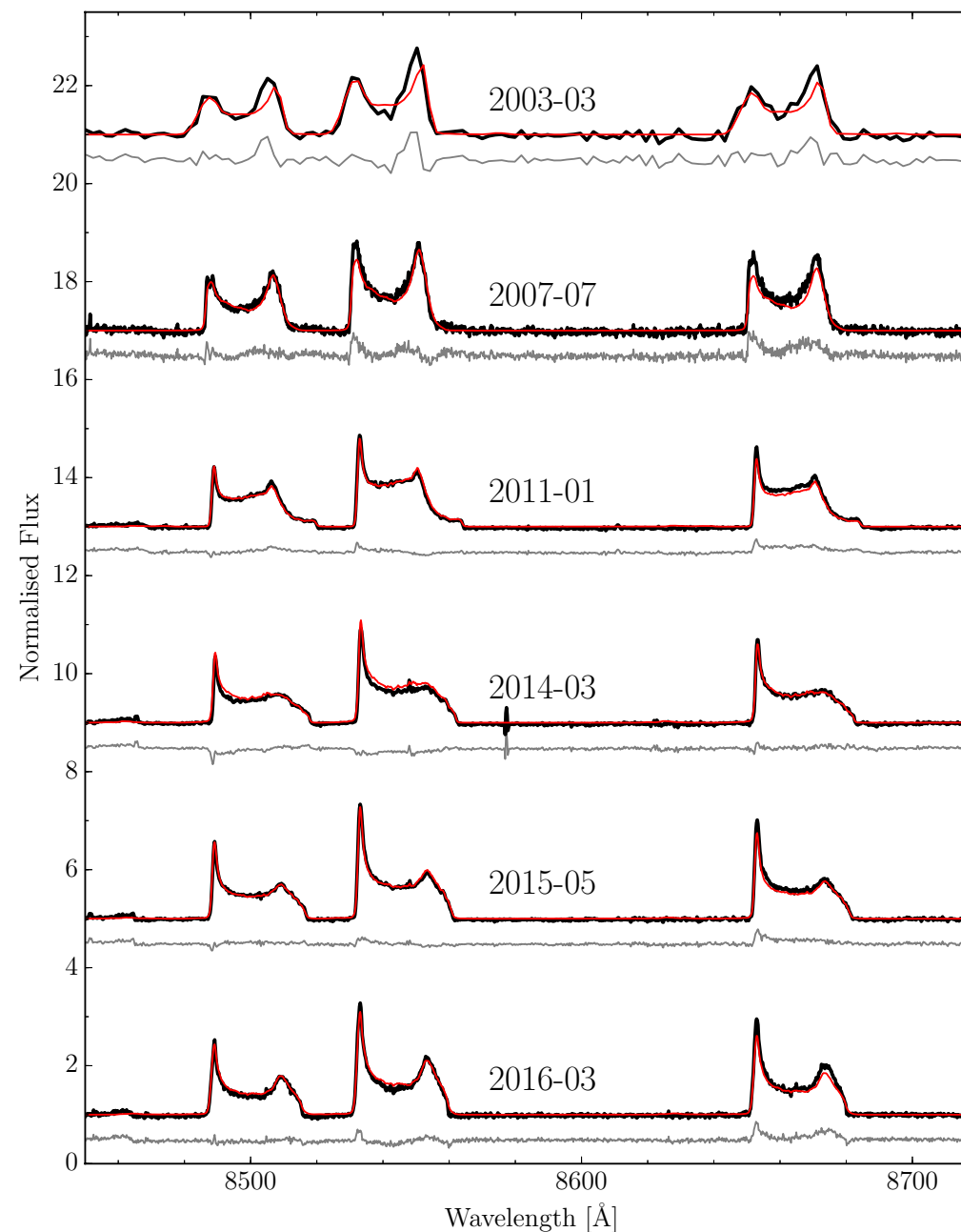


Circumstellar gas



Other variable gas discs

SDSS J1228+1040

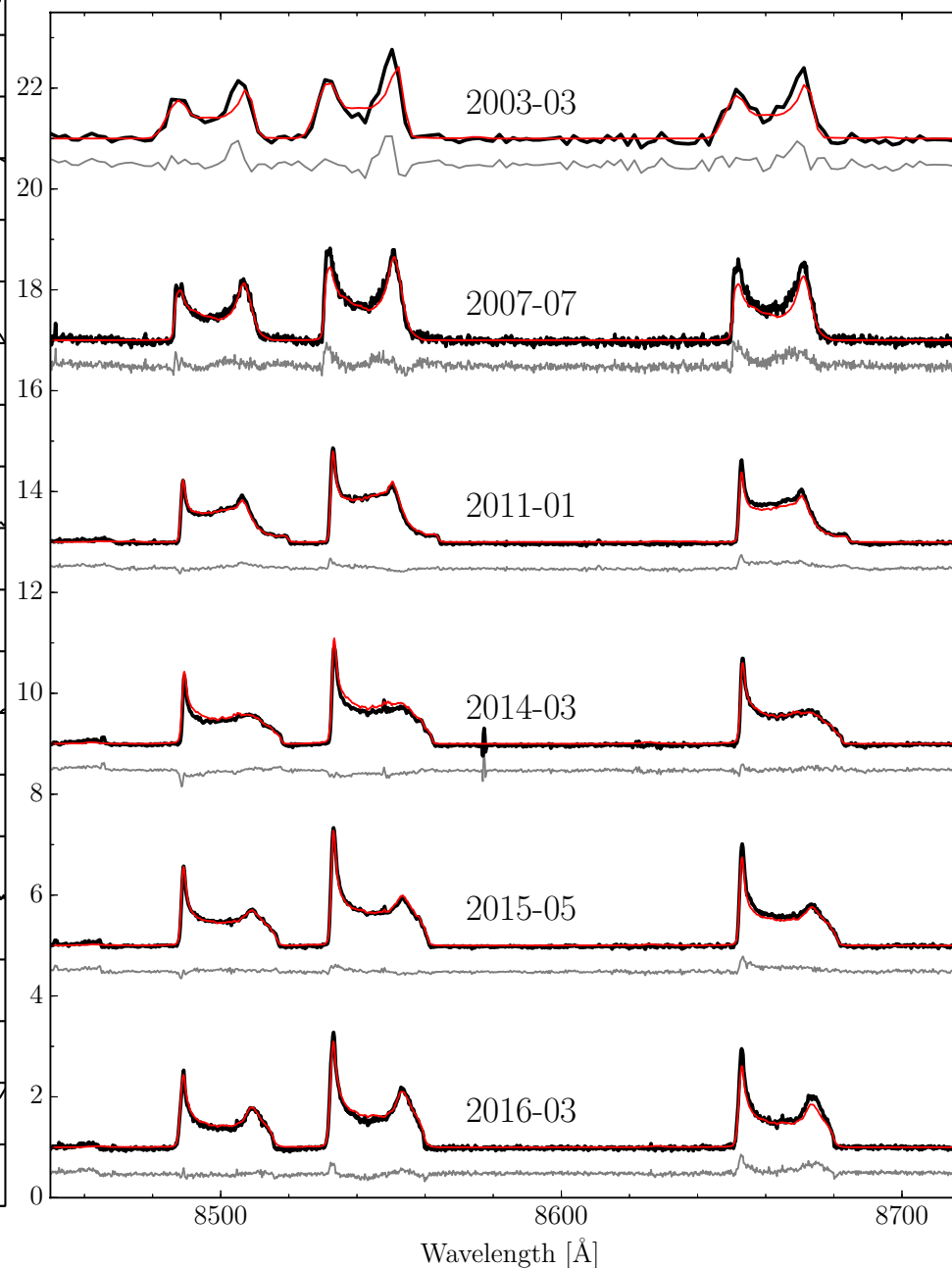
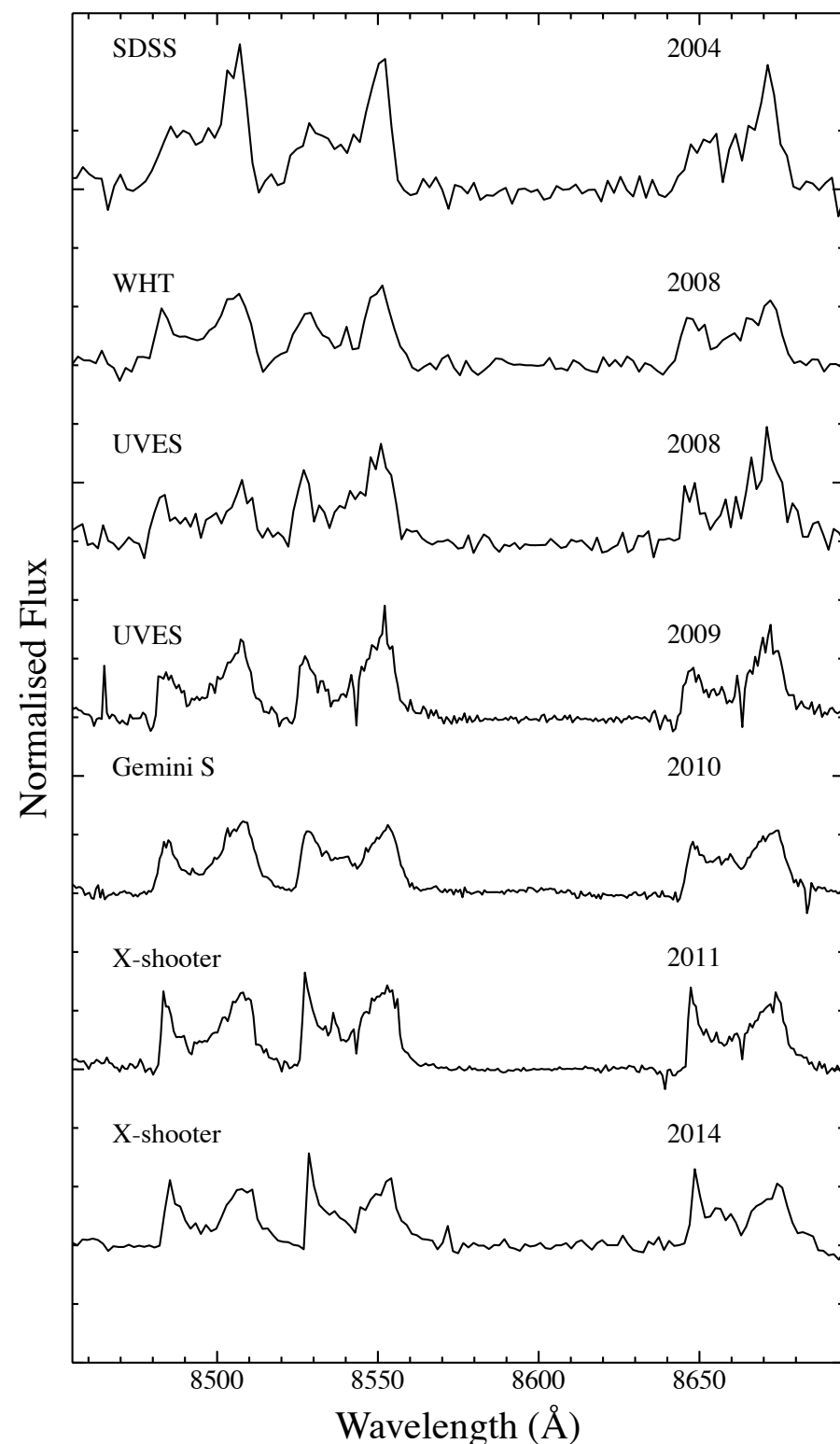


Manser et. al. 2016, MNRAS, 455, 4467

Morphologically variable

SDSS J0845+2257

SDSS J1228+1040



Wilson et. al. 2014, MNRAS, 451, 3237

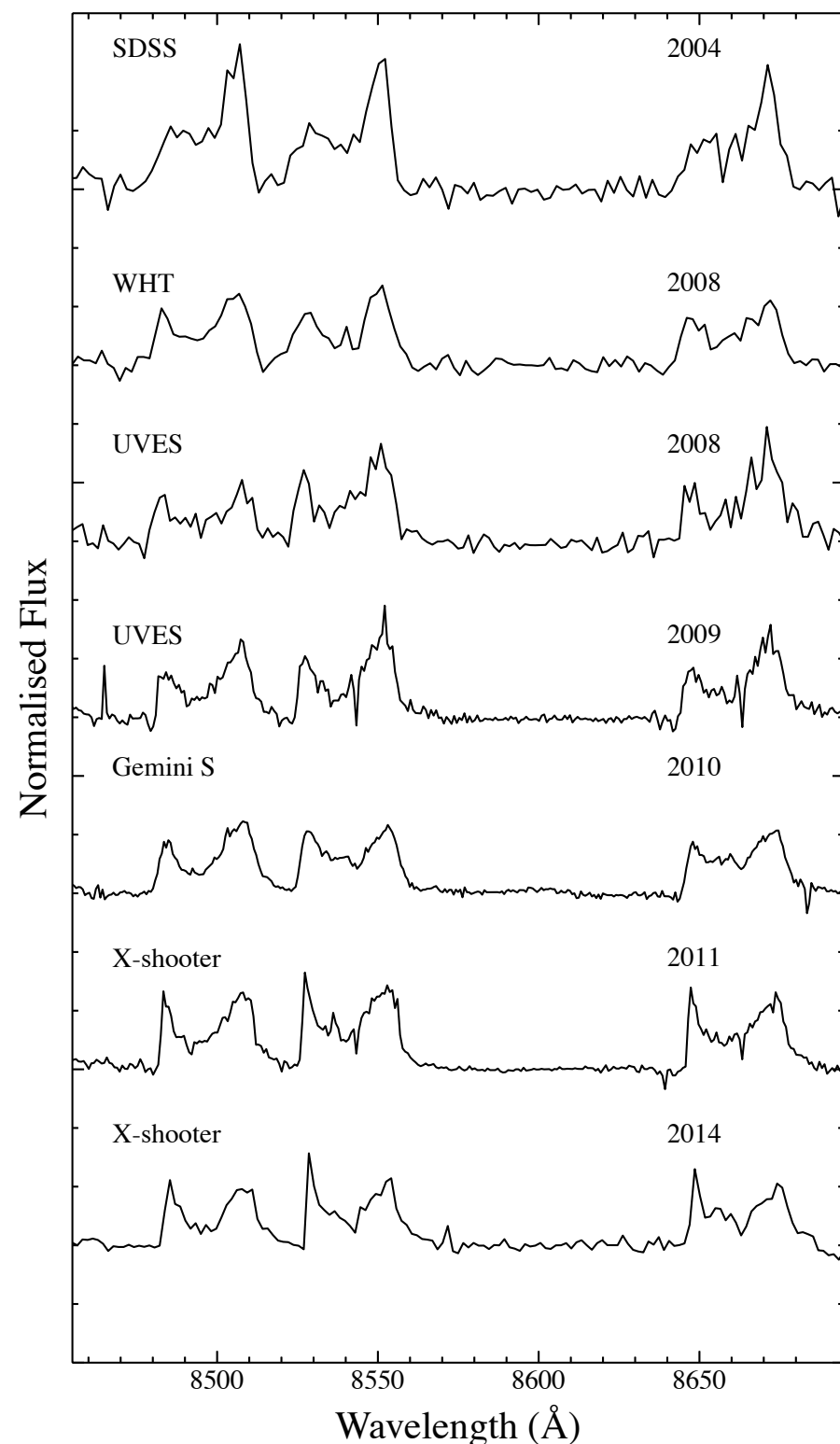
Manser et. al. 2016, MNRAS, 455, 4467

Morphologically variable

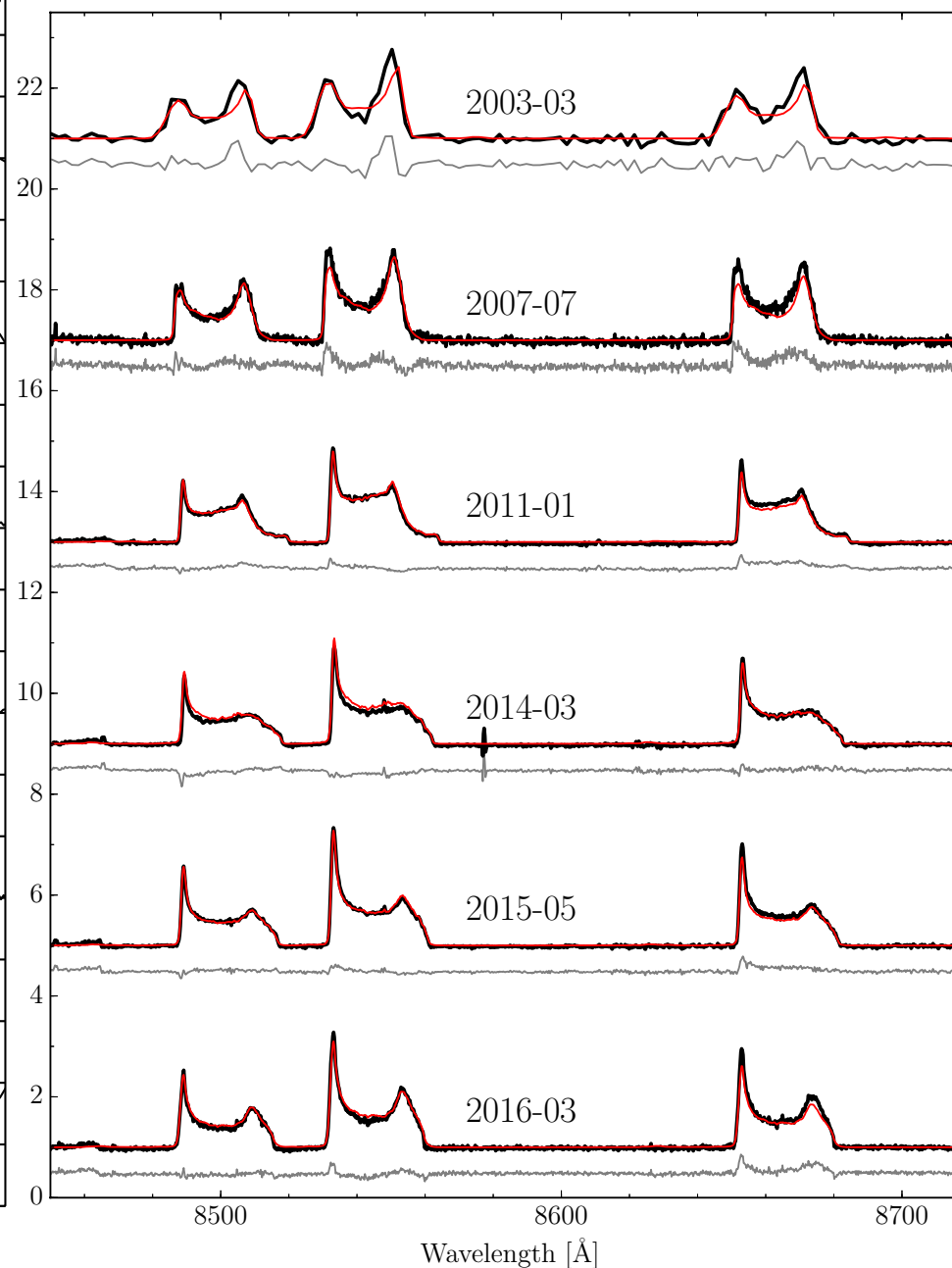
SDSS J0845+2257

SDSS J1228+1040

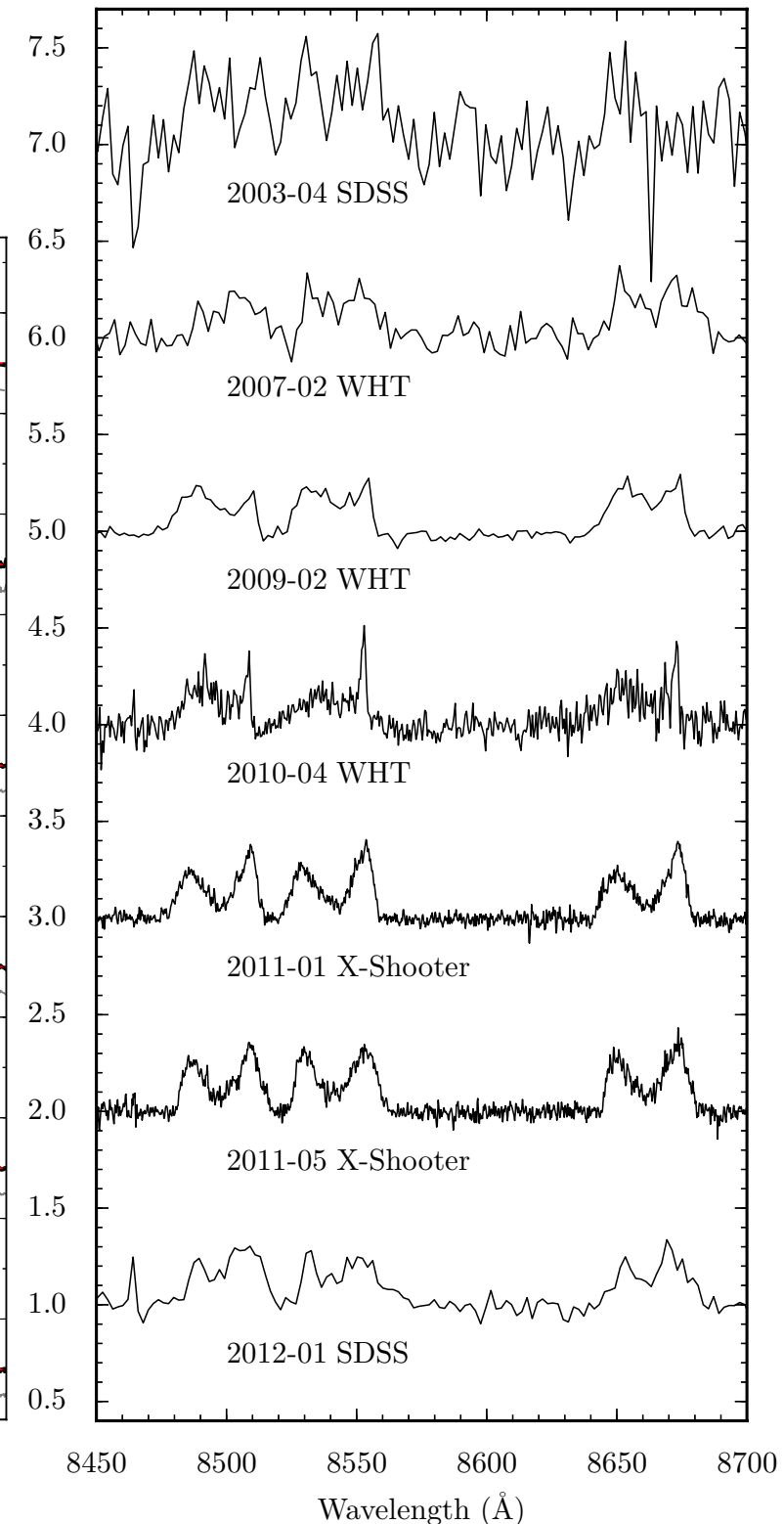
SDSS J1043+0855



Wilson et. al. 2015, MNRAS, 451, 3237



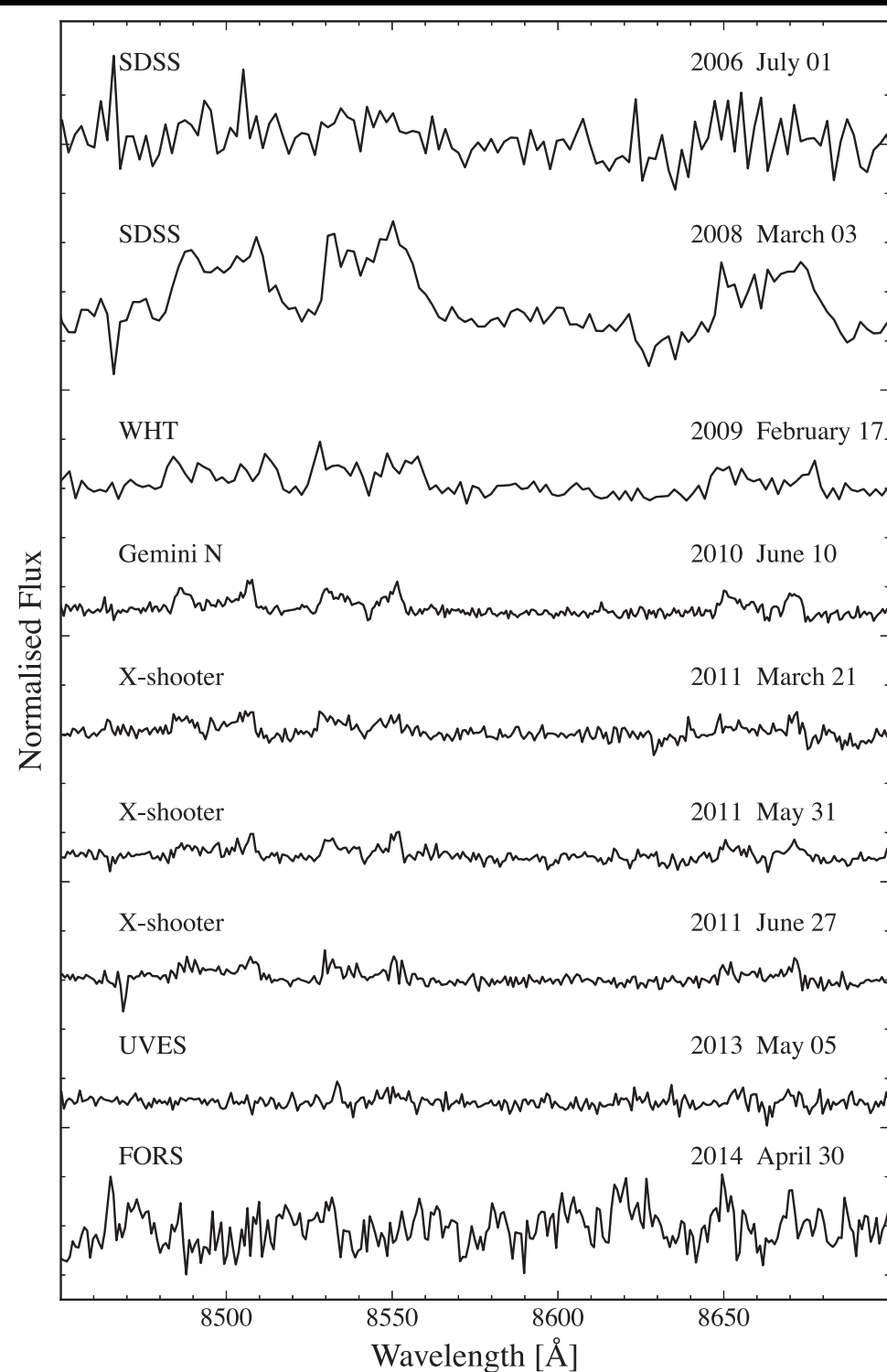
Manser et. al. 2016, MNRAS, 455, 4467



Manser et. al. 2016, MNRAS, 462 1461

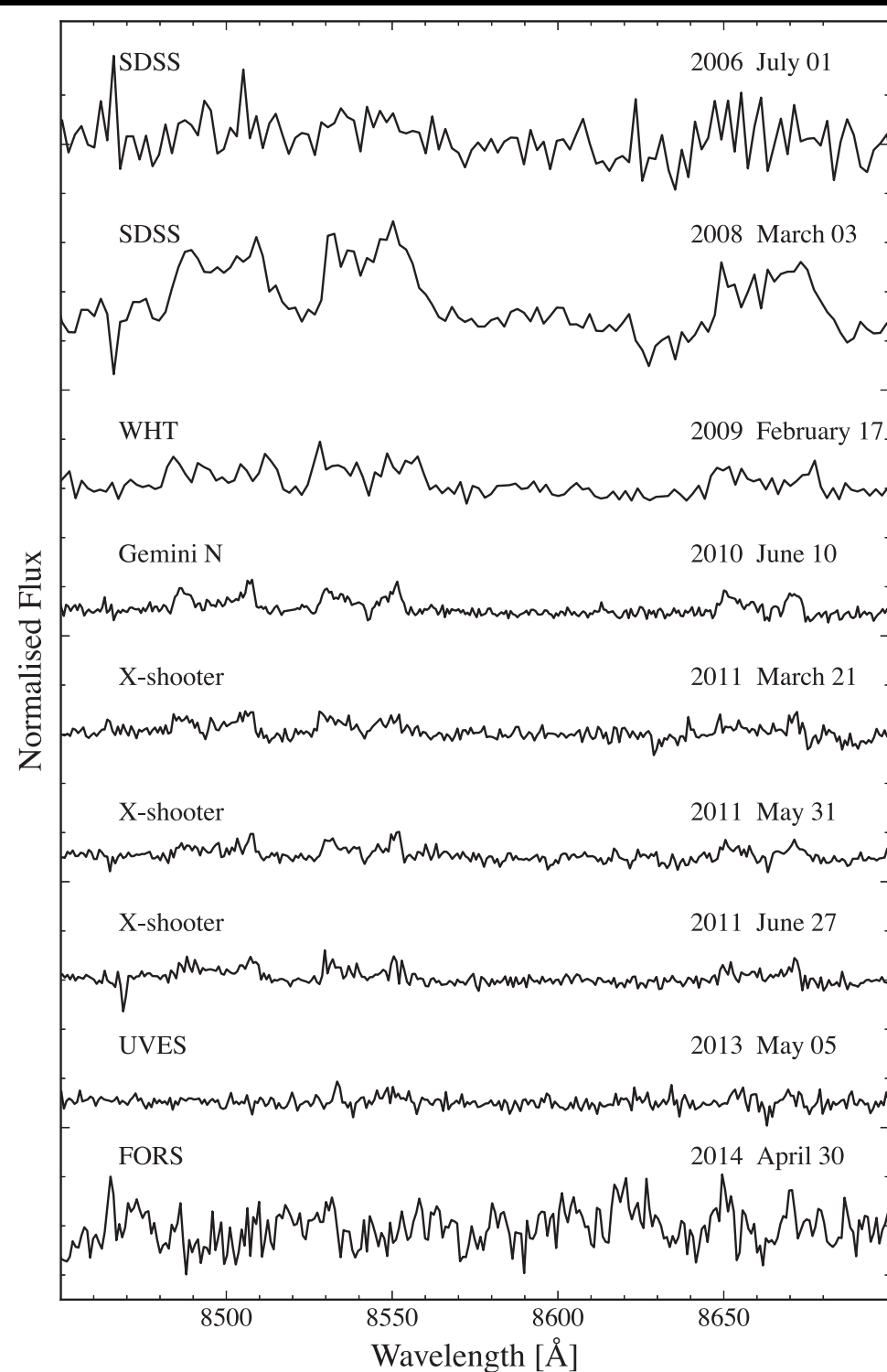
Variable strength

SDSS J1617+1620



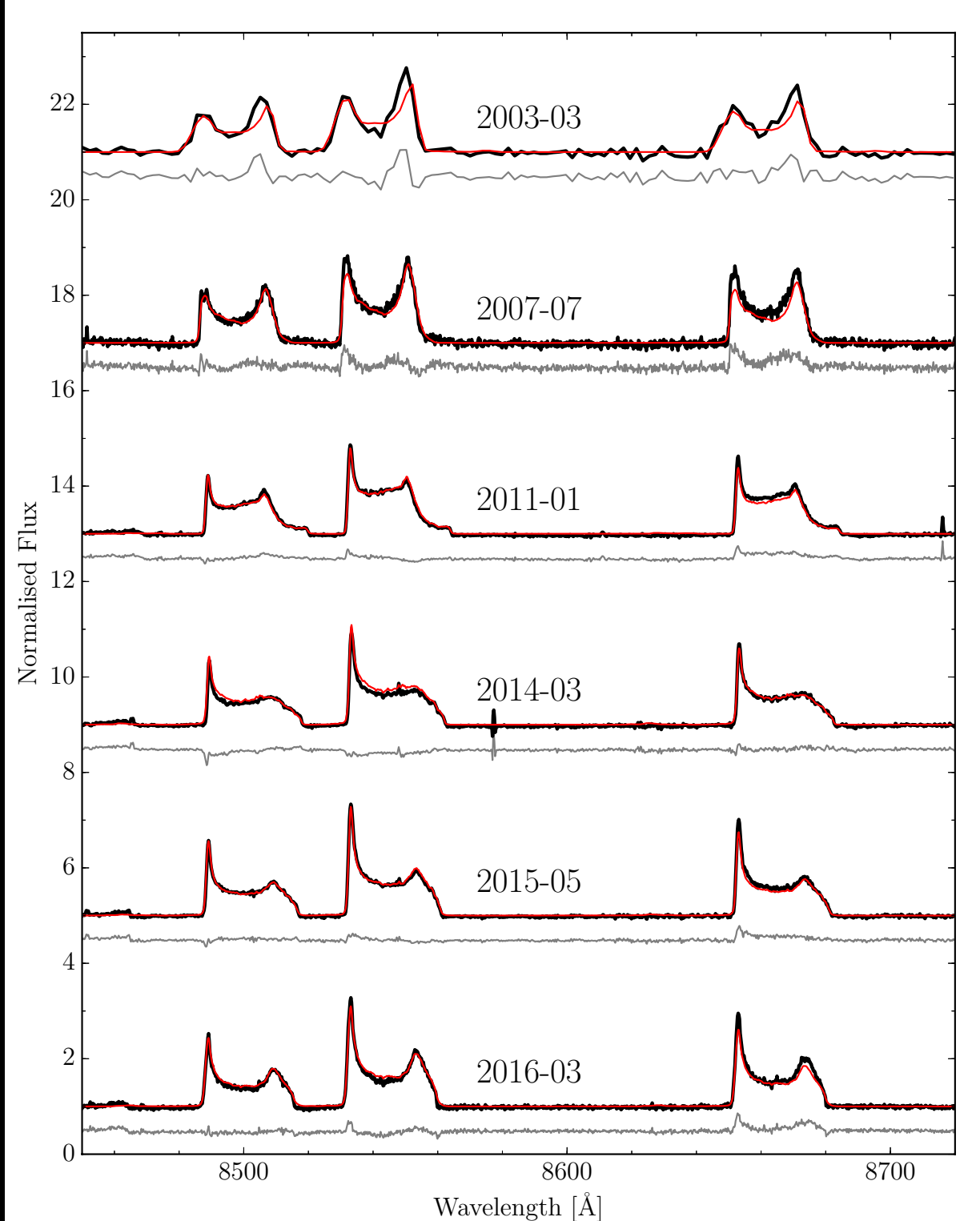
Variable strength

SDSS J1617+1620



Wilson et. al. 2014, MNRAS, 445, 1878

SDSS J1228+1040



Manser et. al. 2016, MNRAS, 455, 4467

Variability

e - Gaseous emission

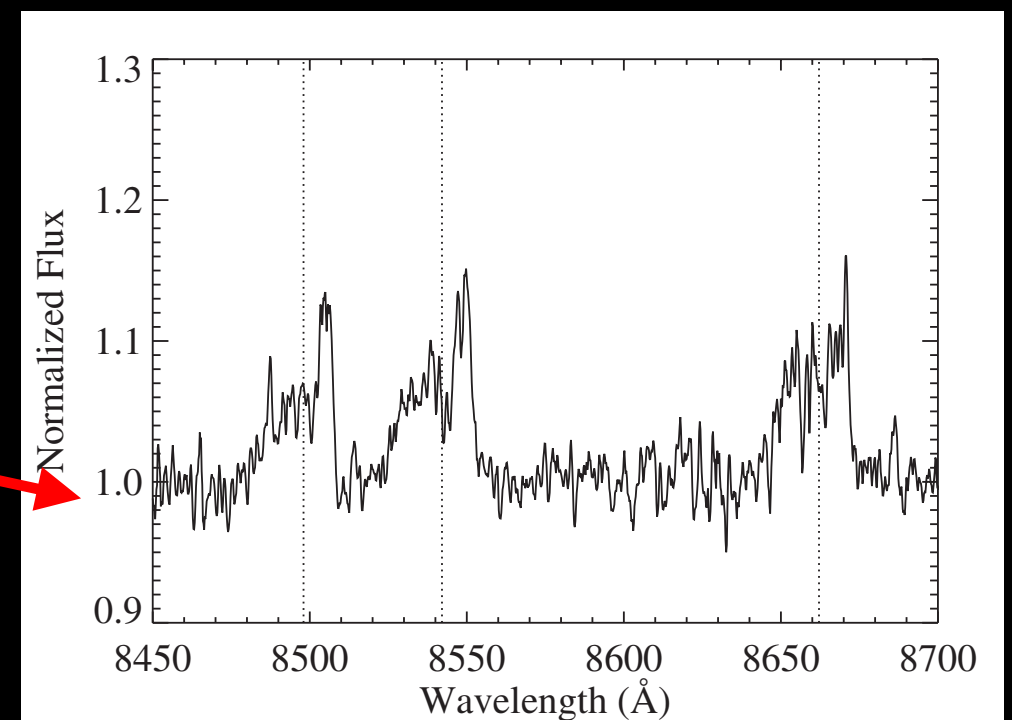
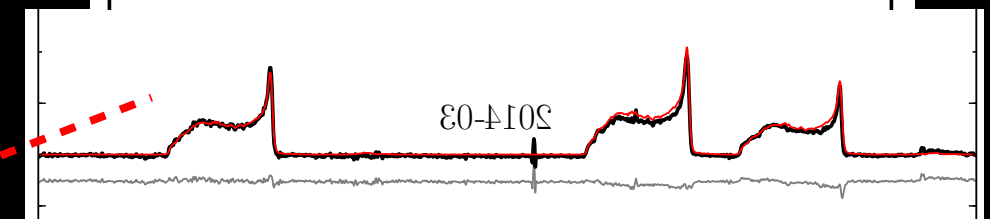
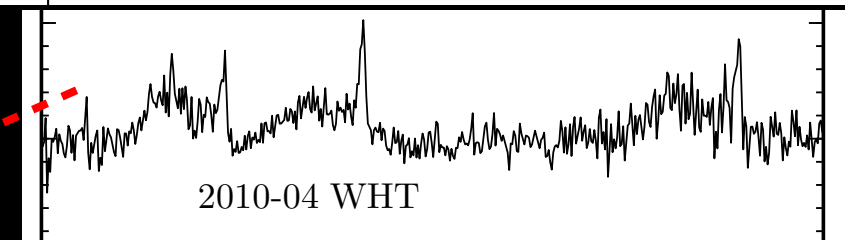
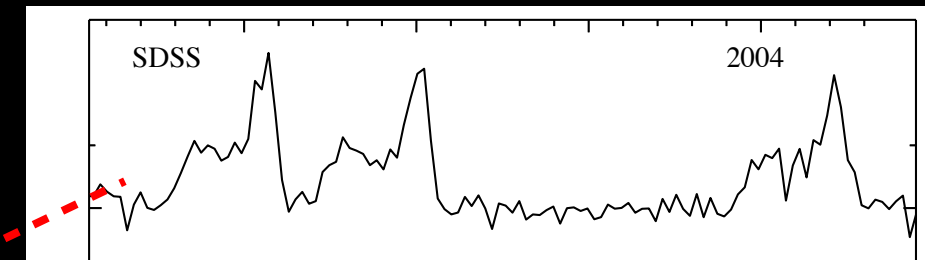
a - Gaseous absorption

v - Spectroscopic or
photometric Variability

Name	Type	Features
SDSS J0738+1835	DB	e
SDSS J0845+2257	DB	e, v
SDSS J0959-0200	DA	e, v
SDSS J1043+0855	DA	e, v
WD 1144+0529	DA	e
SDSS J1228+1040	DA	e, a, v
HE 1349-2305	DBA	e
SDSS J1617+1620	DA	e, v

Variability

Name	Type	Features
SDSS J0738+1835	DB	e
SDSS J0845+2257	DB	e, v
SDSS J0959-0200	DA	e, v
SDSS J1043+0855	DA	e, v
WD 1144+0529	DA	e
SDSS J1228+1040	DA	e, a, v
HE 1349-2305	DBA	e
SDSS J1617+1620	DA	e, v

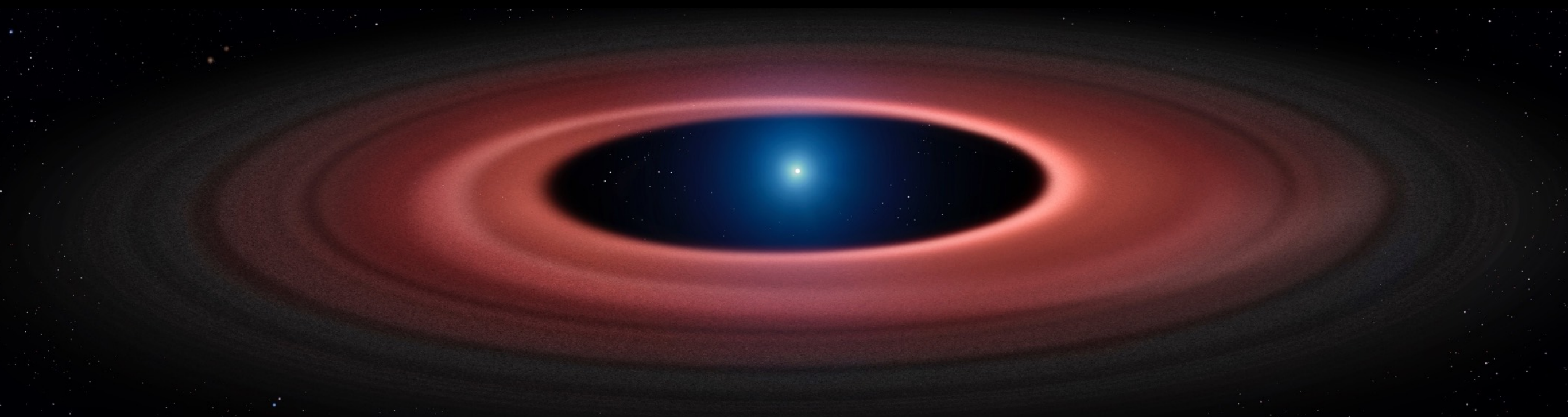


Melis et. al. 2012, ApJL, 715, L4

Detected Remnant Planetary System statistics

Metal pollution Koester et. al. 2014

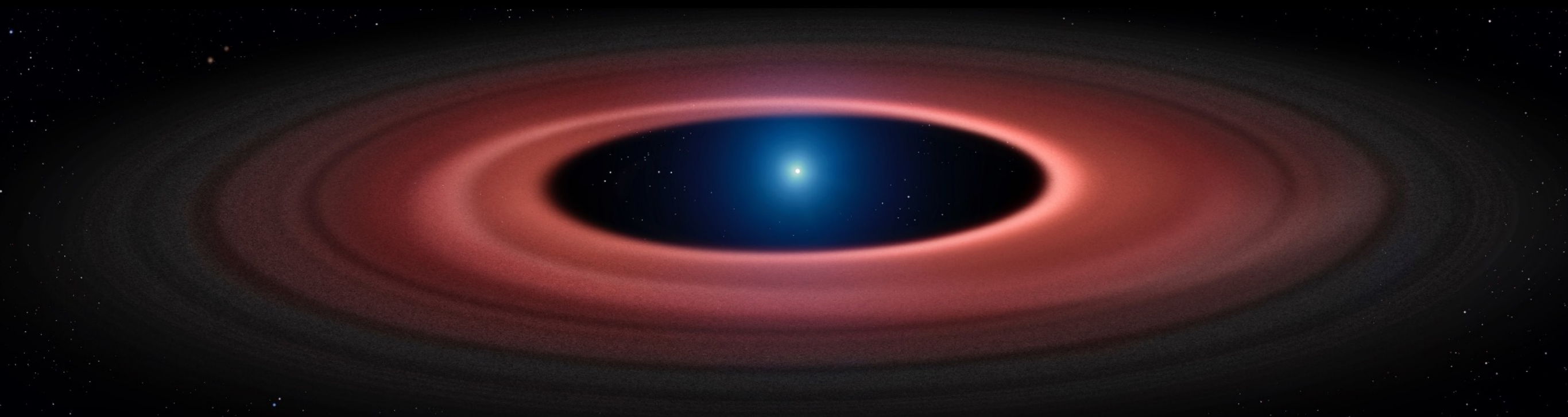
25 - 50 %



Detected Remnant Planetary System statistics

Metal pollution Koester et. al. 2014 25 - 50 %

Dusty disc Farihi et al. 2009 Rocchetto et al. 2015 1 - 3 %

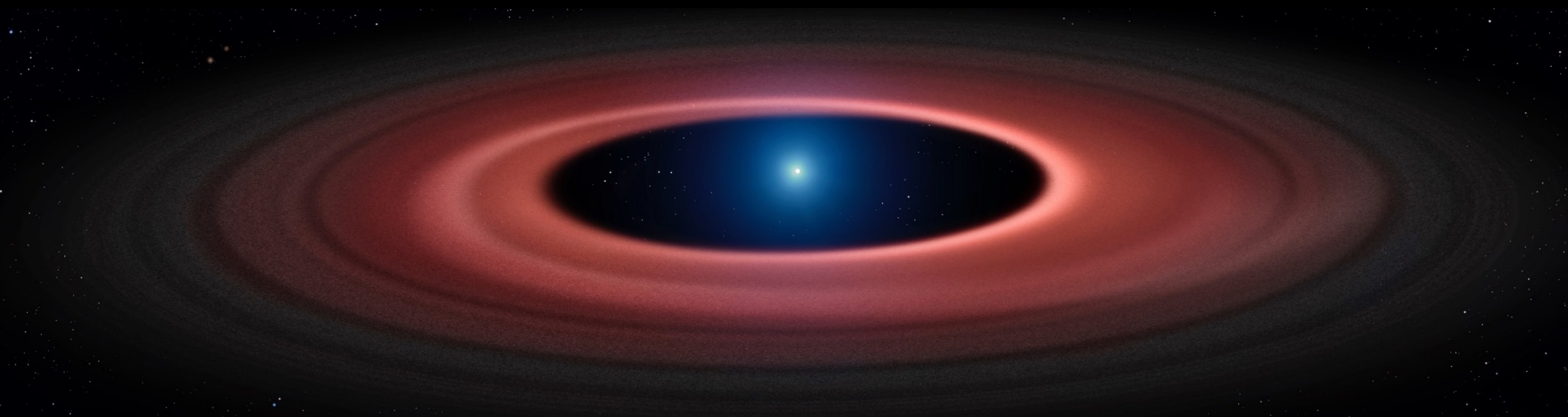


Detected Remnant Planetary System statistics

Metal pollution Koester et. al. 2014 25 - 50 %

Dusty disc Farihi et al. 2009 Rocchetto et al. 2015 1 - 3 %

Gaseous component ??? %



The sample

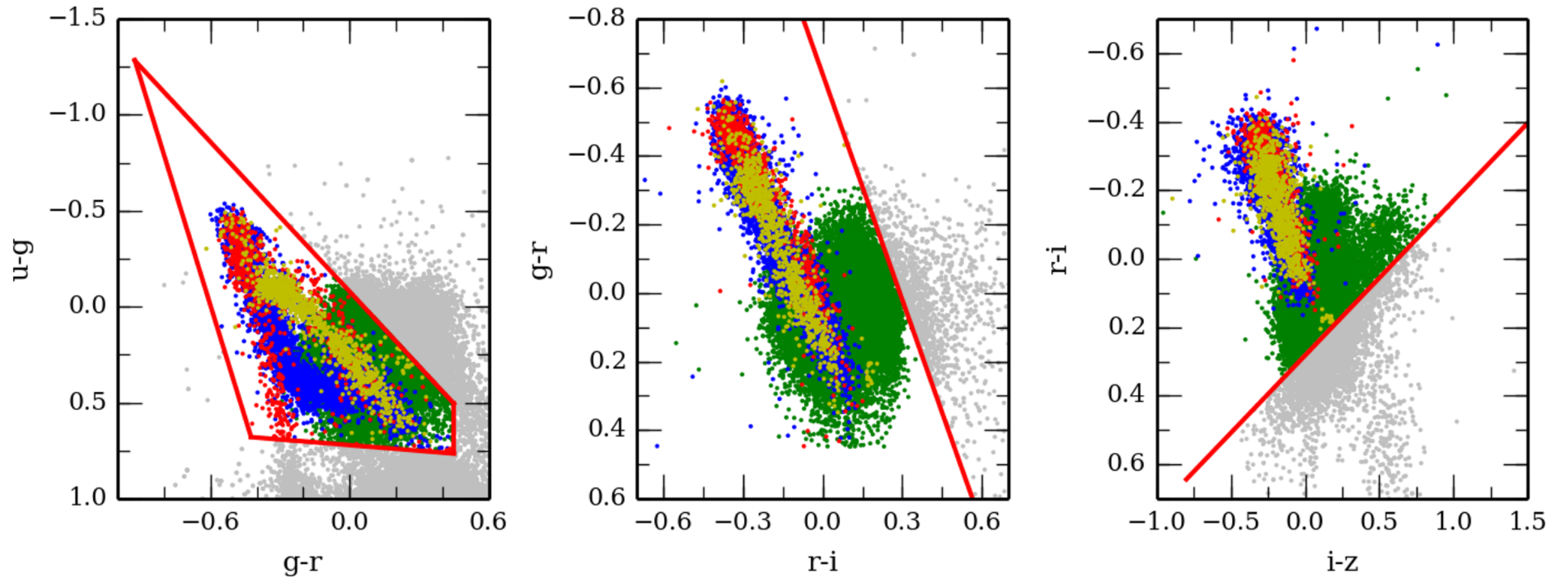
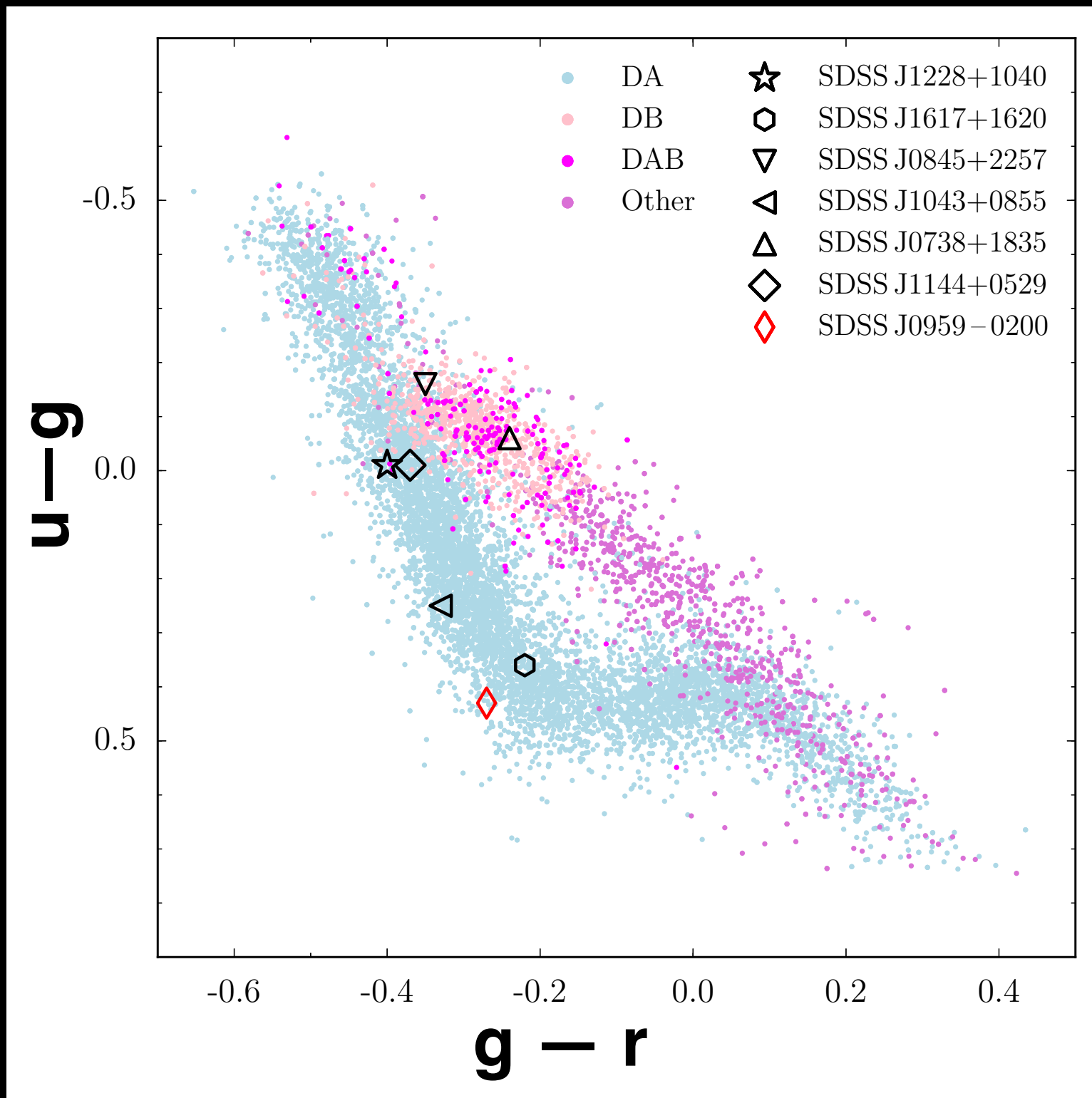


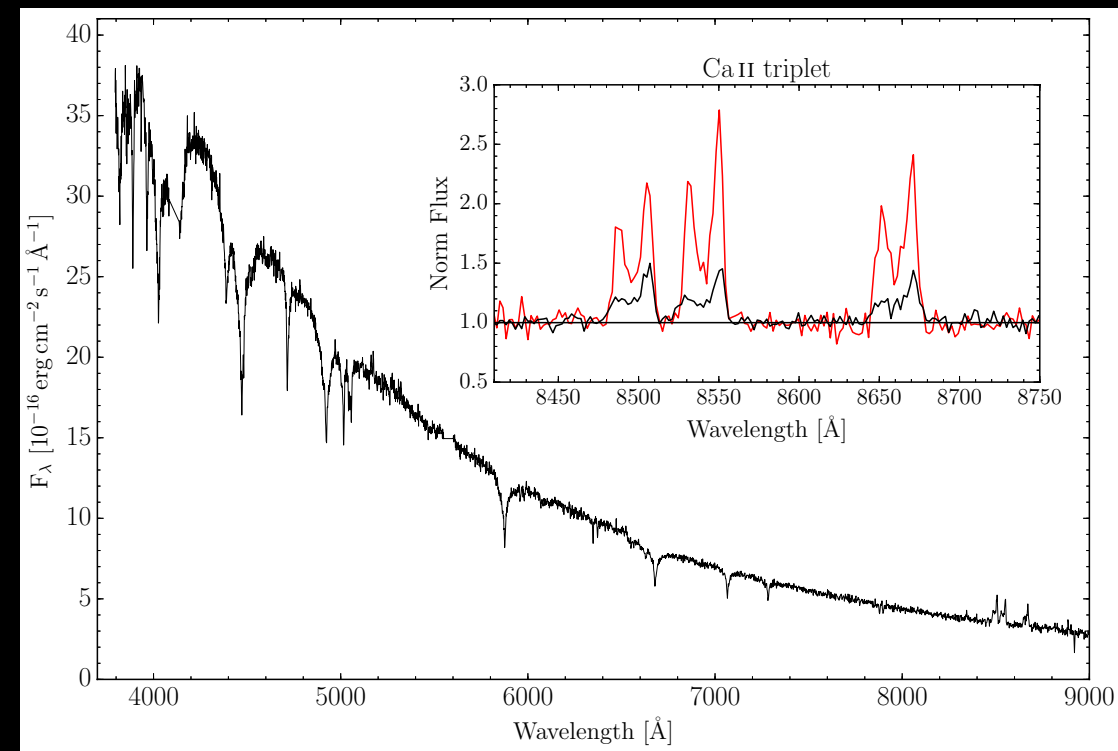
Figure 2. Colour-colour diagrams illustrating the location of the 27 639 DR7 spectroscopic objects that we used as training sample for our selection method. DA white dwarfs, non-DA white dwarfs, NLHS and quasars are shown as blue, yellow, red and green dots, respectively. The colour cuts that define our initial broad selection from Table 2 are overlaid as red lines. Objects outside this selection were not classified and are therefore plotted as grey dots.

The sample



The frequency of gaseous discs

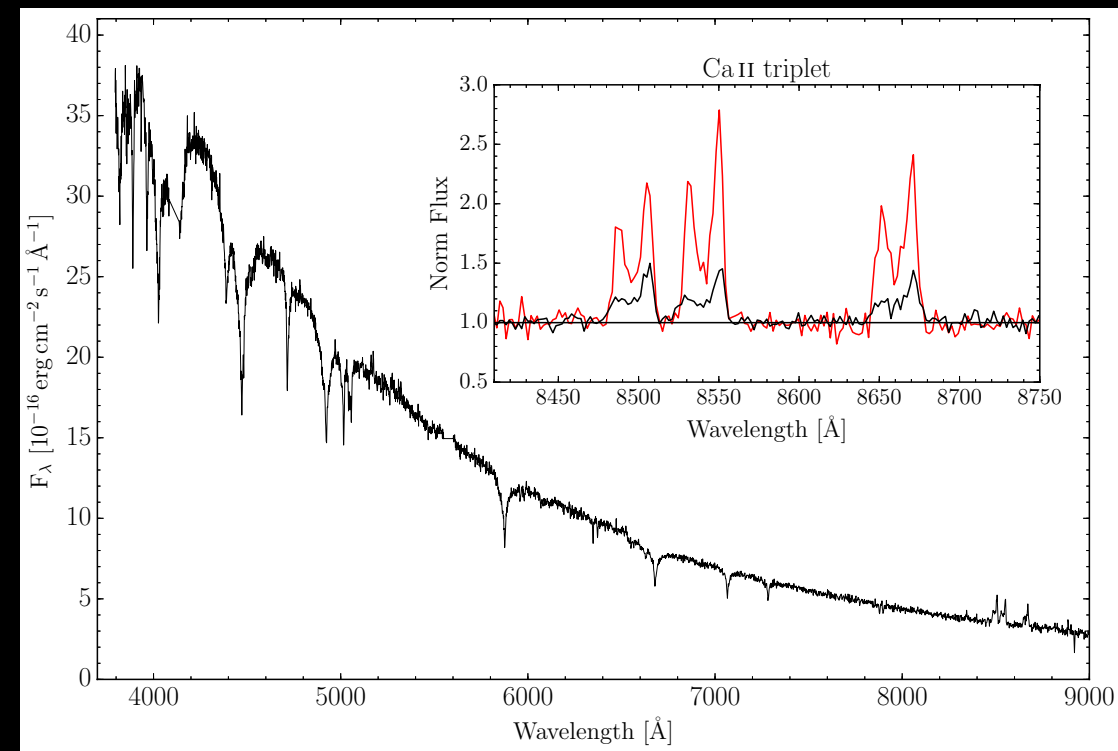
9079 single white dwarfs



The frequency of gaseous discs

9079 single white dwarfs

6 Gaseous components

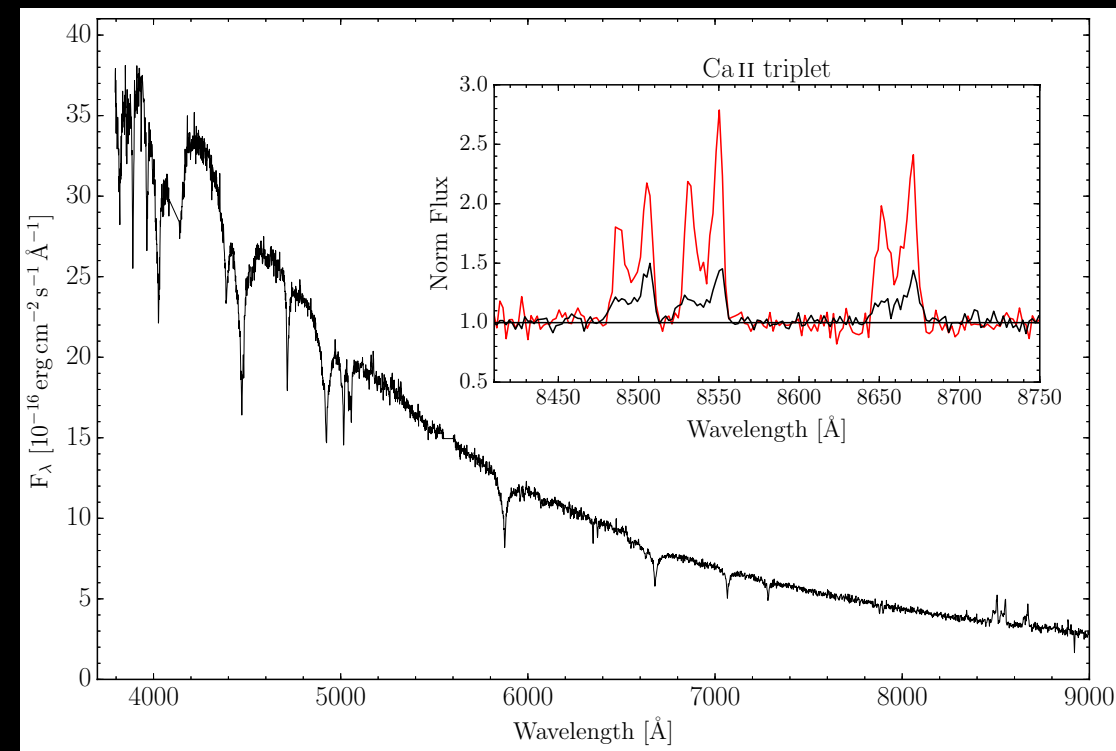


The frequency of gaseous discs

9079 single white dwarfs

6 Gaseous components

Frequency of observable
gaseous debris discs at
white dwarfs



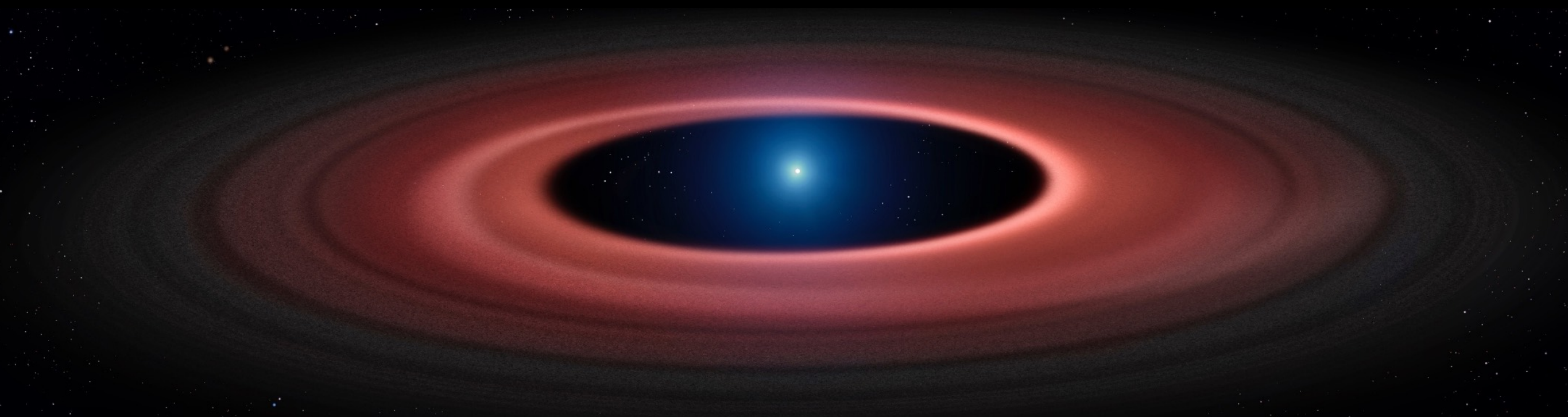
$0.07^{+0.03}_{-0.02} \%$

Detected Remnant Planetary System statistics

Metal pollution Koester et. al. 2014 25 - 50 %

Dusty disc Farihi et al. 2009 Rocchetto et al. 2015 1 - 3 %

Gaseous component 0.07 %



Detected Remnant Planetary System statistics

Metal pollution Koester et. al. 2014 25 - 50 %

Dusty disc Farihi et al. 2009 Rocchetto et al. 2015 1 - 3 %

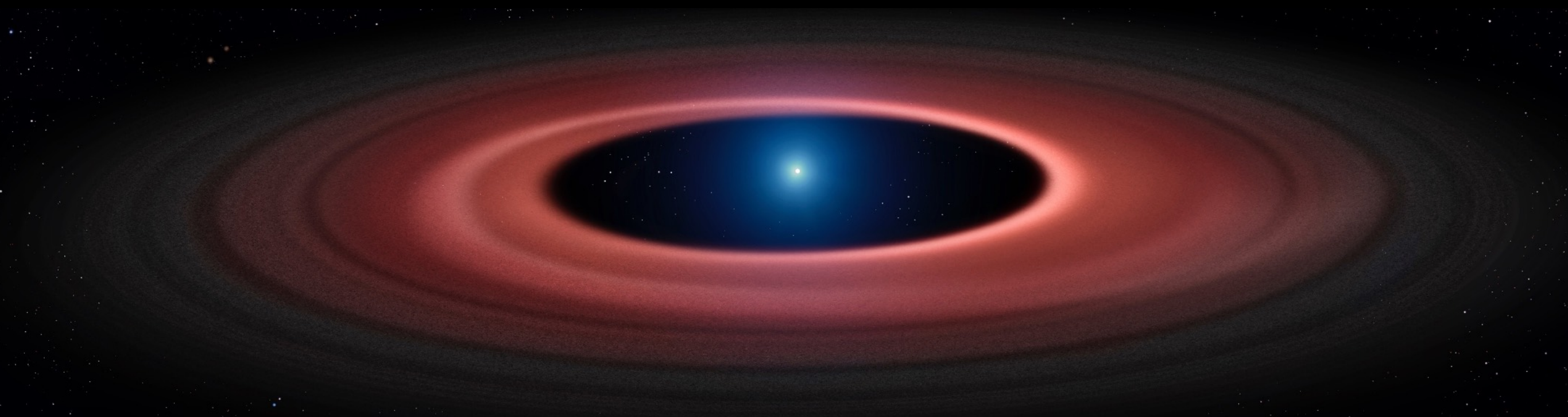
Gaseous component 0.07 %

Debris discs with a gaseous component 2 - 10 %



Summary

- SDSS J1228+1040 is well studied, but still many unanswered questions.
- An observable gaseous component appears to be linked with variability
- Determined the frequency of a gaseous component to a debris disc at a white dwarf.



Summary

- SDSS J1228+1040 is well studied, but still many unanswered questions.
- An observable gaseous component appears to be linked with variability
- Determined the frequency of a gaseous component to a debris disc at a white dwarf.

A white dwarf star, a small and hot remnant of a star, is shown at the center. It is surrounded by a large, flat, glowing red disc of gas and dust, known as a debris disc. The disc is illuminated by the star, creating a bright ring of light. The background is a dark, starry space.

Thanks for listening!